

The Challenge of Interpreting Direct Detection Data

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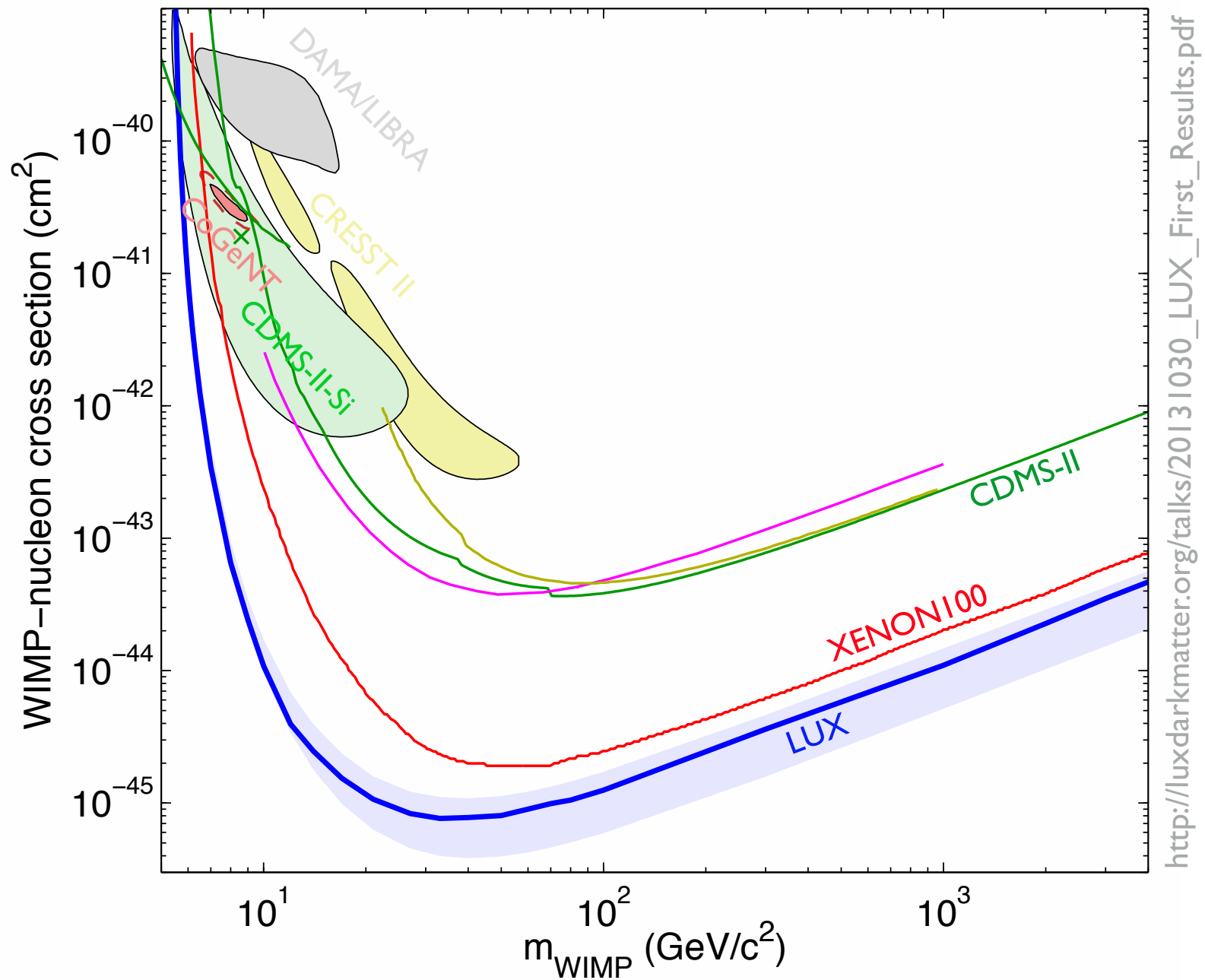
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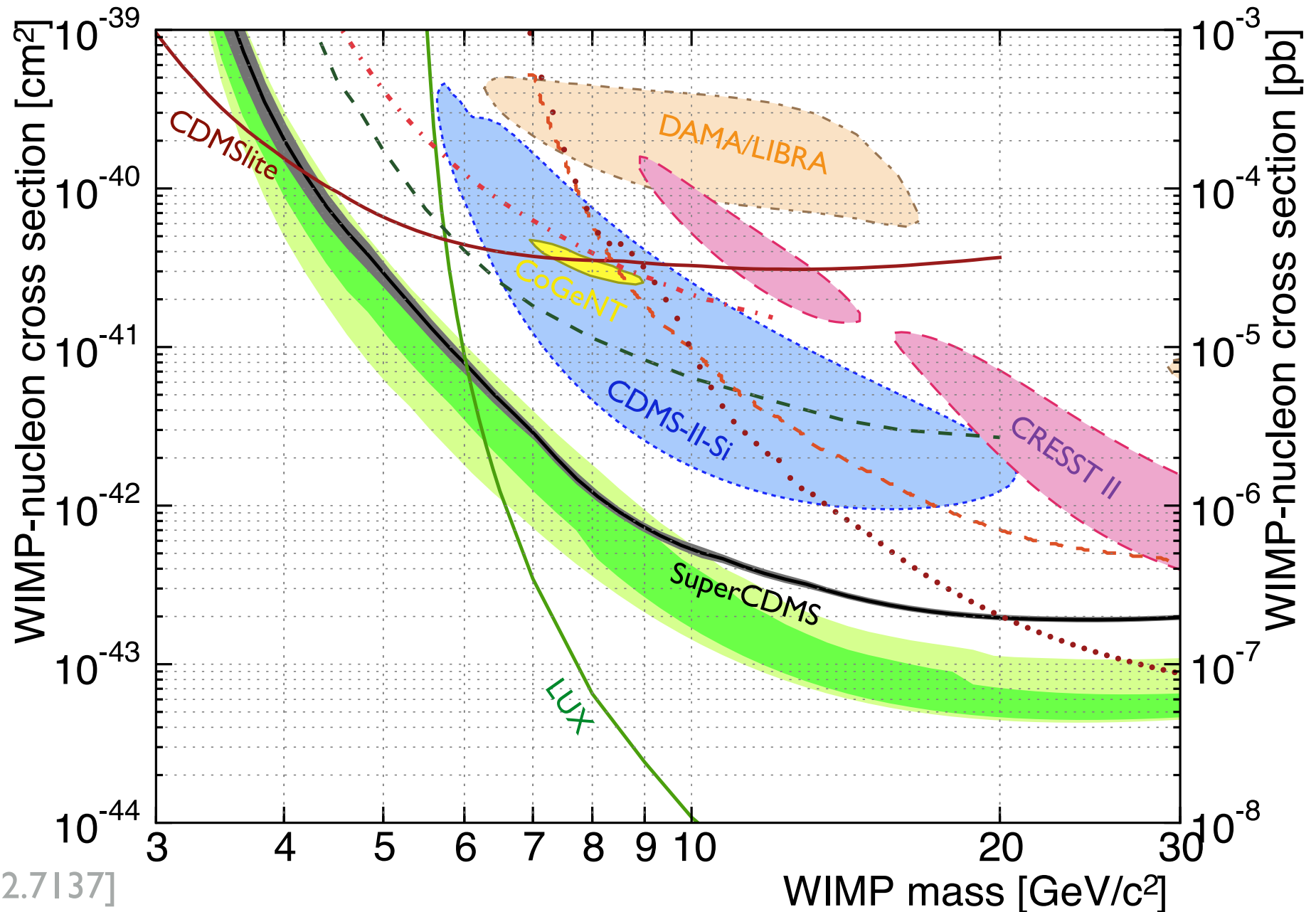
Assumptions

