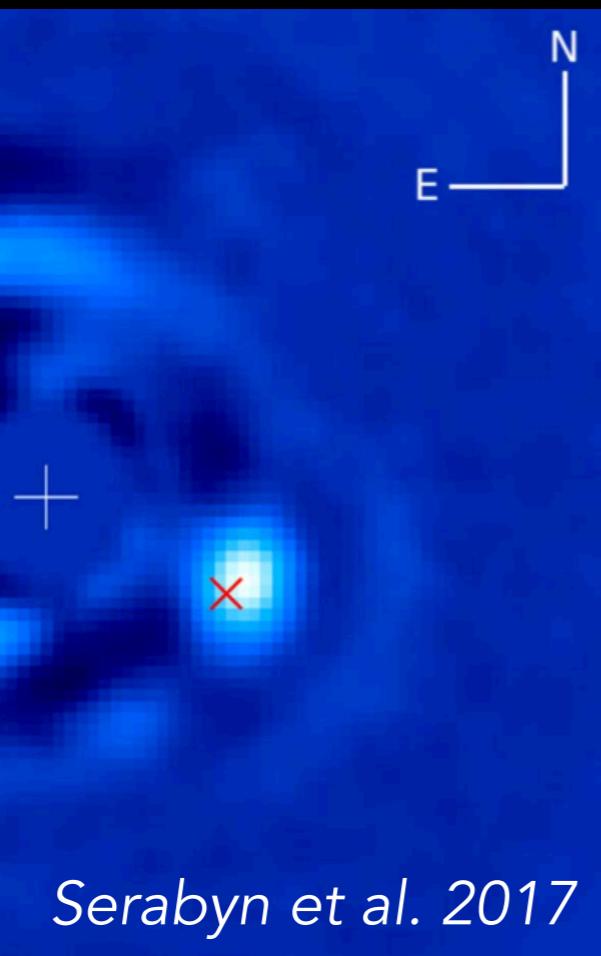
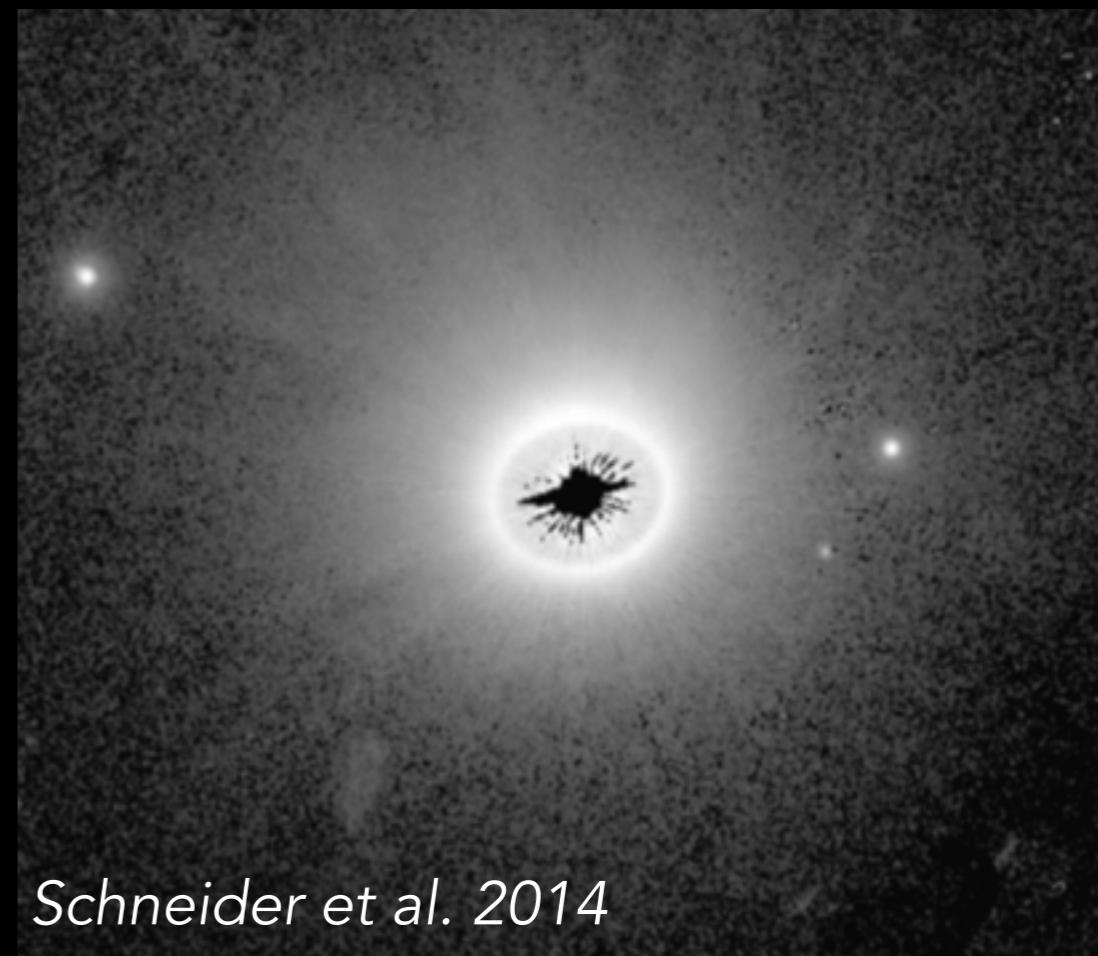


RDI-importance of PSF libraries

Elodie Choquet
(Hubble Fellow, Caltech)



Serabyn et al. 2017

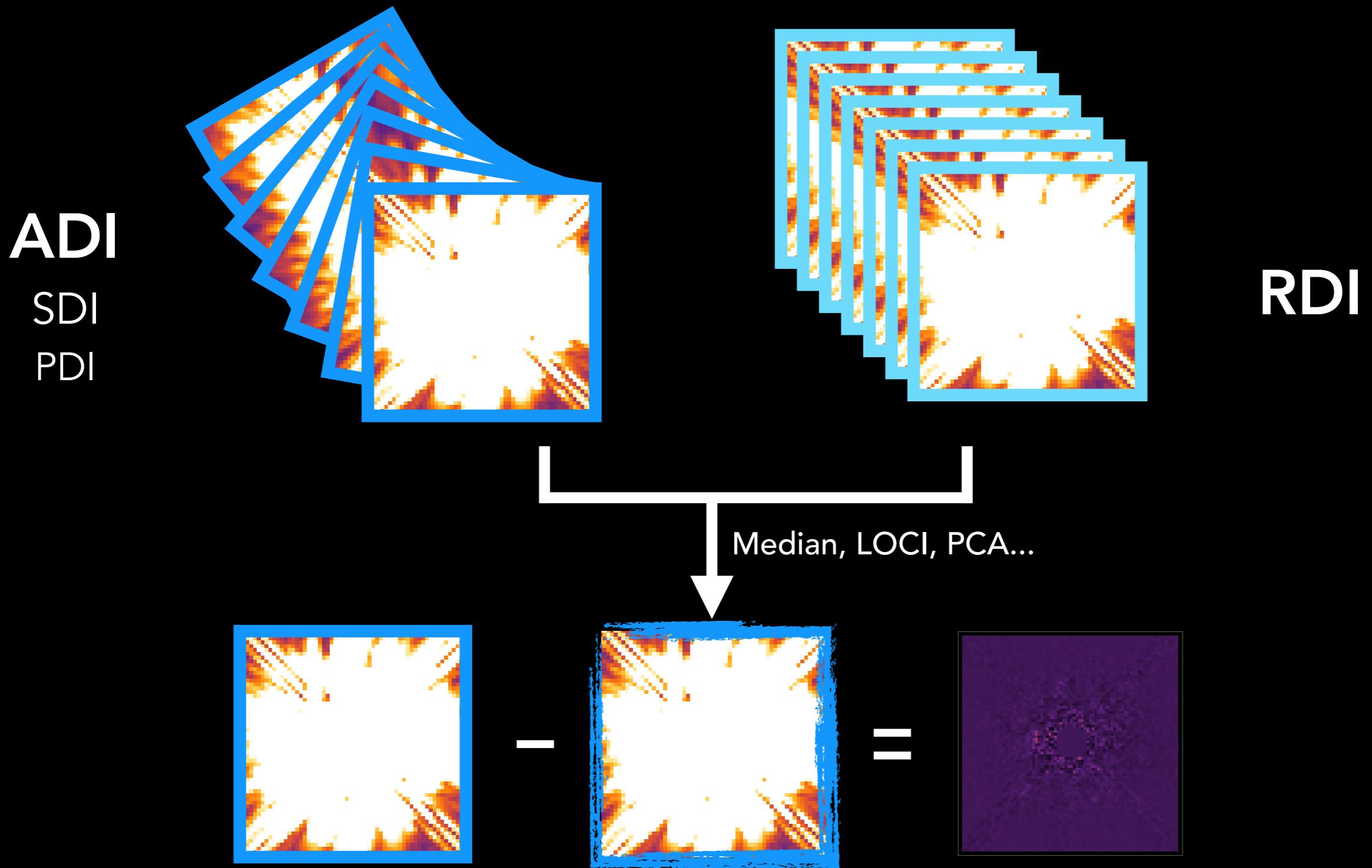


Schneider et al. 2014

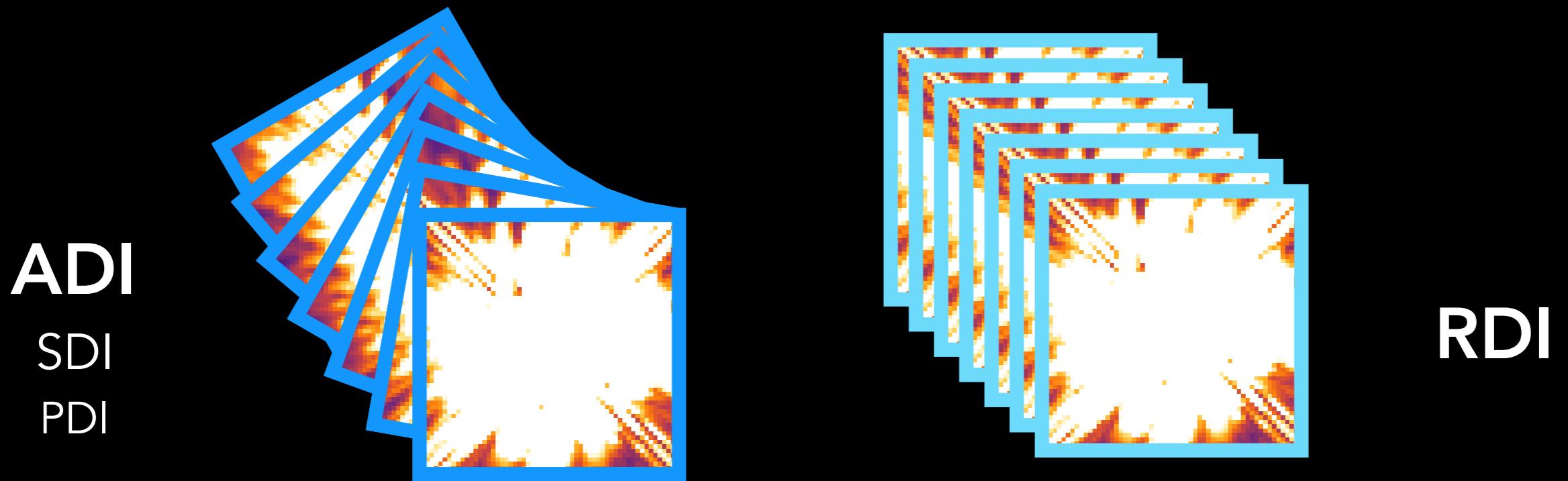
Outline

- ▶ RDI vs. Others (Ground-based case)
- ▶ RDI for Space Telescopes
- ▶ MRDI for Space Telescopes