- ◎ **WIMPs: Weak-scale interacting DM particles**
non-relativistic, with mass $1 \text{ GeV} - 10 \text{ TeV}$
- ◎ **Actually, only one kind of particle**
like approximating $SU(3) \times SU(2) \times U(1)$ as “just hydrogen”

Limits and fits

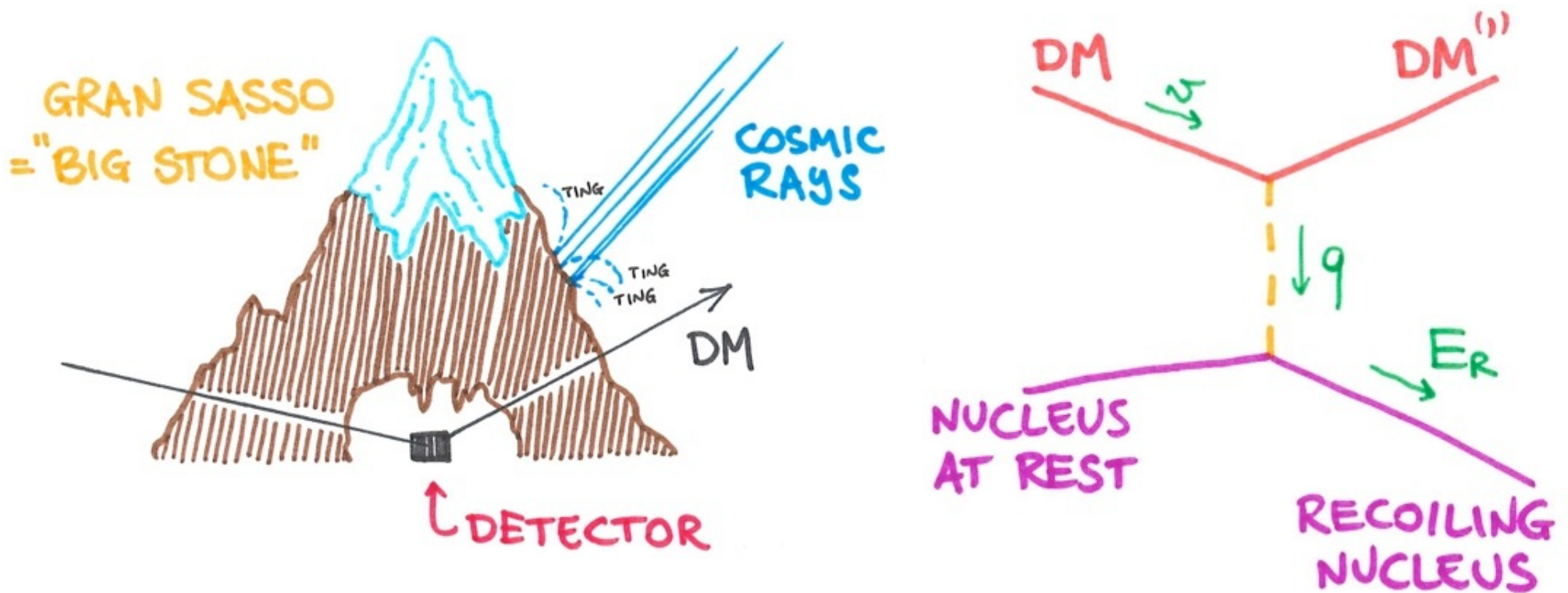


Limits and fits



DM direct detection

Basically explained by



$$\frac{dR_T}{dE_R} = n_T \Phi_{\text{DM}}(t) \frac{d\sigma_T}{dE_R}$$

DM direct detection

The basic ingredient is the recoil rate

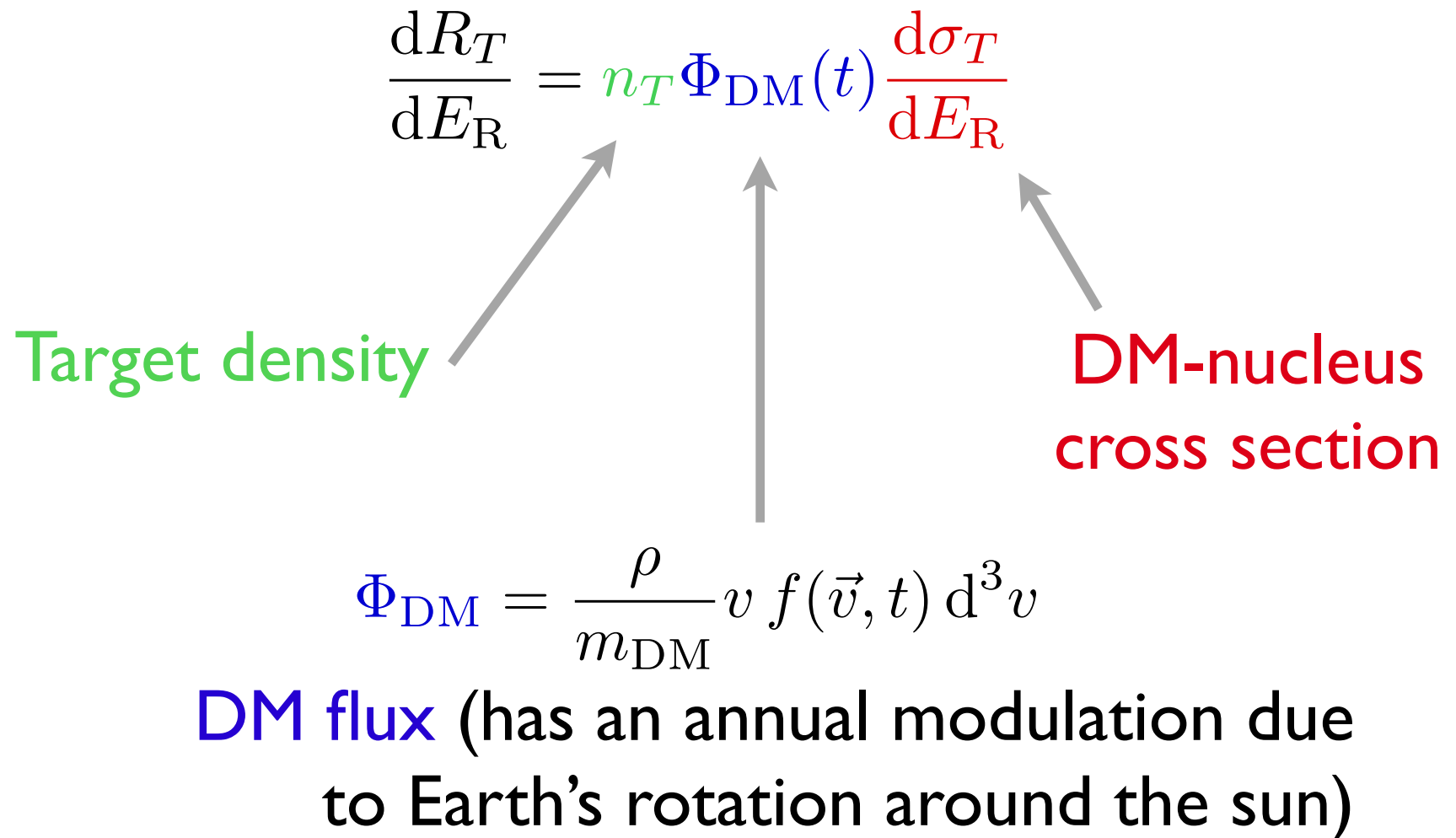
$$\frac{dR_T}{dE_R} = n_T \Phi_{\text{DM}}(t) \frac{d\sigma_T}{dE_R}$$

Target density

DM-nucleus cross section

$$\Phi_{\text{DM}} = \frac{\rho}{m_{\text{DM}}} v f(\vec{v}, t) d^3v$$

DM flux (has an annual modulation due to Earth's rotation around the sun)



Direct detection rate

$$\frac{dR_T}{dE_R} = \frac{\xi_T}{m_T} \frac{\rho}{m_{\text{DM}}} \int_{v_{\min}(E_R)}^{v_{\text{esc}}} d^3v f(\vec{v}, t) v \frac{d\sigma_T}{dE_R}(E_R, \vec{v})$$

ρ DM local density

$f(\vec{v}, t)$ DM velocity distribution

ξ_T target mass fraction

v_{esc} galactic escape velocity

$$R_{[E'_1, E'_2]}(t) = \sum_T \int_{E'_1}^{E'_2} dE' \epsilon_1(E') \int_0^\infty dE_R \epsilon_2(E_R) G_T(E', Q_T E_R) \frac{dR_T}{dE_R}$$

E' detected energy

$G_T(E', Q_T E_R)$ detector resolution

ϵ_1, ϵ_2 efficiencies

Q_T quenching factor

Differential cross section

$$\frac{d\sigma_T}{dE_R} \sim |\langle N | \mathcal{L}_{q,g} | N \rangle|^2 \mathcal{A}_T^2 F_T^2(E_R)$$

$\mathcal{L}_{q,g}$ fundamental Lagrangian

$F_T(E_R)$ nuclear form factor

$|N\rangle$ nucleon state

$\mathcal{A}_T$ coherence factor

$$\mathcal{A}_T \cong \begin{cases} A_T & \text{spin-independent} \\ Z_T & \text{millicharged DM} \\ Z_T f_p + (A_T - Z_T) f_n & \text{isospin-violating} \\ Z_T, \mu_T & \text{magnetic moment DM} \\ J_T & \text{spin-dependent} \\ \dots & \dots \end{cases}$$

Uncertainties

$$R_{[E'_1, E'_2]}(t) = \sum_T \int_{E'_1}^{E'_2} dE' \epsilon_1(E') \int_0^\infty dE_R \epsilon_2(E_R) \underbrace{G_T(E', Q_T E_R)}_{\text{Experiment}} \frac{dR_T}{dE_R}$$

$$\frac{dR_T}{dE_R} = \frac{\xi_T}{m_T} \frac{\rho}{m_{\text{DM}}} \int_{v_{\min}(E_R)}^{v_{\text{esc}}} d^3v f(\vec{v}, t) v \frac{d\sigma_T}{dE_R}(E_R, \vec{v}) \quad \text{Astrophysics}$$

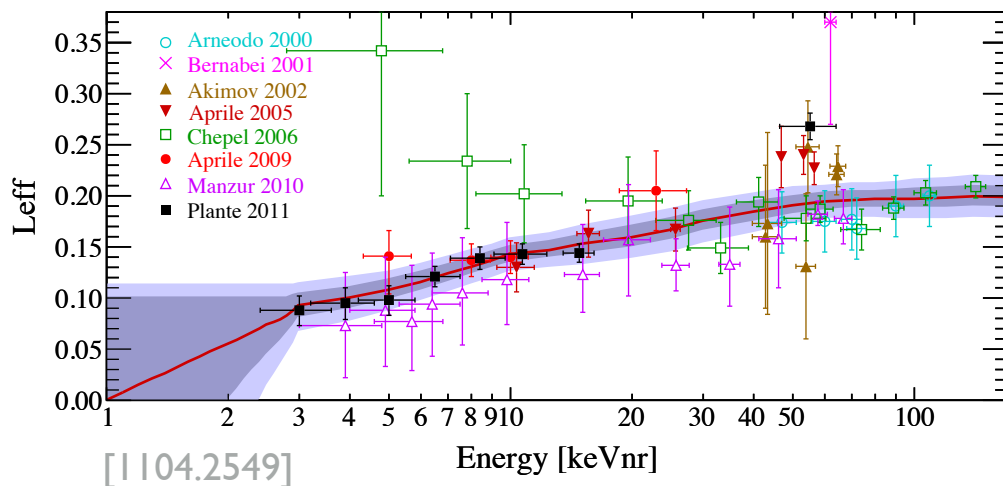
$$\frac{d\sigma_T}{dE_R} \sim |\langle N | \mathcal{L}_{q,g} | N \rangle|^2 \mathcal{A}_T^2 F_T^2(E_R) \quad \text{Particle physics}$$