Target vs. Reference

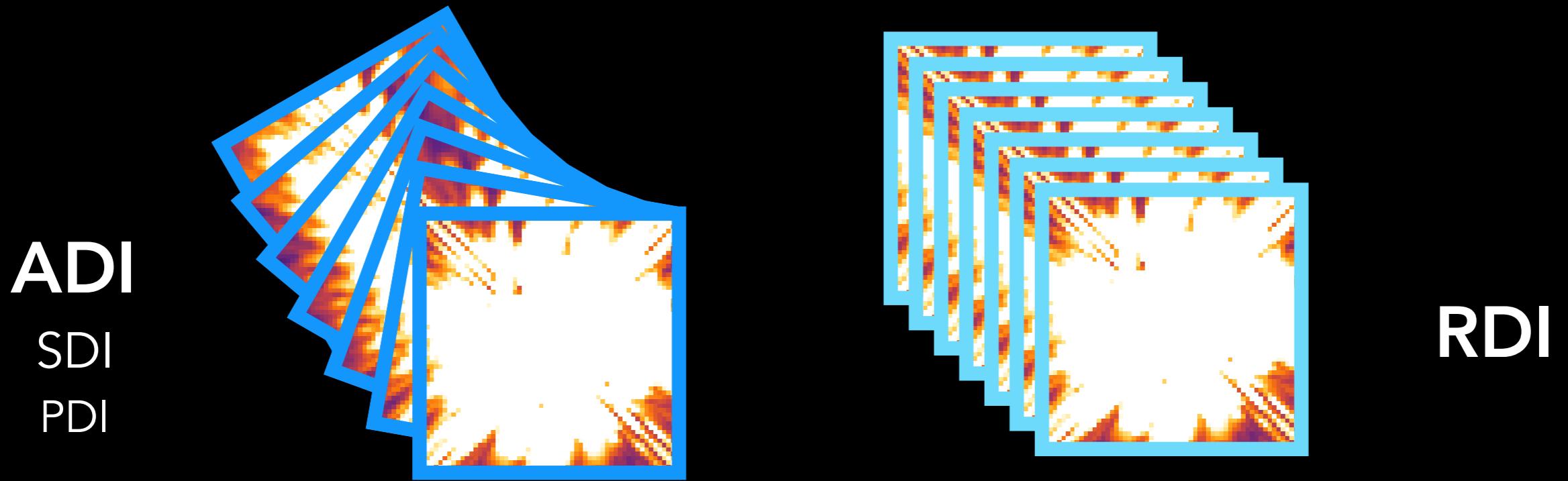


Target vs. Reference



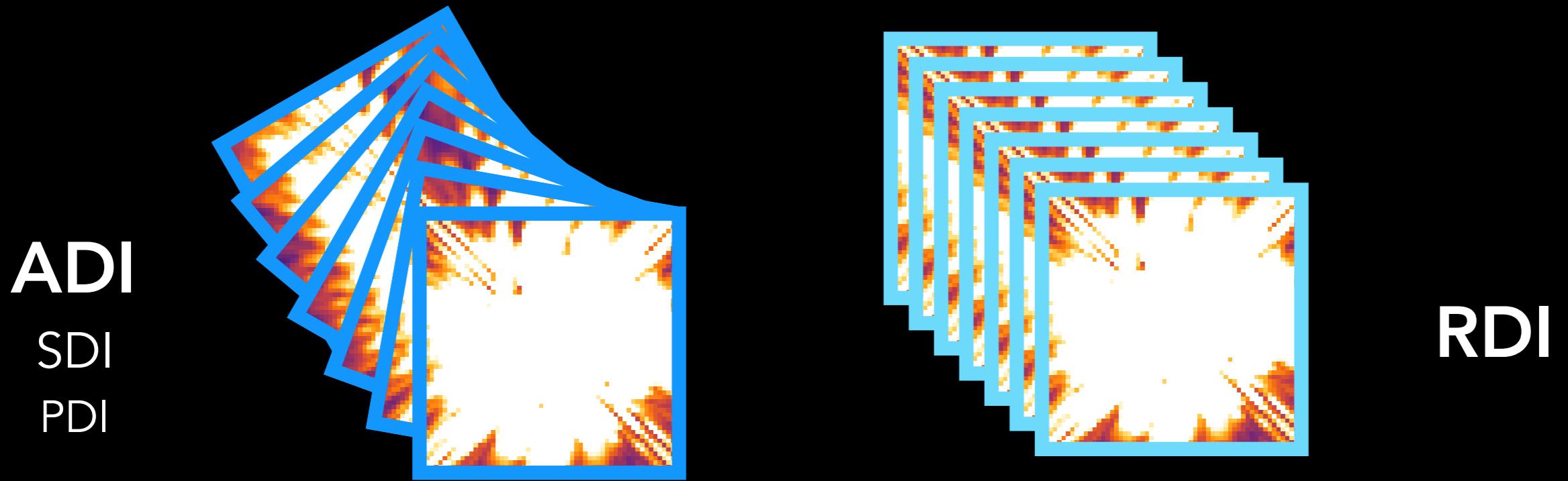
	Target-based	Reference-based
WFE-matching	✓	✗
λ -matching	✓	✗
Self-subtraction	✗	✓

Target vs. Reference



	Target-based	Reference-based
WFE-matching	✓	✗
λ -matching	✓	✗
Self-subtraction	✗	✓

Target vs. Reference



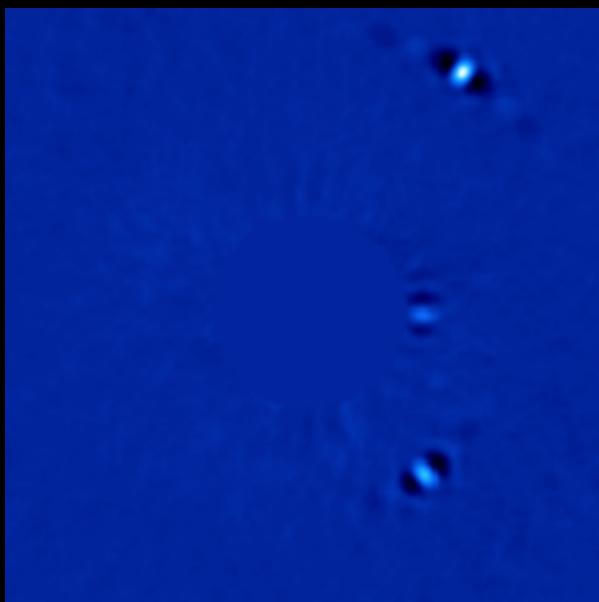
	Target-based	Reference-based	
WFE-matching	✓	✗	$d < 1\text{--}5^\circ$
λ -matching	✓	✗	Color match
Self-subtraction	✗	✓	

$d < 1\text{--}5^\circ$
Color match

For Planet searches

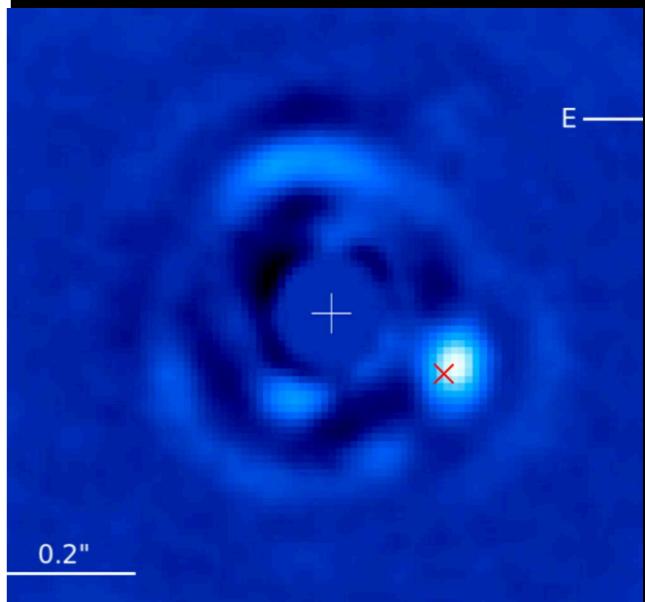
Access to small separations

ADI

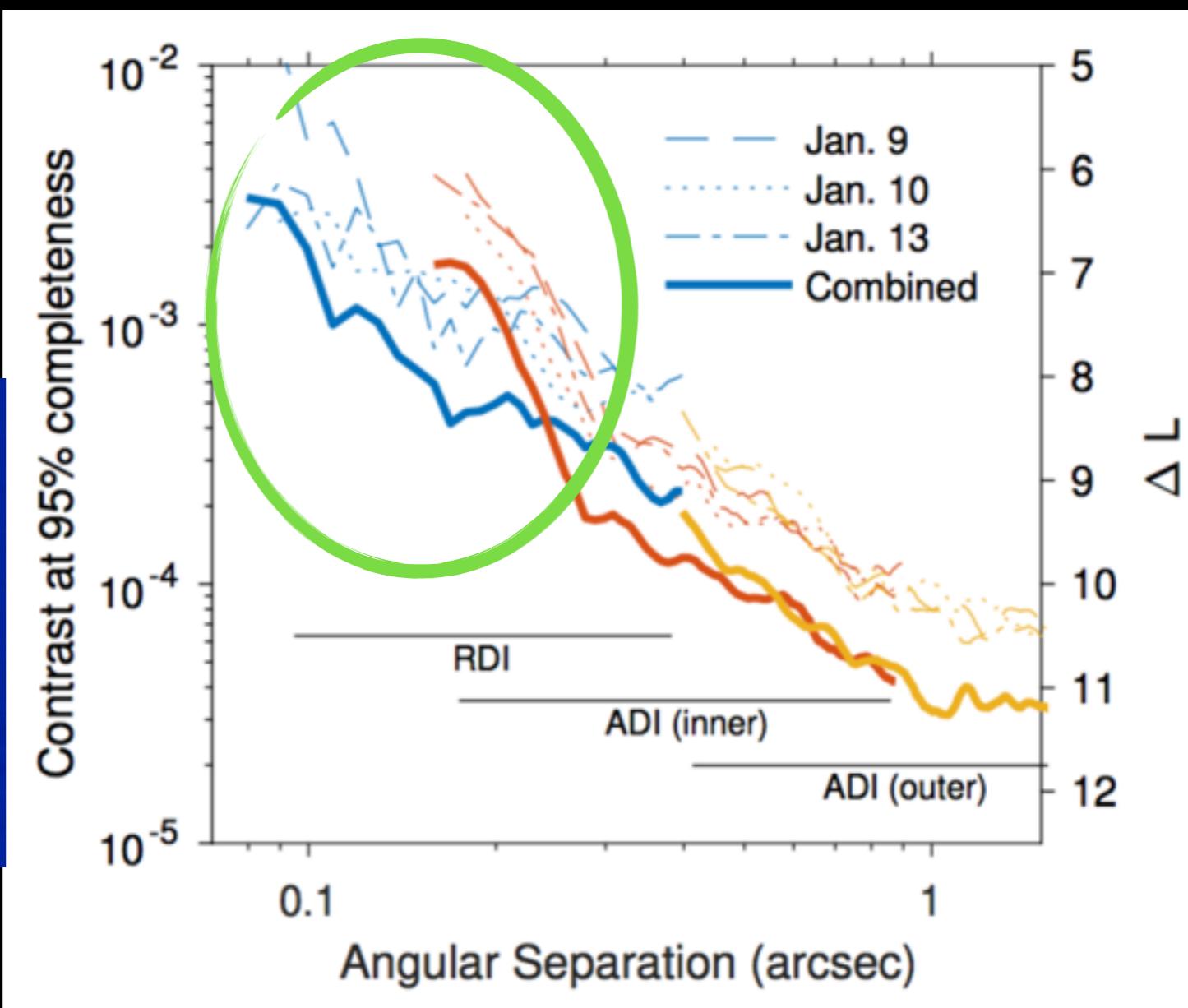


Zurlo et al. 2016

RDI



Serabyn et al. 2017

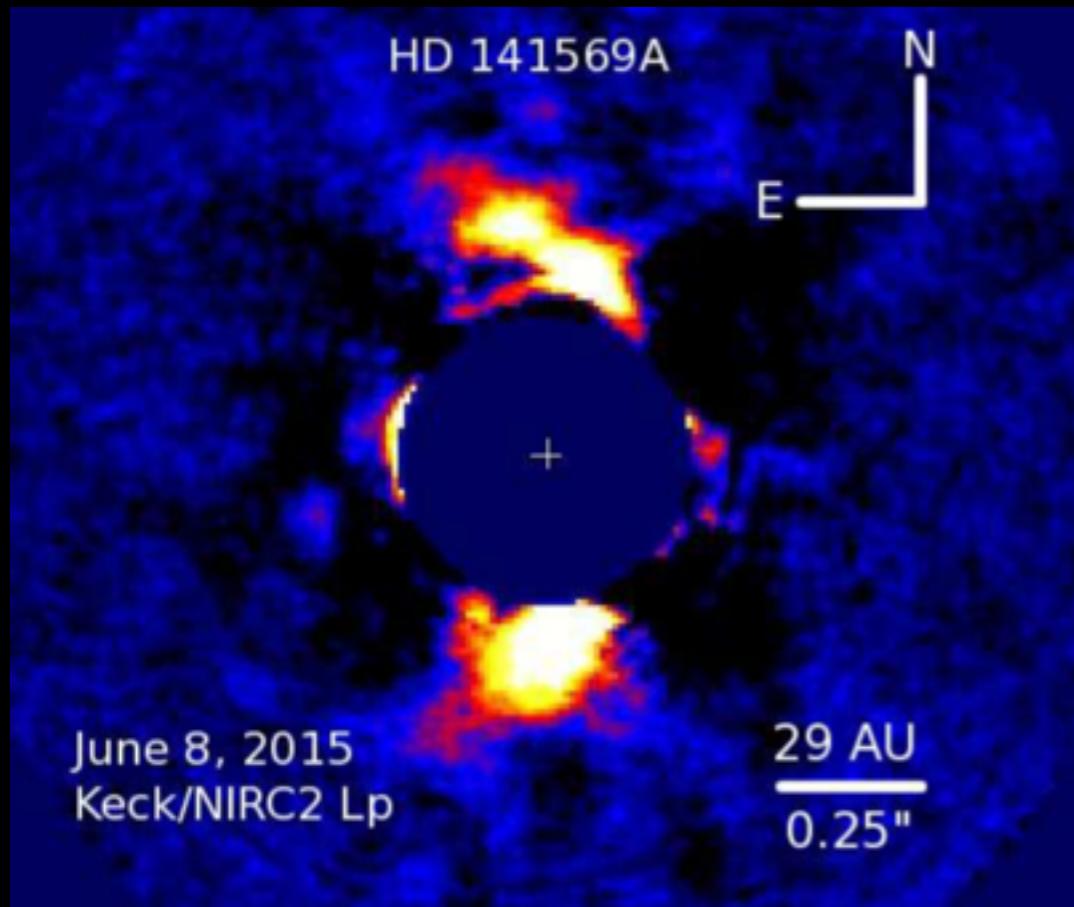


Ruane et al. 2017

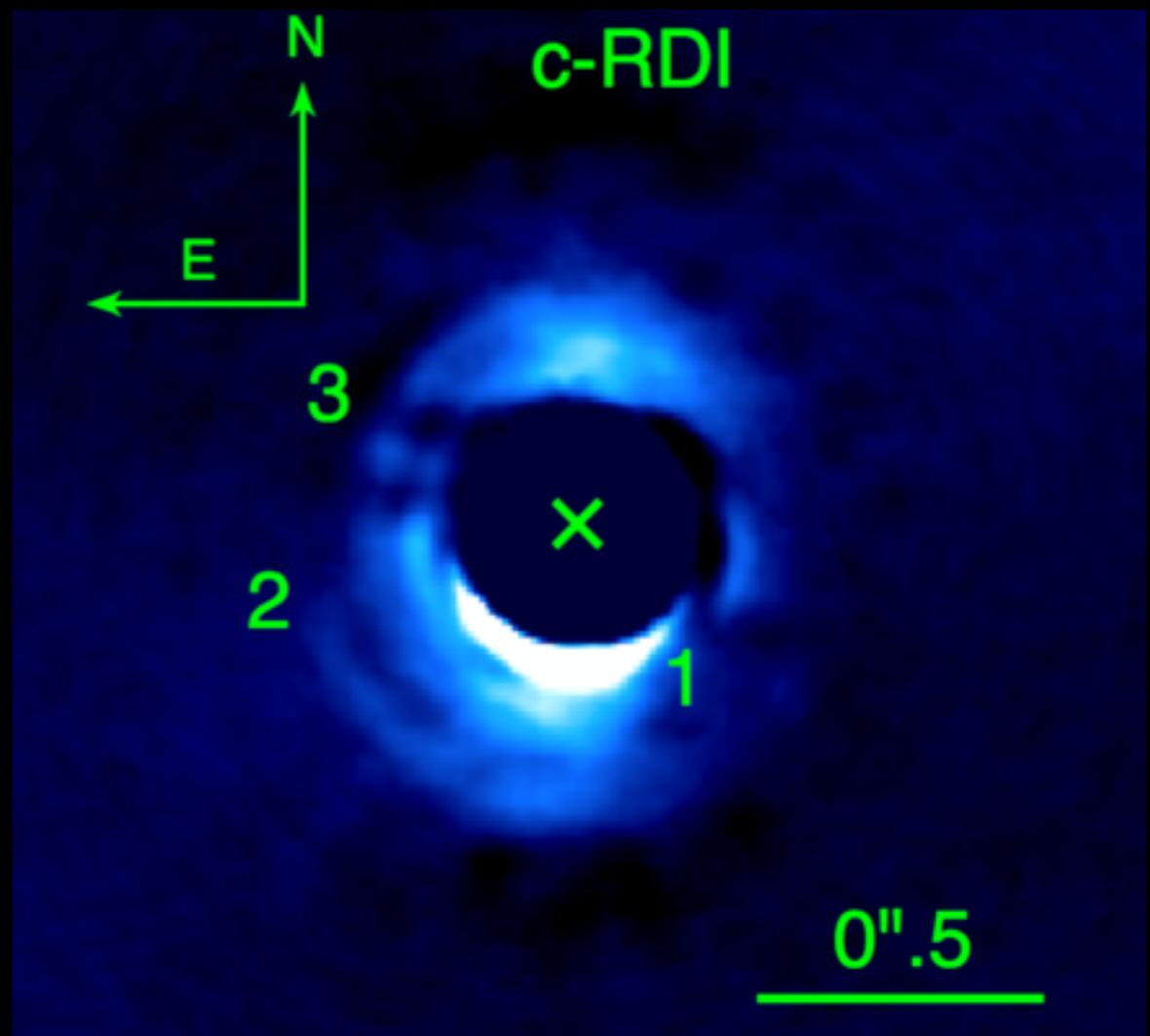
For Disk science

Unbiased morphology

ADI



RDI



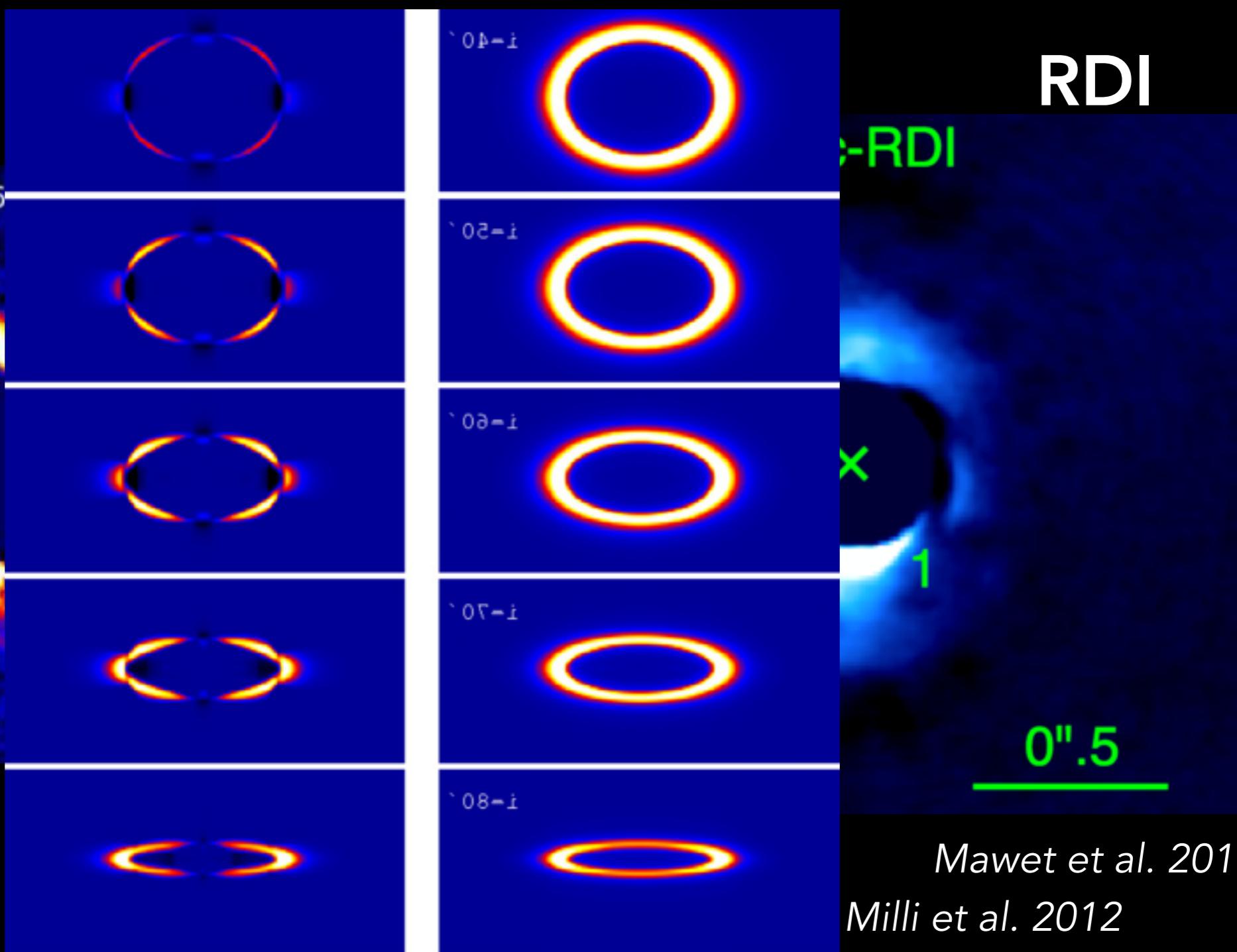
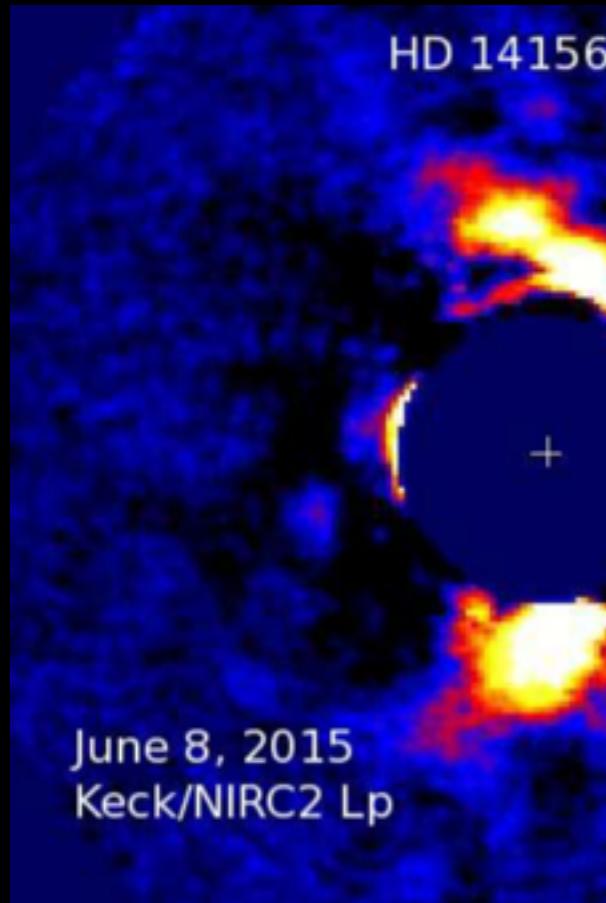
Currie et al. 2016

Mawet et al. 2017

For Disk science

Unbiased morphology

ADI

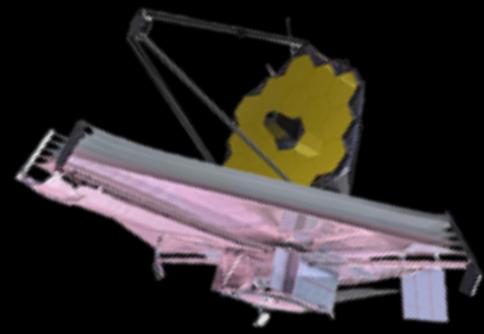


Currie et al. 2016

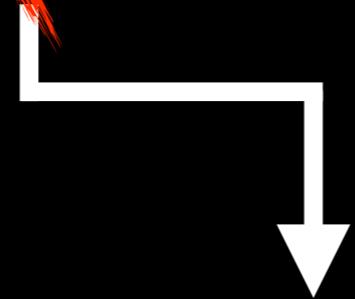
Mawet et al. 2017
Milli et al. 2012



Space Telescopes



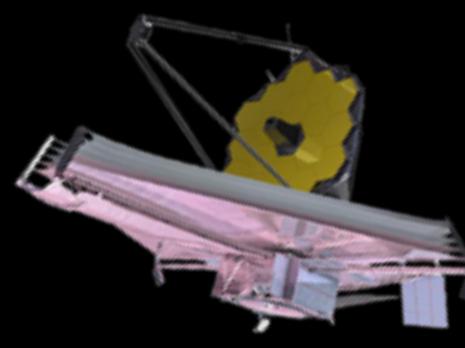
ADI
SDI
PDI



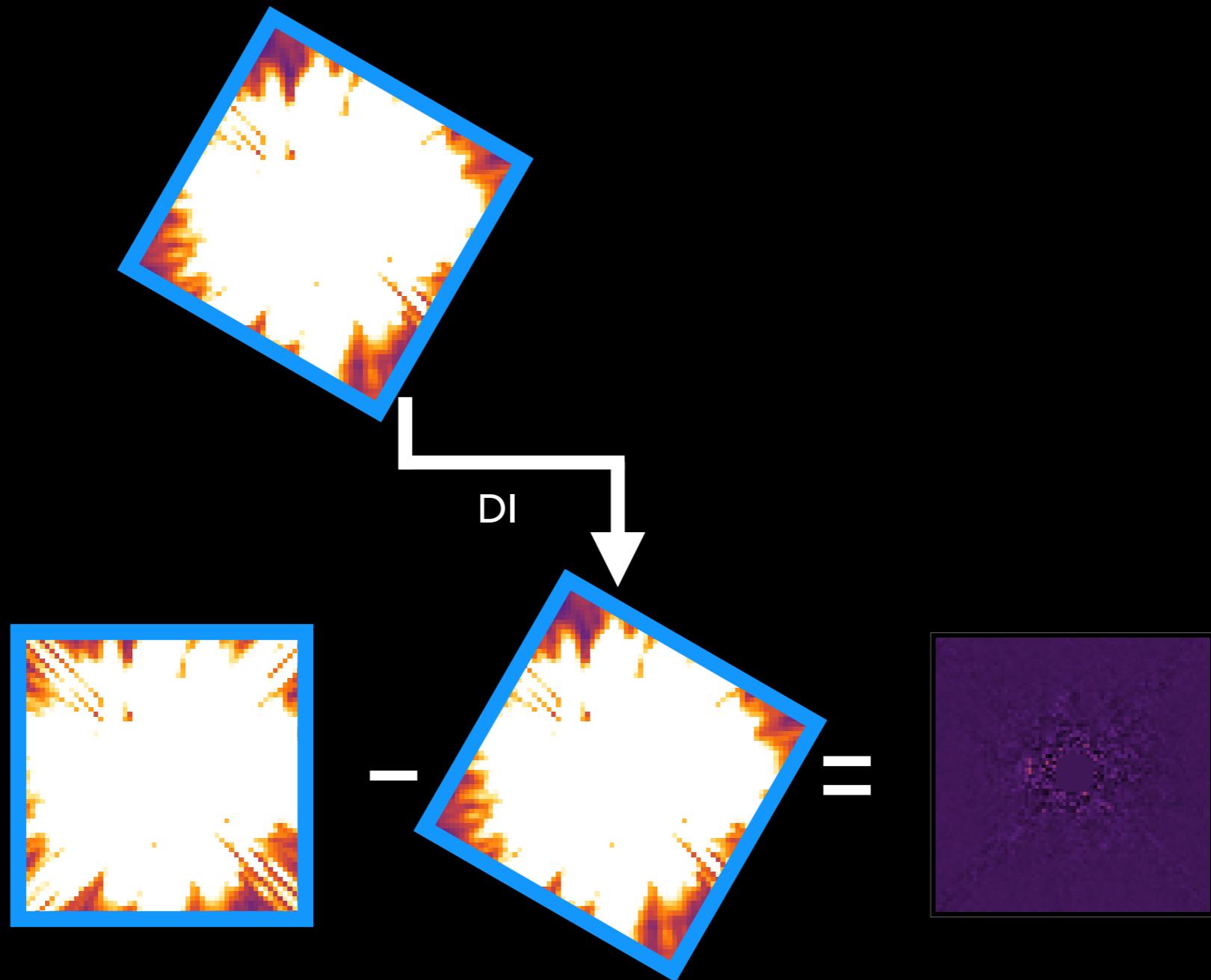
$$\boxed{\text{[Image]}} - \boxed{\text{[Image]}} = \boxed{\text{[Residual Image]}}$$



Space Telescopes

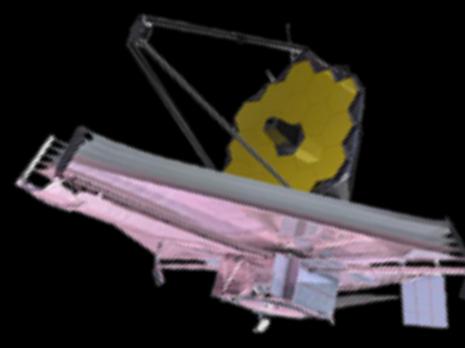


Rolls

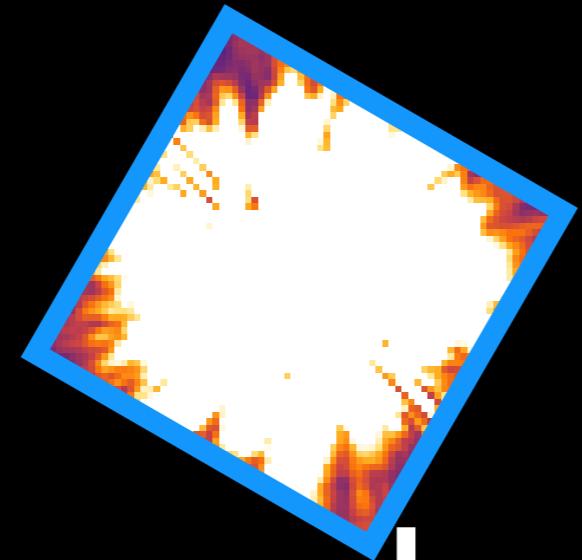




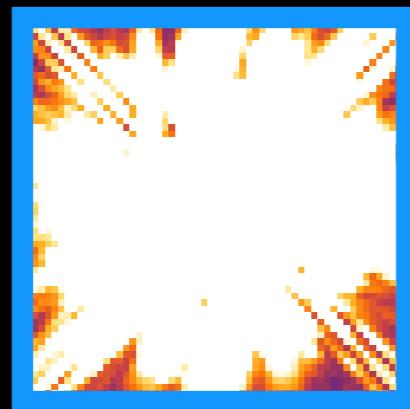
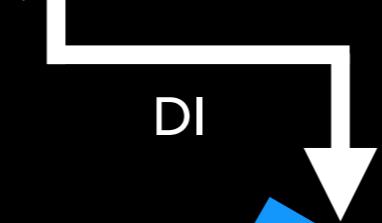
Space Telescopes