Uncertainties

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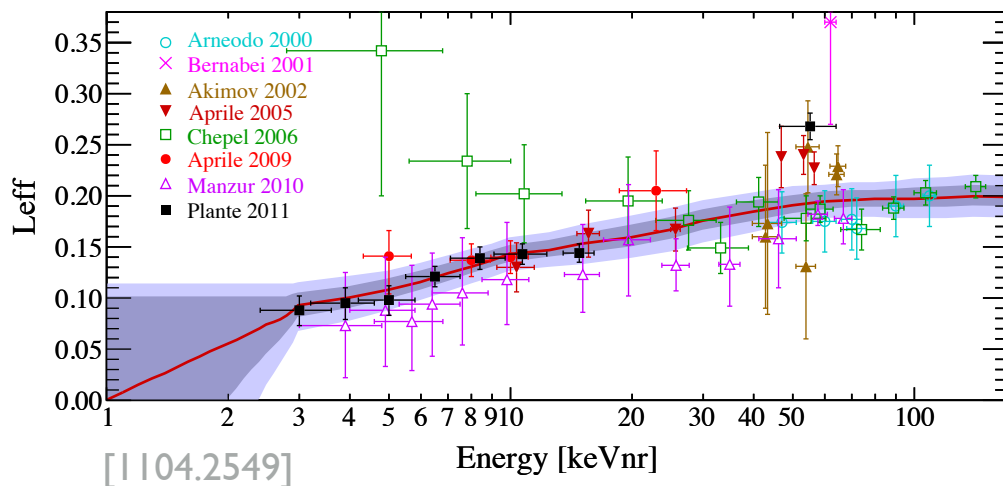


Uncertainties

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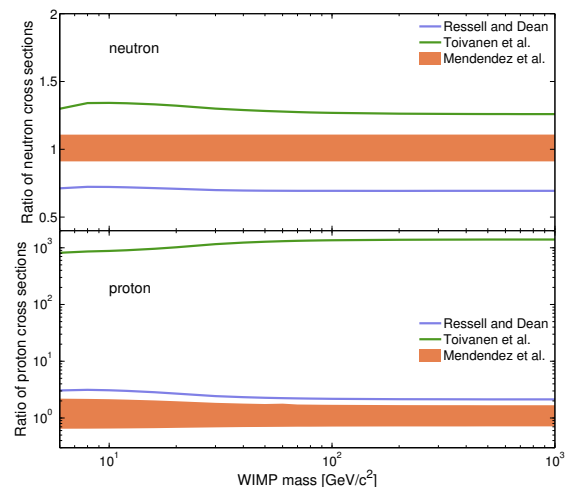
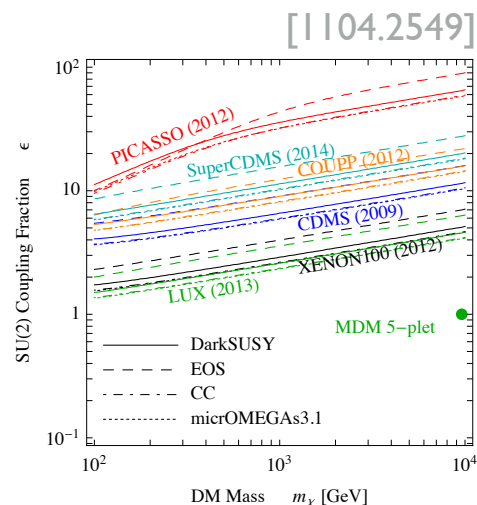
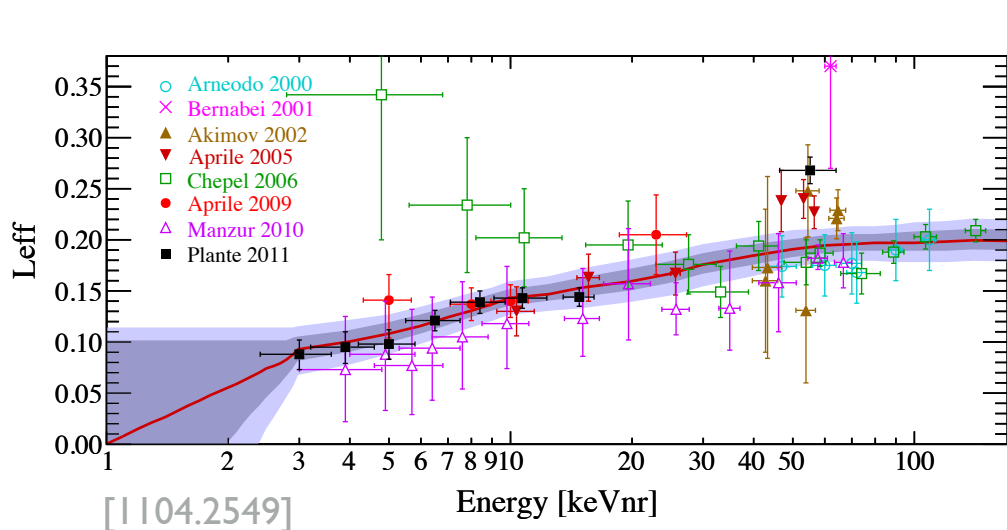
Experiment