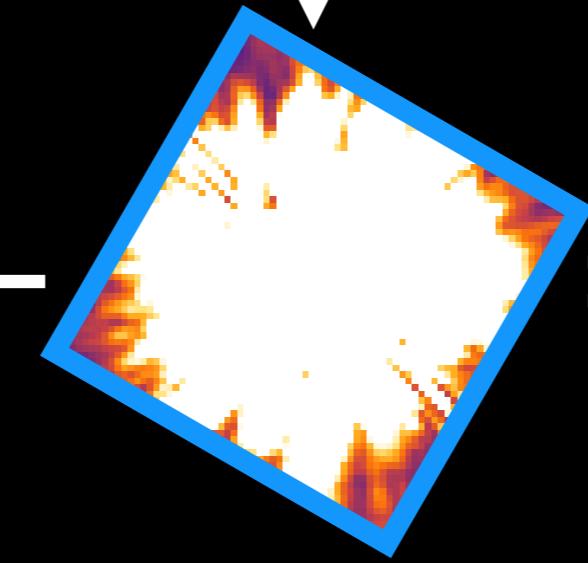
Rolls



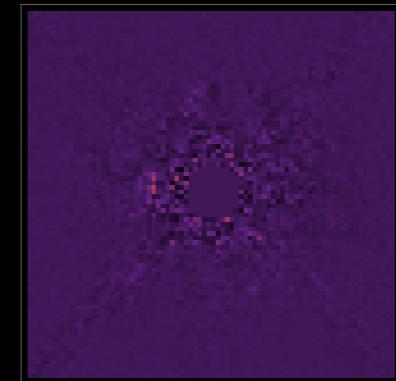
DI



-



=

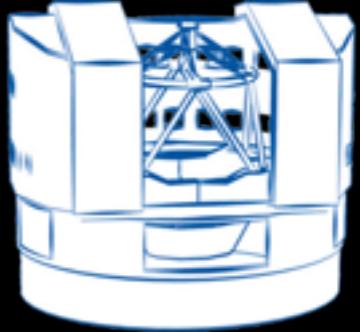


Discontinuous (new acc)

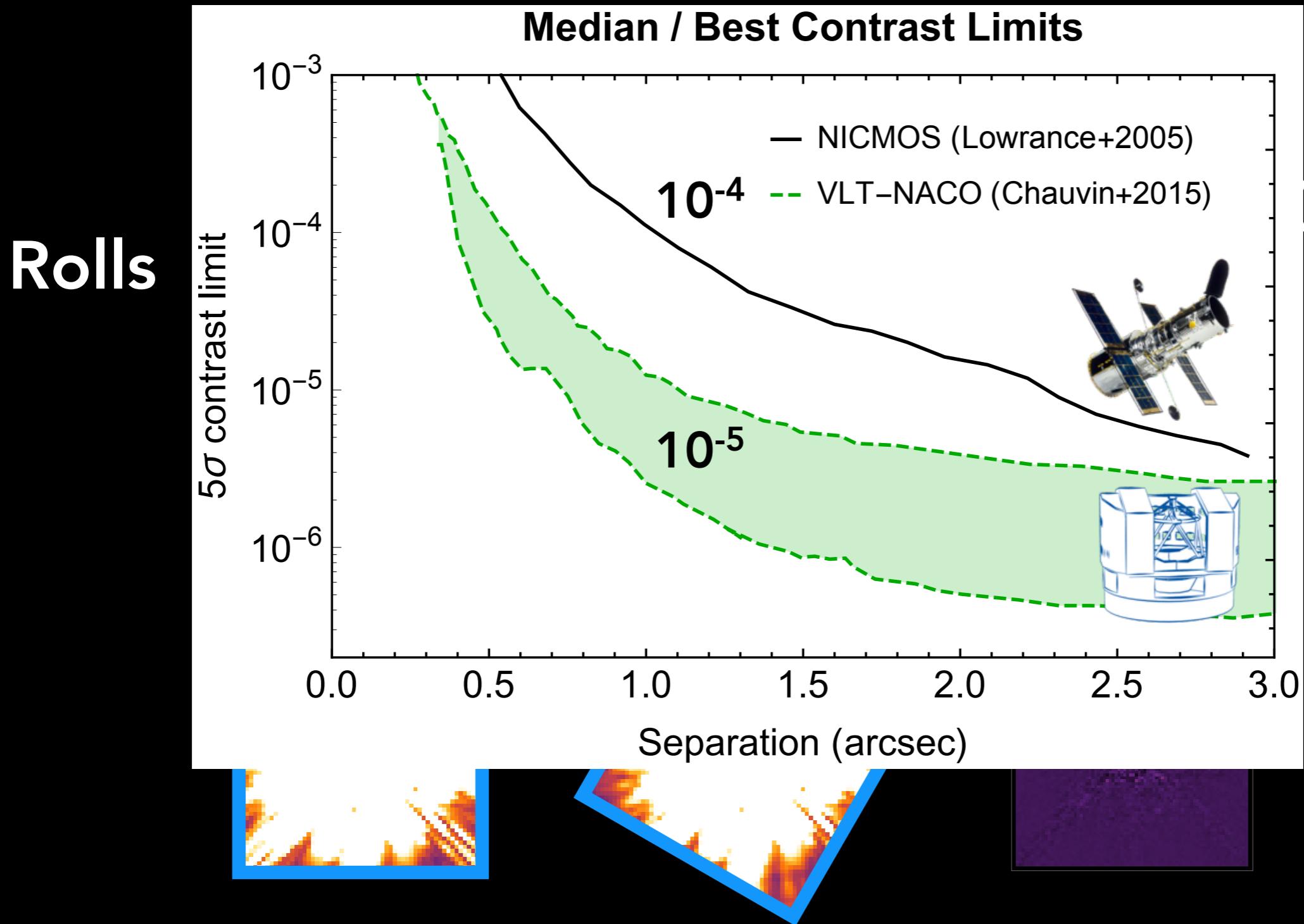
Limited range:

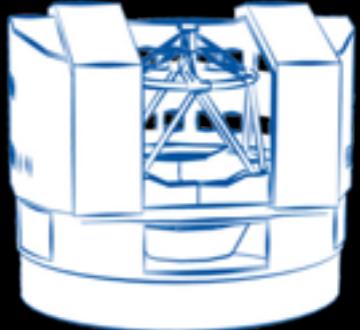
- HST: 30°
- JWST: 10°

But great PSF stability
(no atmosphere, no AO)

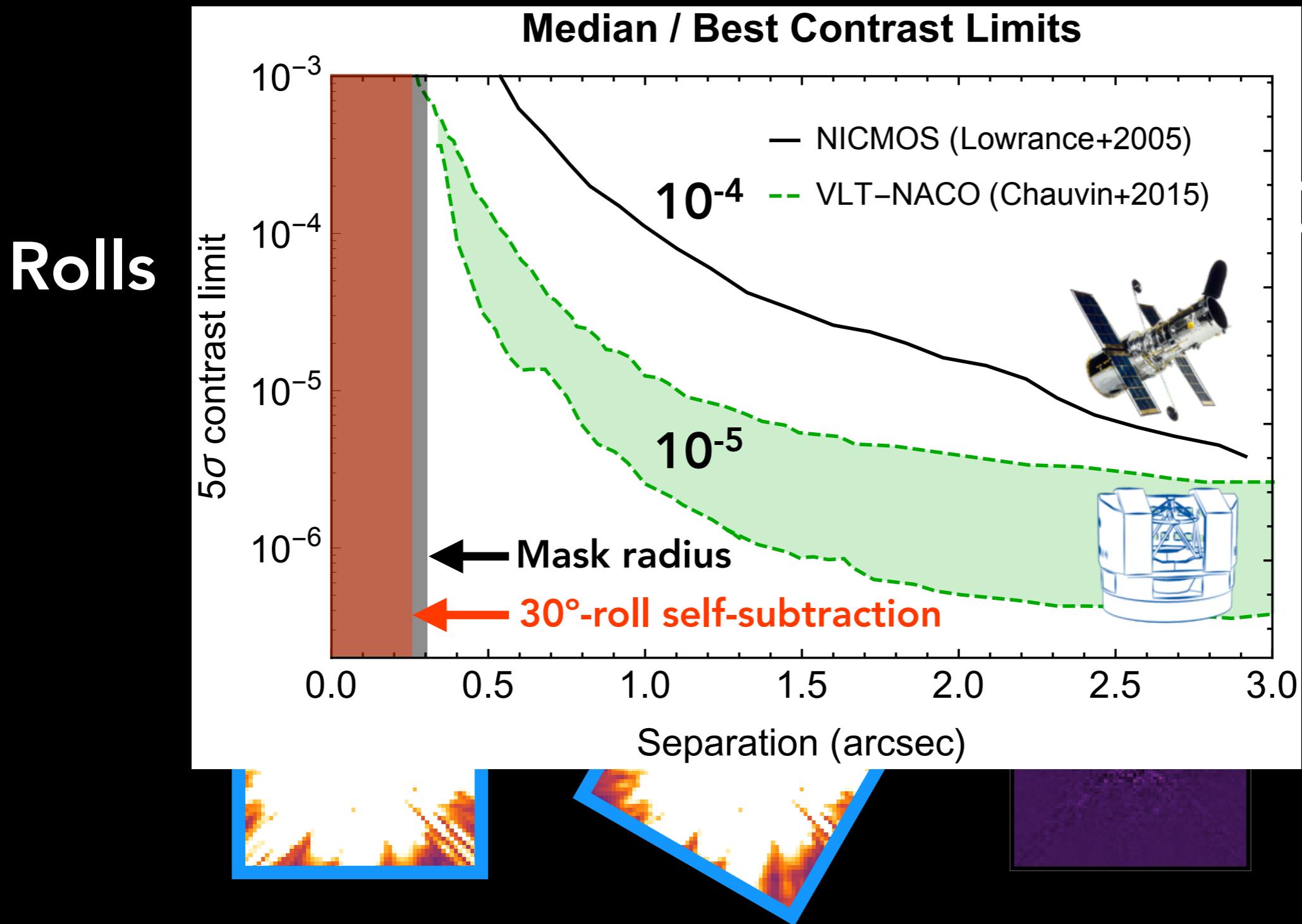


HST vs Ground



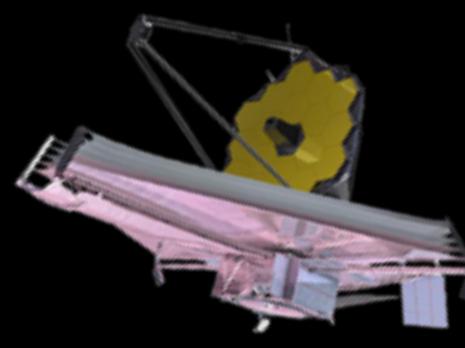


HST vs Ground



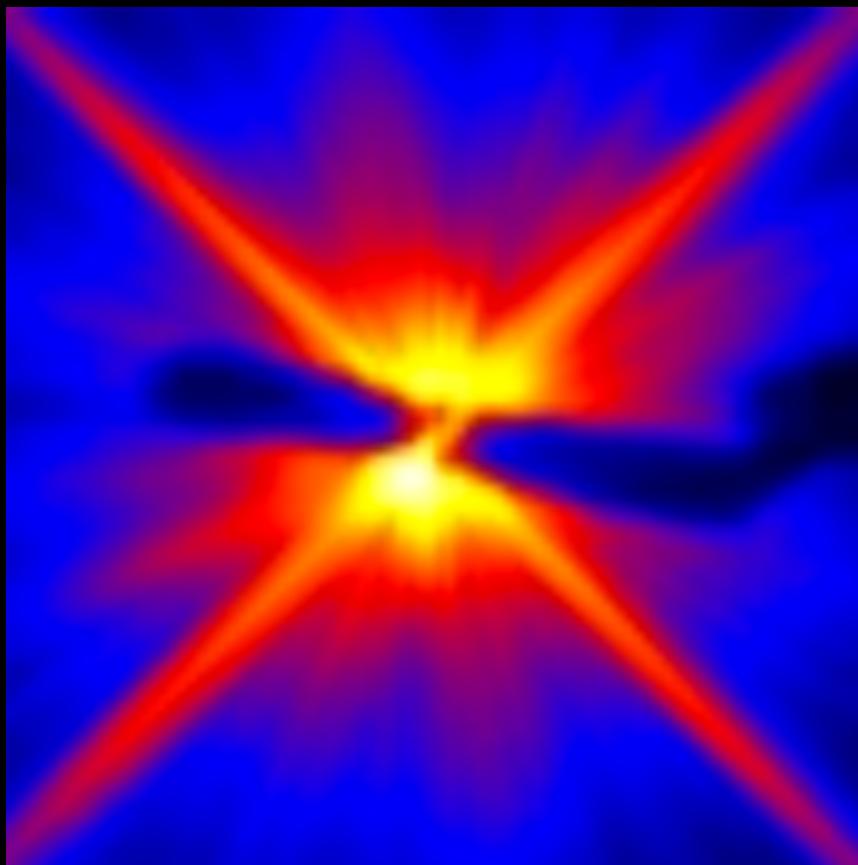


Space Telescopes



Small IWA coronagraphs
(or telescopes with limited roll-angle)

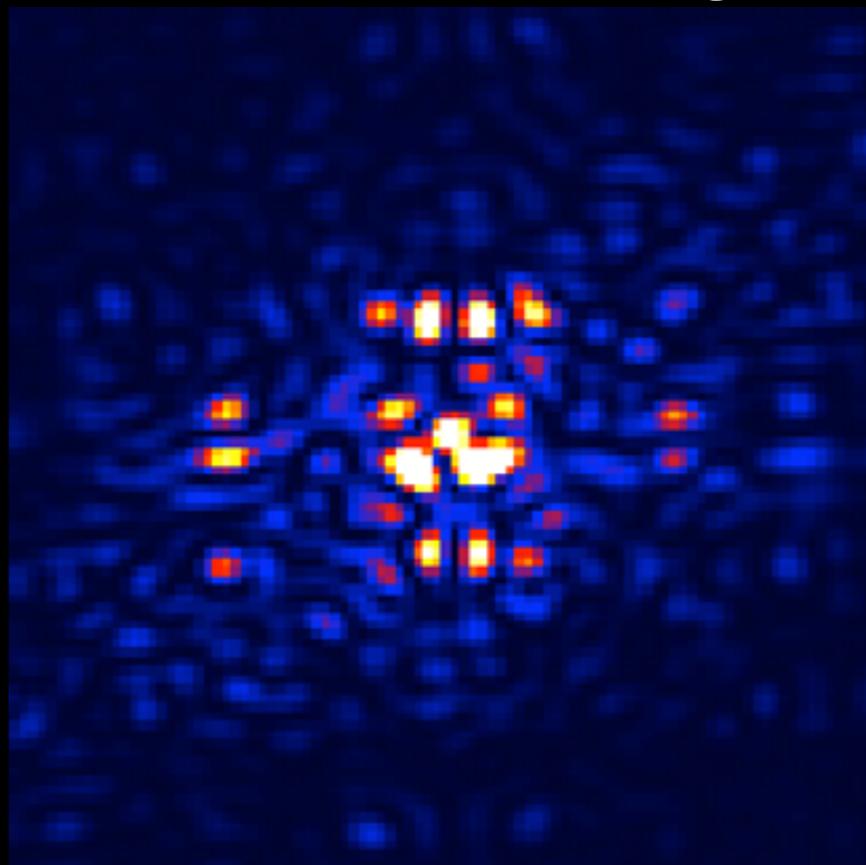
STIS : BAR5

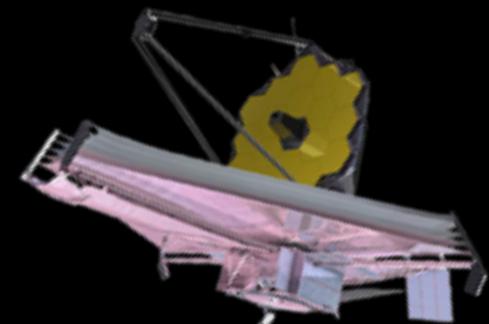


RDI needed

... but limited by
pointing errors

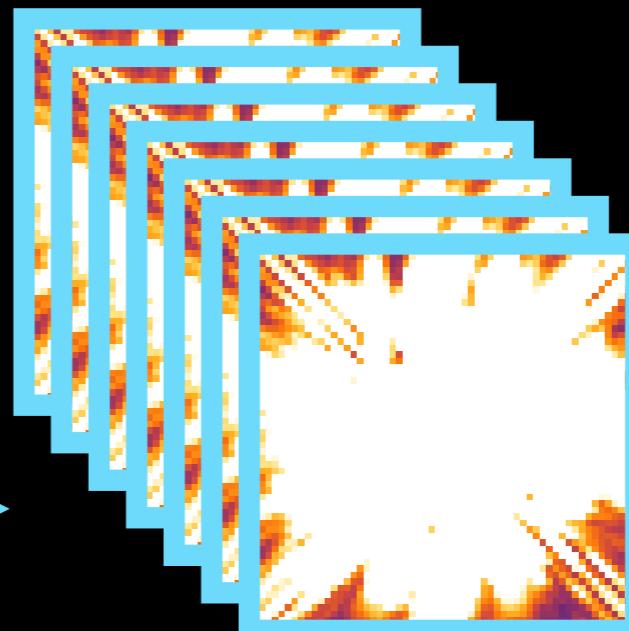
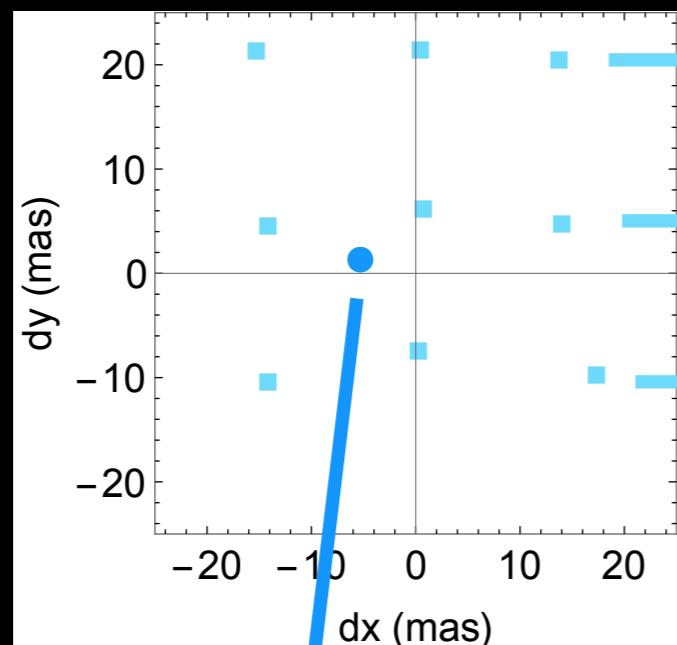
MIRI : 4QPMs
NIRCam : wedge





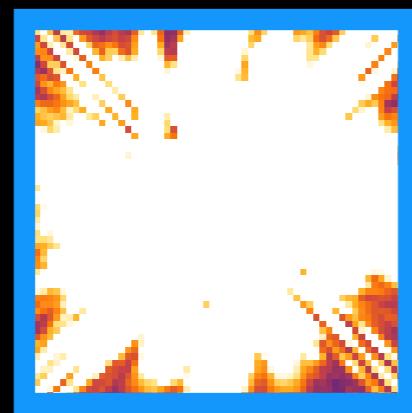
RDI with dithers

Small-grid dither

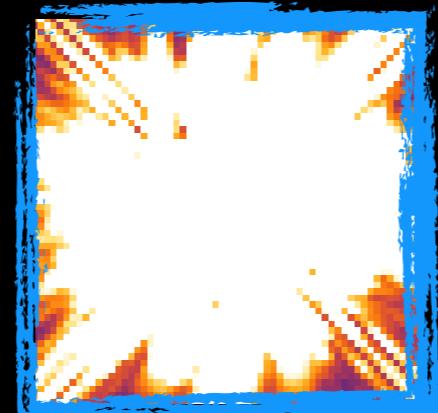


RDI

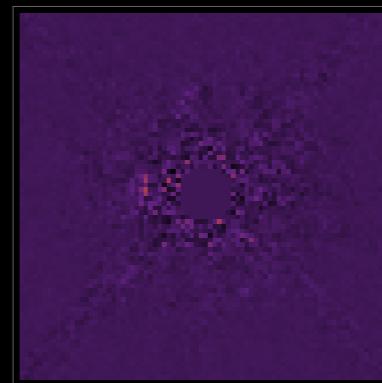
PCA



-



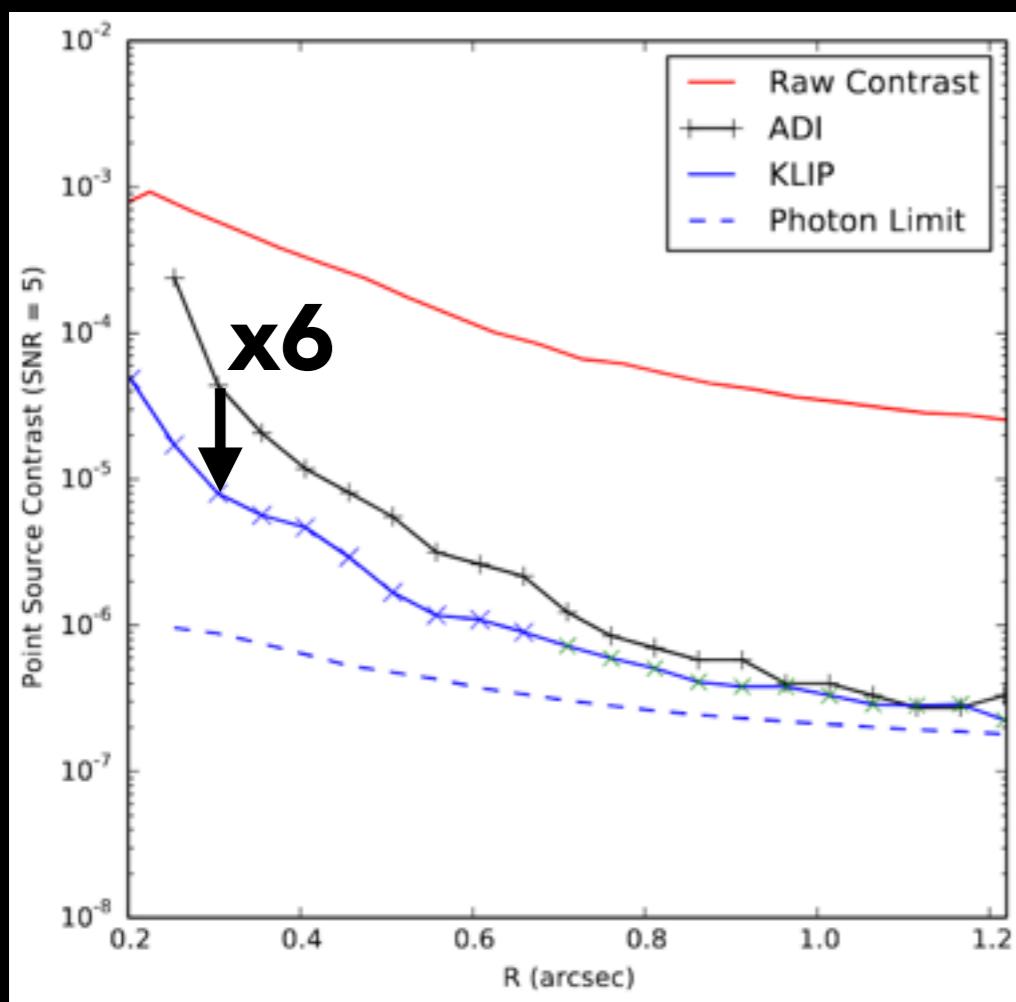
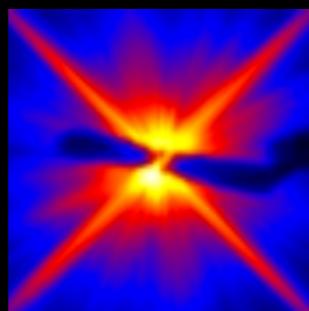
=