$$\frac{dR_T}{dE_R} = \frac{\xi_T}{m_T} \frac{\rho}{m_{\text{DM}}} \int_{v_{\text{min}}(E_R)}^{v_{\text{esc}}} d^3v f(\vec{v}, t) v \frac{d\sigma_T}{dE_R}(E_R, \vec{v})$$

Astrophysics

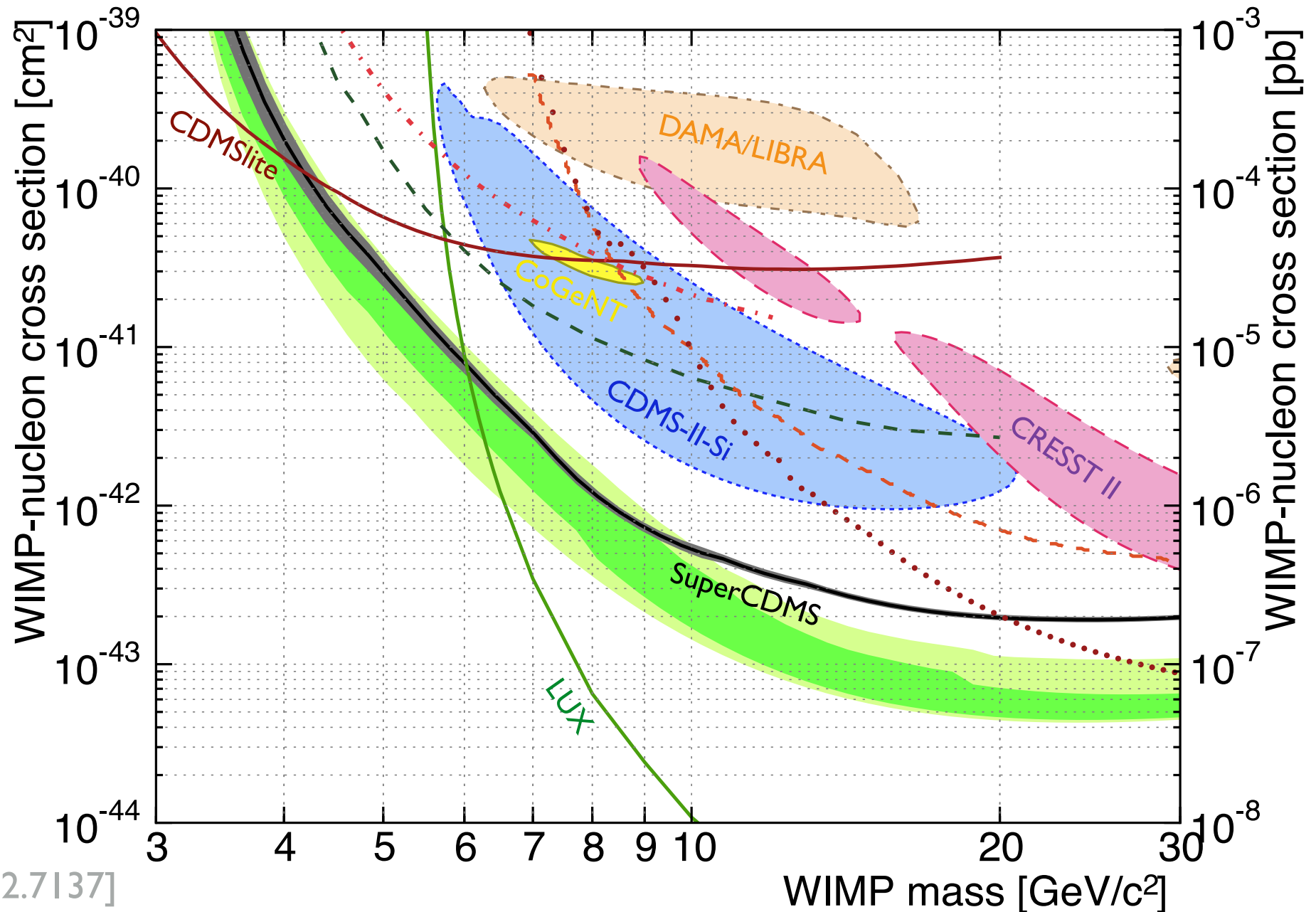
$$\frac{d\sigma_T}{dE_R} \sim |\langle N | \mathcal{L}_{q,g} | N \rangle|^2 \mathcal{A}_T^2 F_T^2(E_R)$$