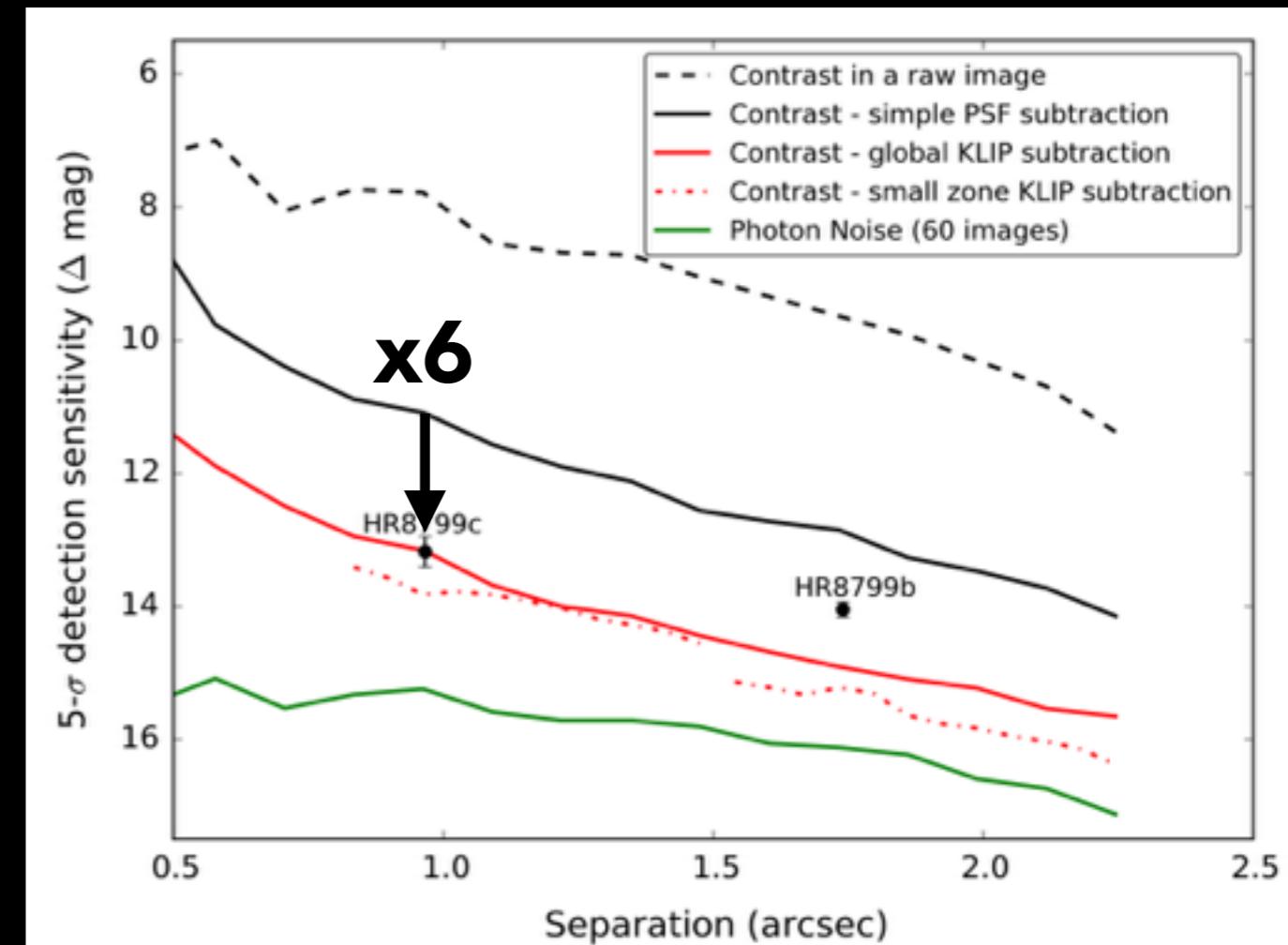


RDI with dithers

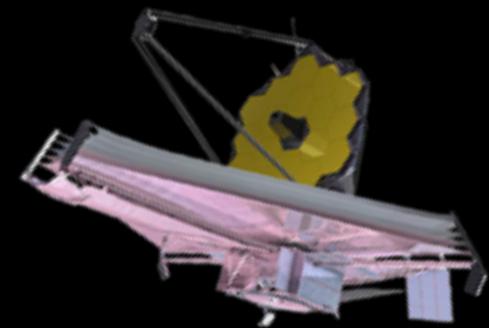
STIS : BAR5



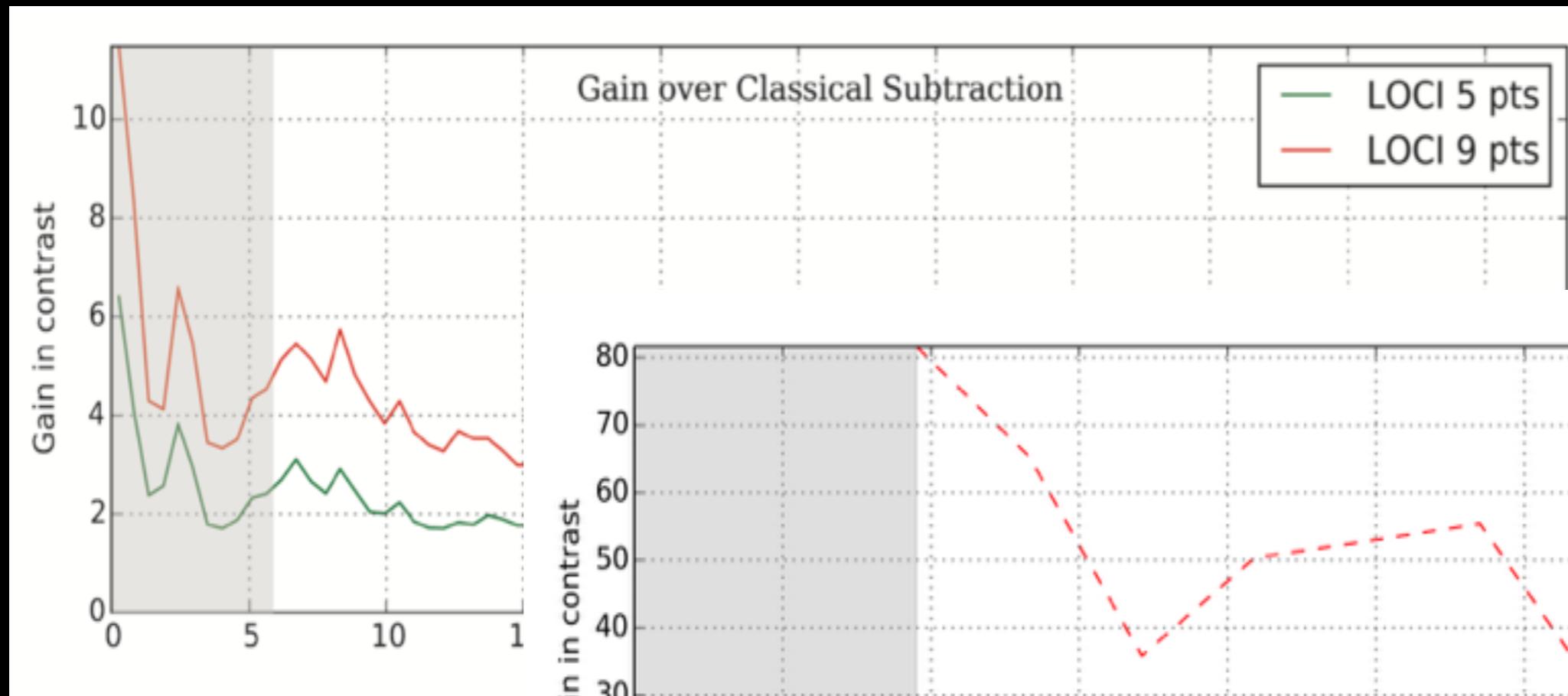
WFC3 : No coronagraph



RDI with dithers

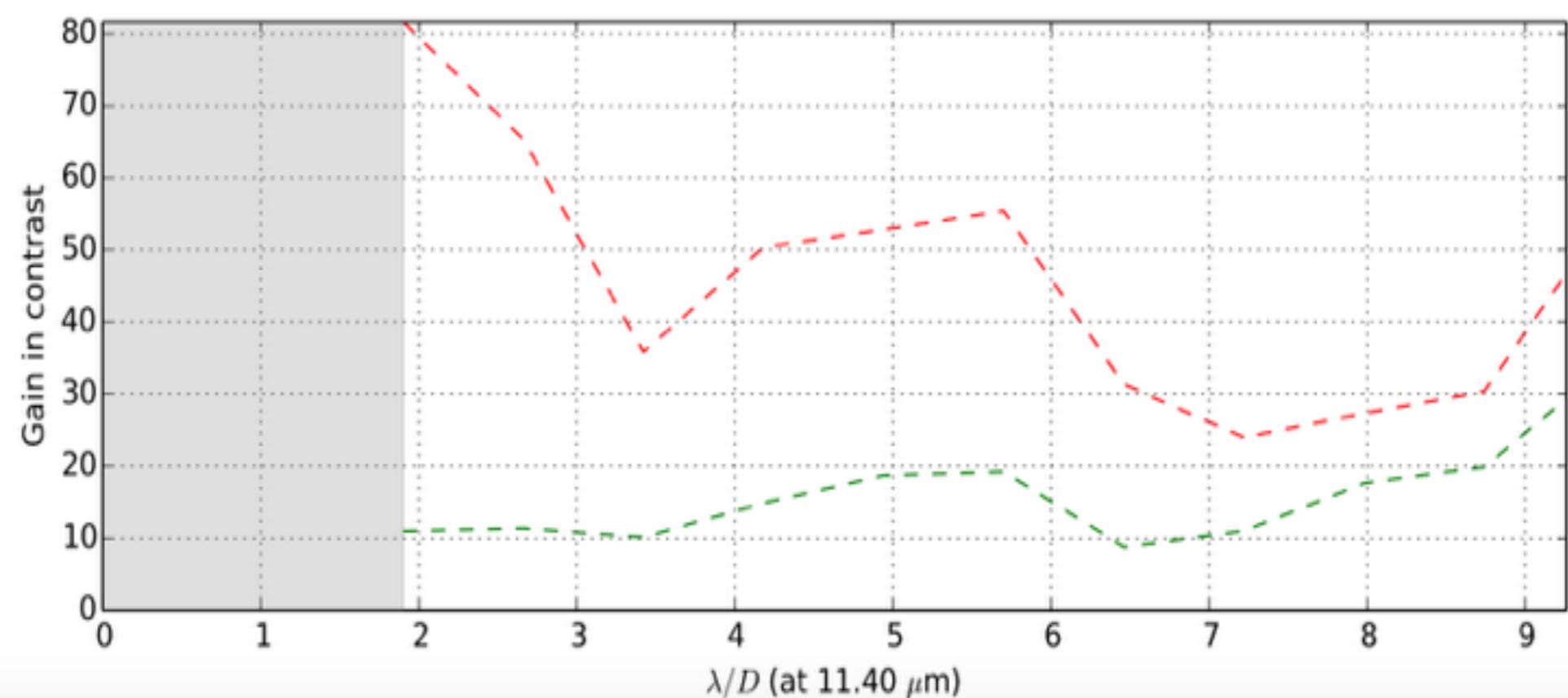


NIRCam



Lajoie et al. 2016

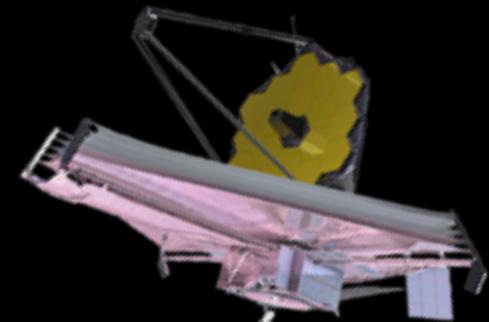
MIRI



Soummer et al. 2016

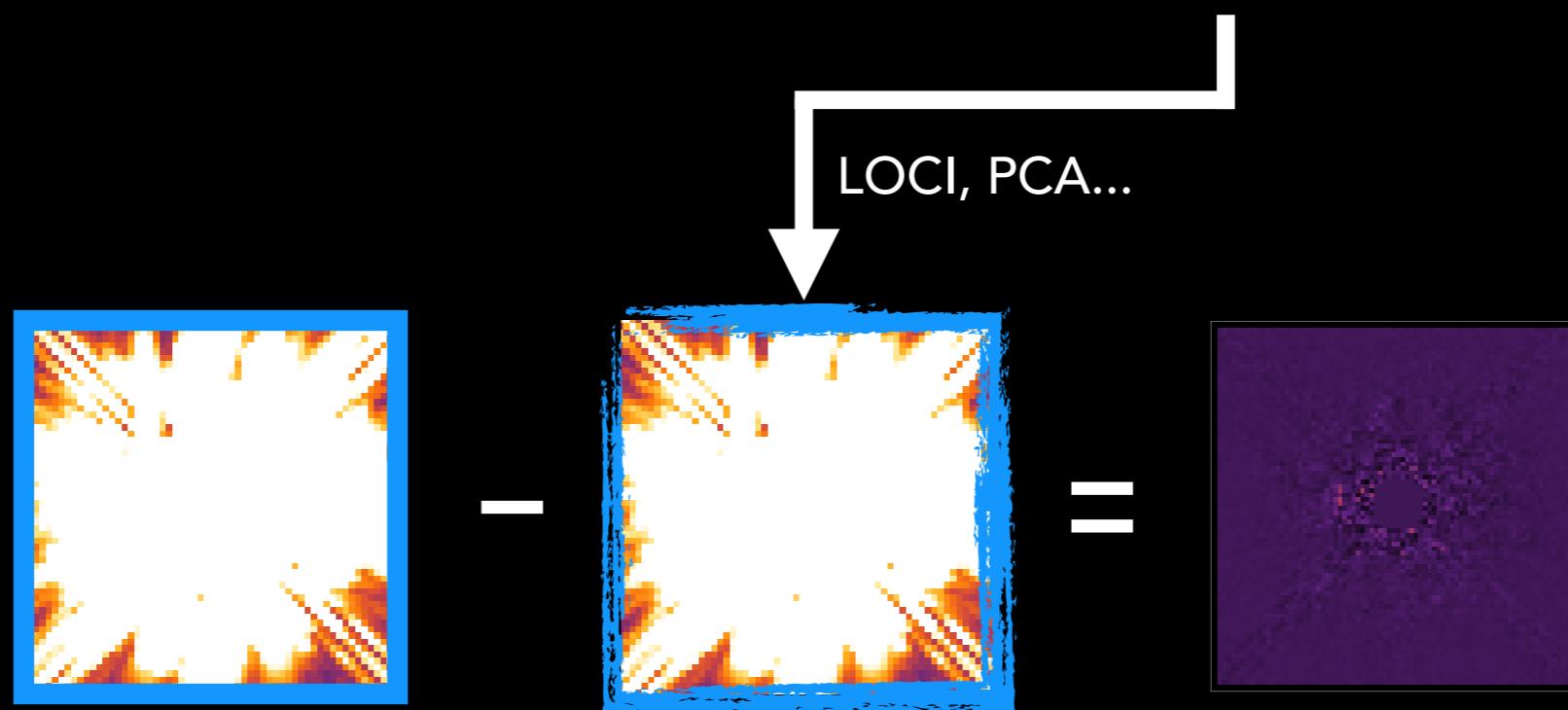
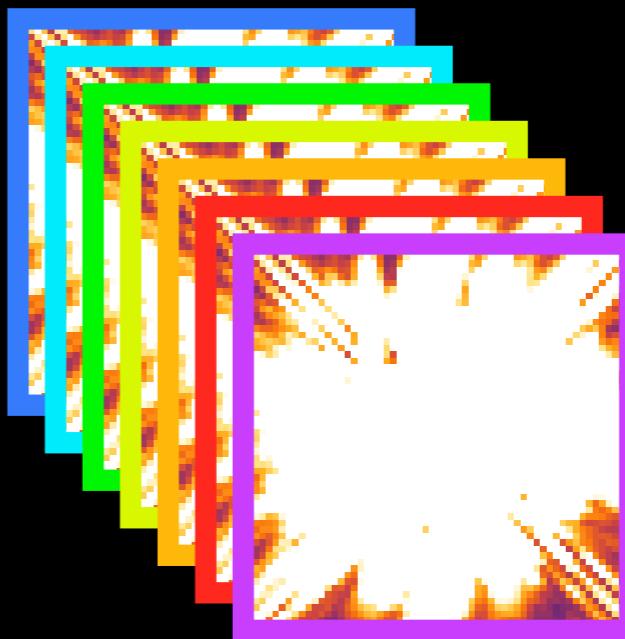


MRDI



Multiple Reference stars

- *Self-referencing surveys*
- *Rich archives*



LOCI, PCA...



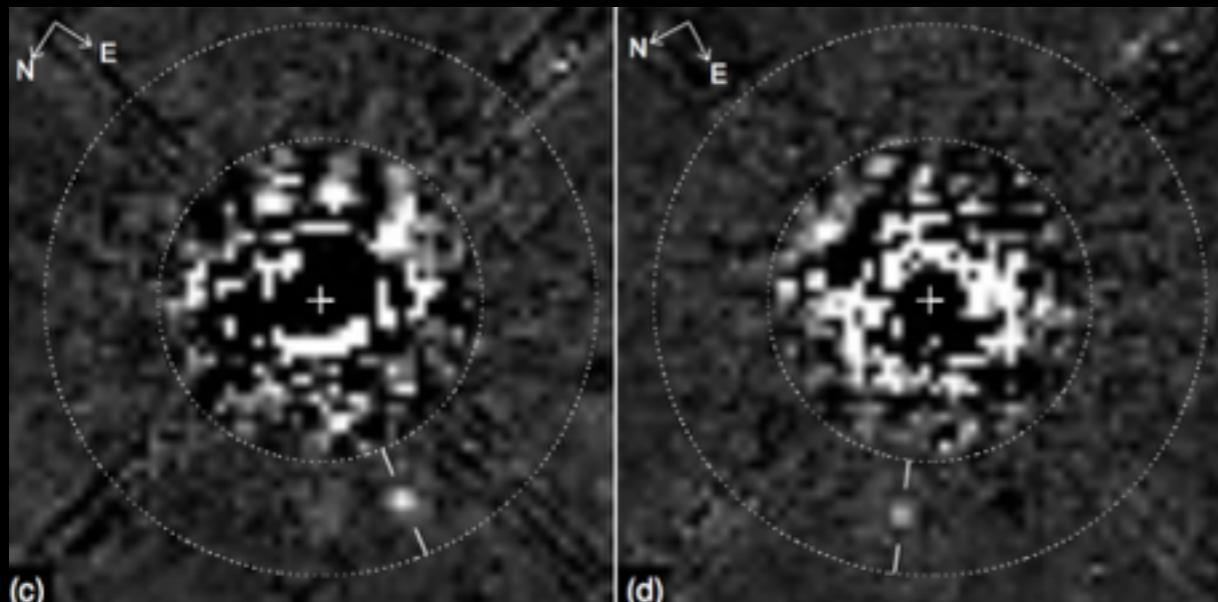
MRDI

Self-referencing surveys:

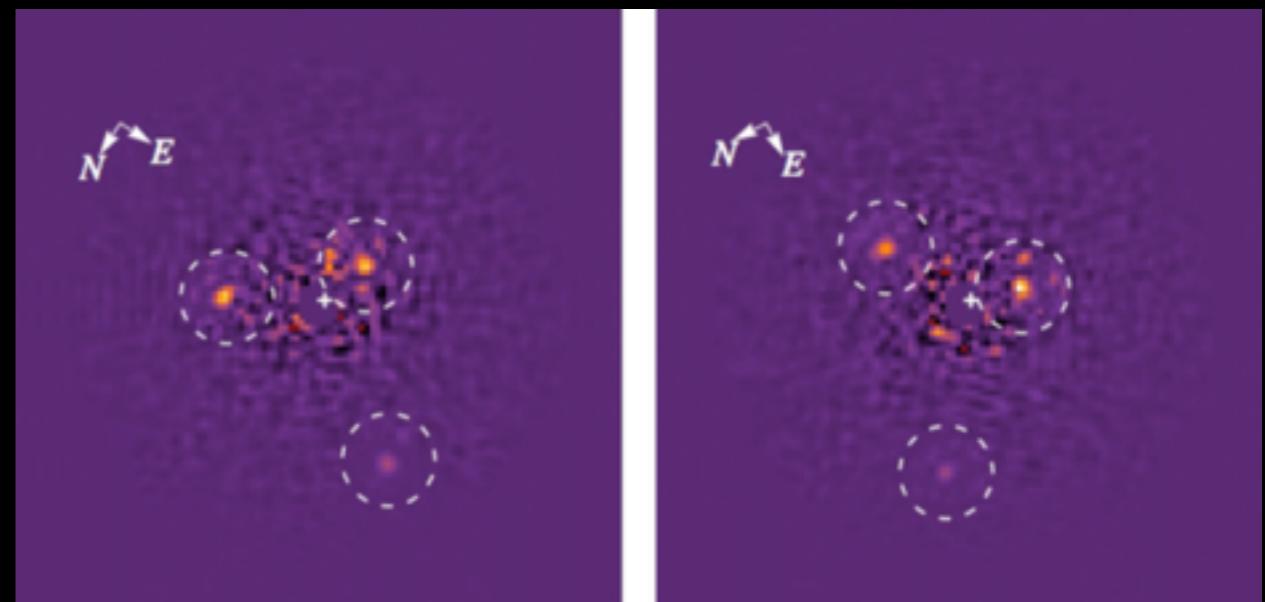
- ▶ *Consistent data sets*
- ▶ *Same noise levels*

HR 8799:

1998 NICMOS survey



Lafrenière et al. 2009



Soummer et al. 2011

Images	Ref Stars
203	23



MRDI

NICMOS archives

- ▶ *Reference stars only*
- ▶ *Frame selection (50-90%)*

F110W (~ J band)

	Images	Ref Stars
N2 Cooling	54	7
Cryo-cooling	655	74

F160W (H band)

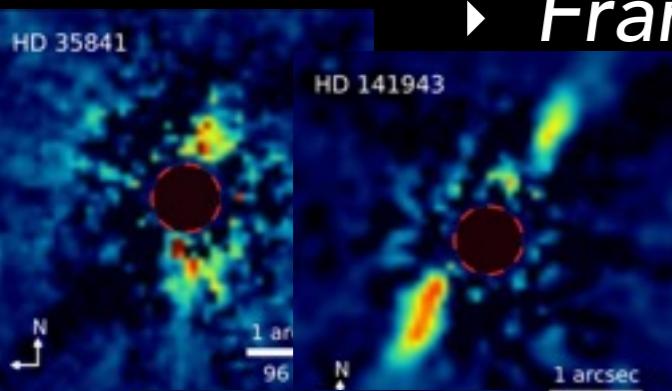
	Images	Ref Stars
N2 Cooling	360	55
Cryo-cooling	809	66



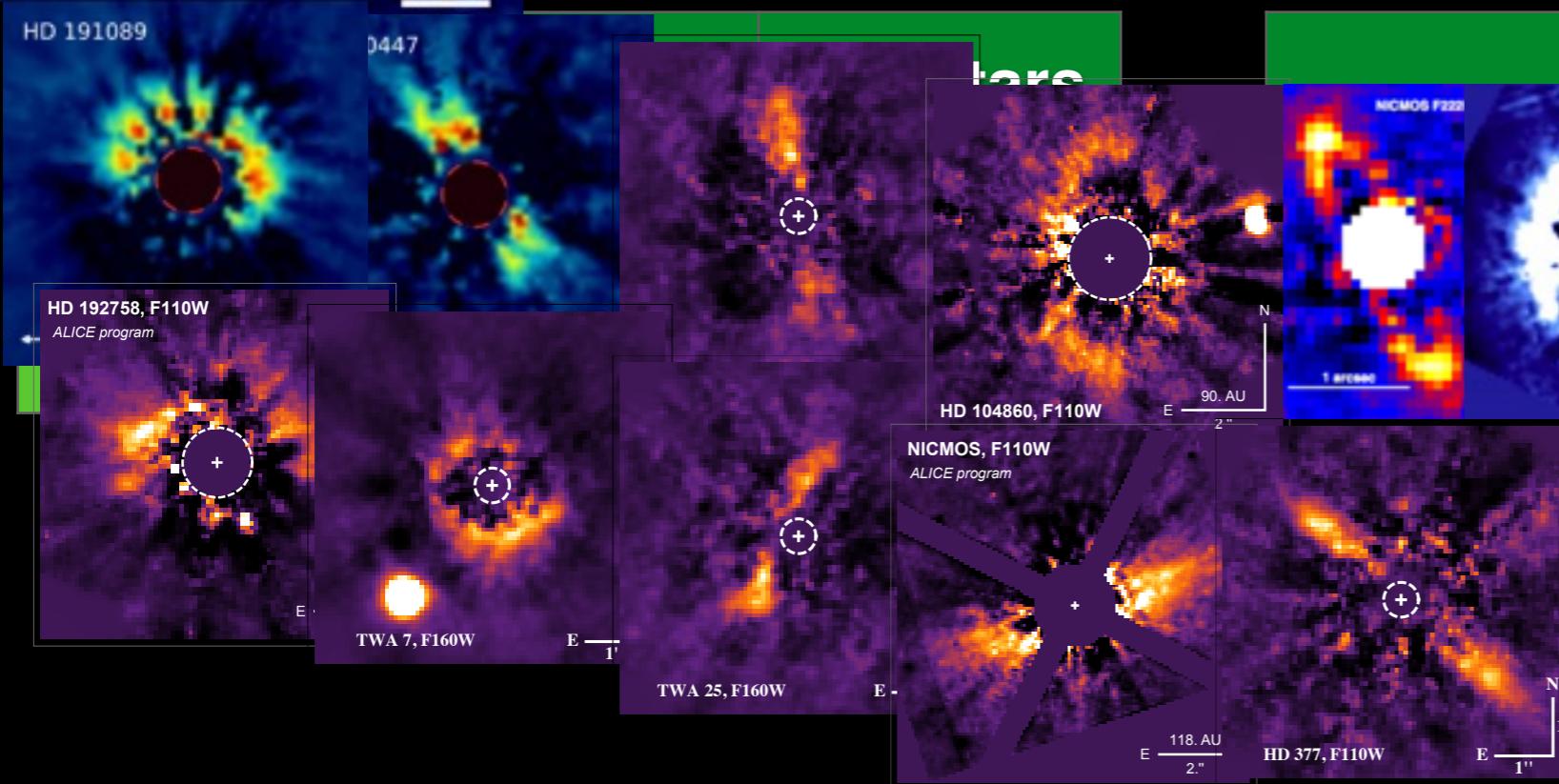
MRDI

NICMOS archives

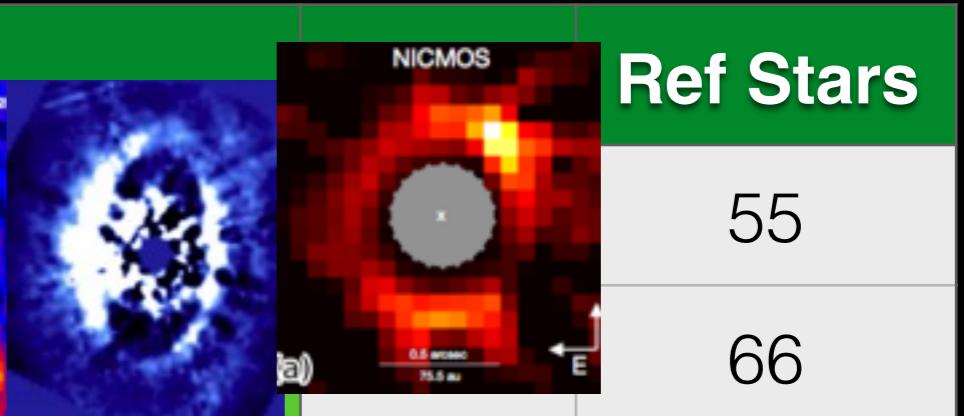
- ▶ Reference stars only
- ▶ Frame selection (50-90%)



(~ J band)



F160W (H band)



Soummier et al. 2014

Choquet et al. 2016, 2017, 2018

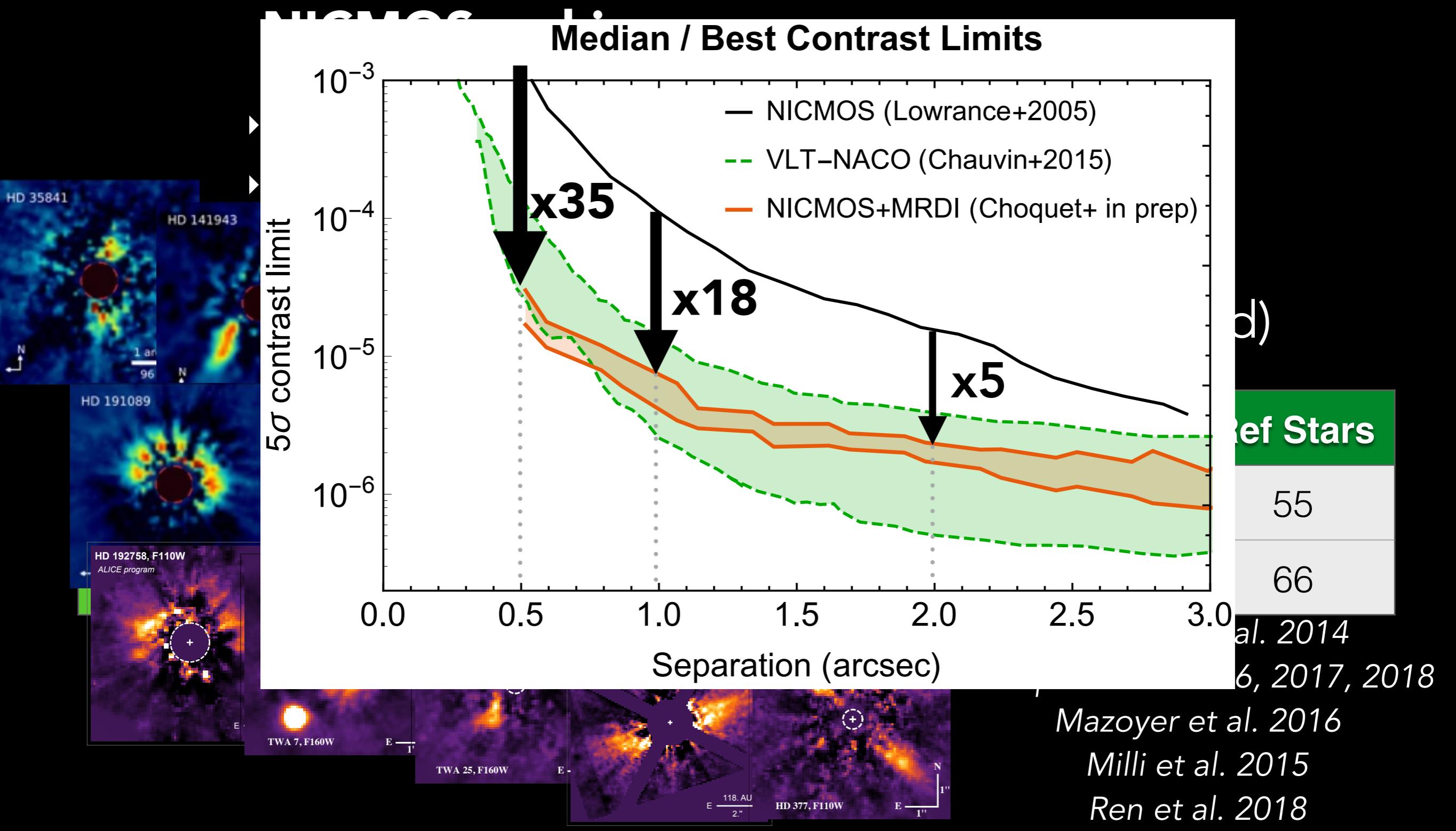
Mazoyer et al. 2016

Milli et al. 2015

Ren et al. 2018



MRDI

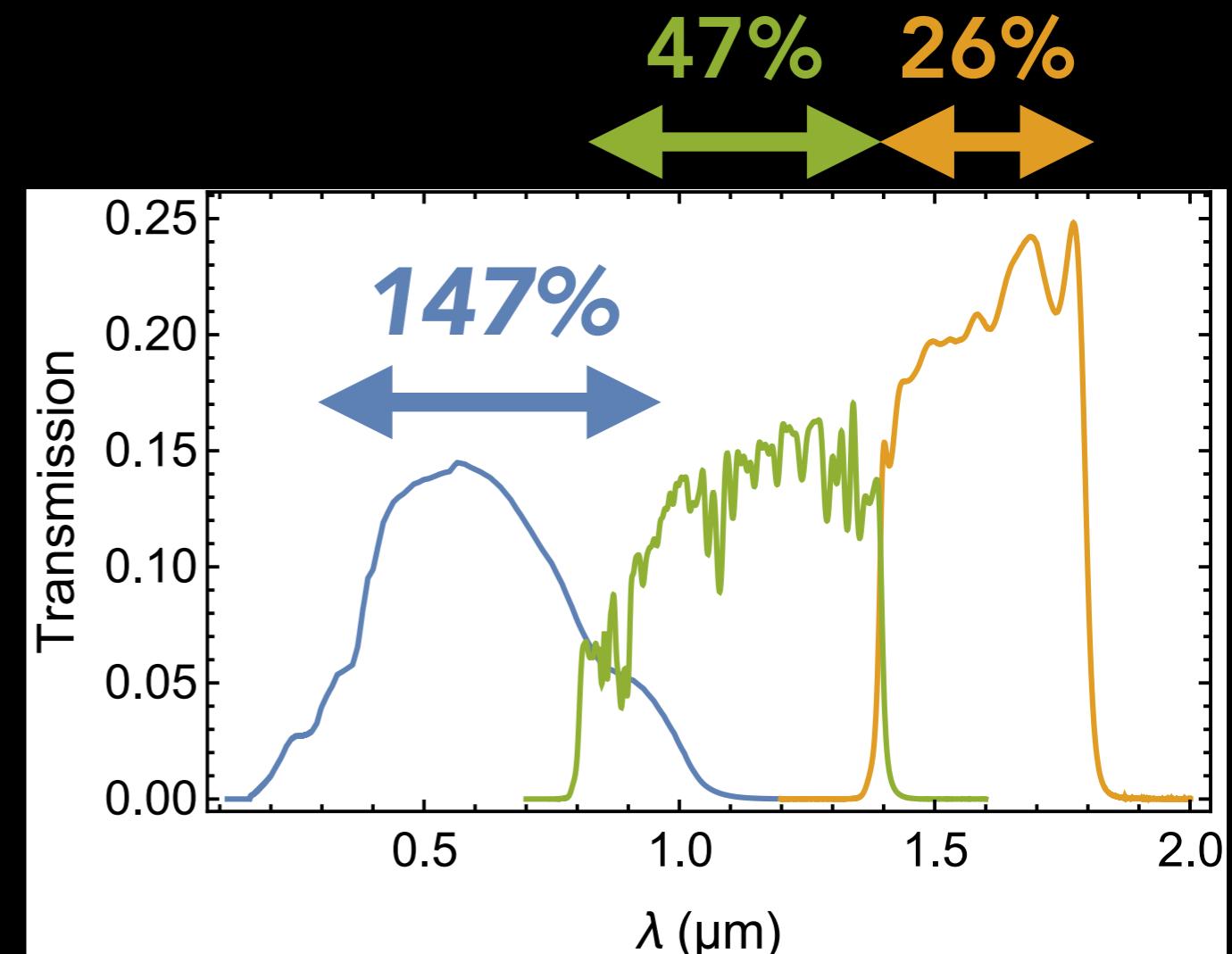
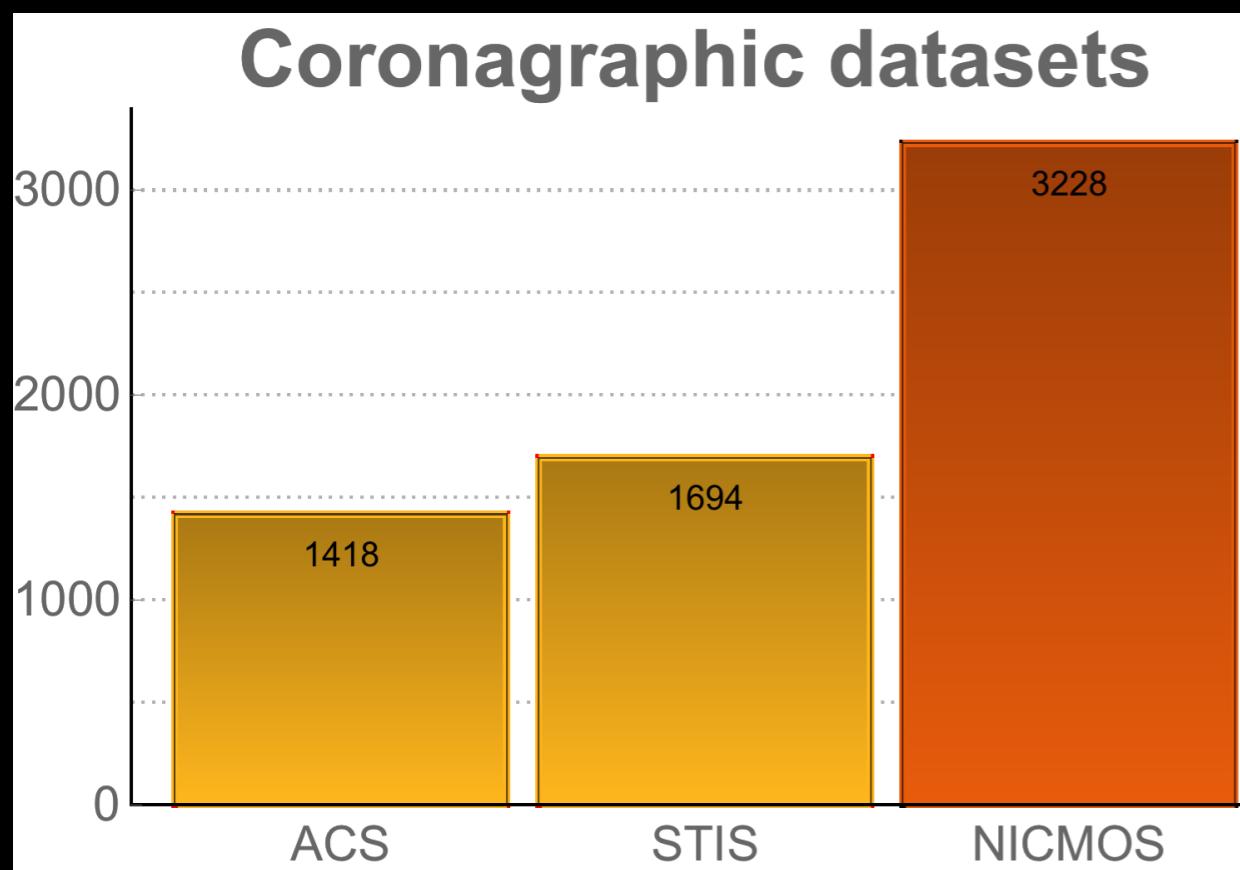




MRDI

STIS archives

Coronagraphic datasets

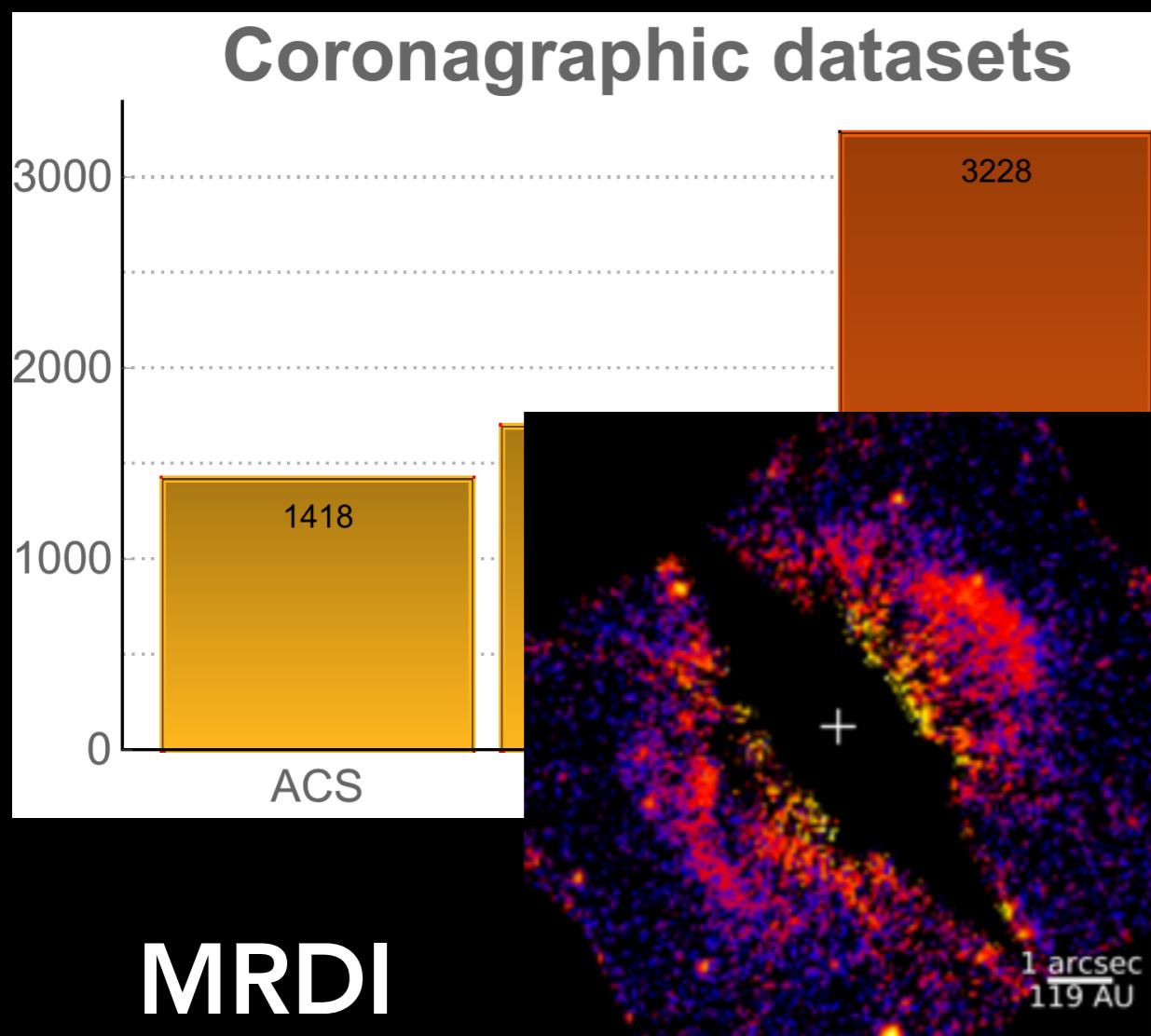




MRDI

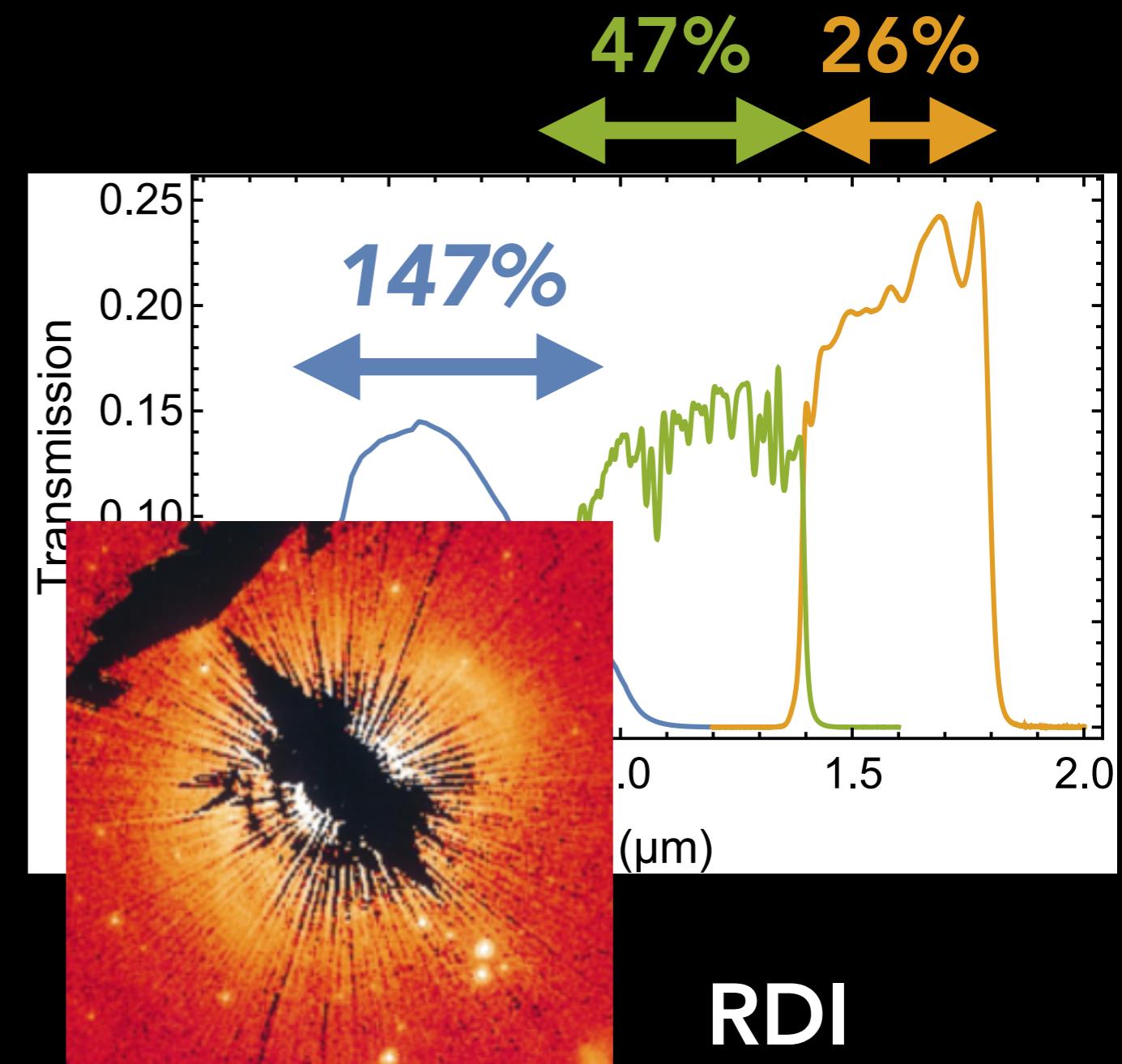
STIS archives

Coronagraphic datasets



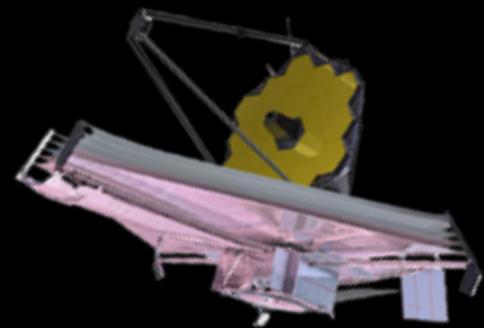
MRDI

Ren et al. 2017



RDI

Ren et al. 2017



Summary

Importance of (M)RDI :

- ✓ Access to small separations
- ✓ Un-biased morphologies & photometries
- ➡ Better sensitivity limits

MRDI + Space Tel.

Needs :

- ✓ Good matching reference stars
- ✓ Diversity (dithering, archives)