Particle physics



[1301.6620]

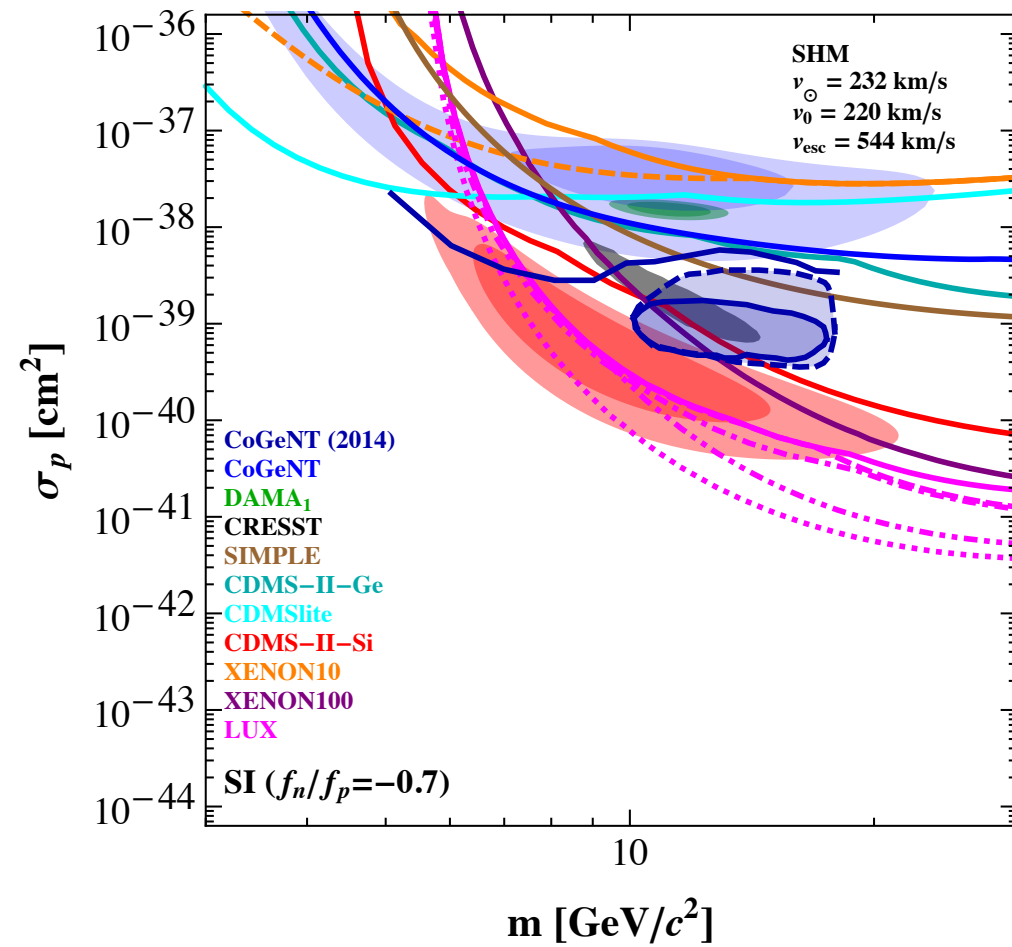
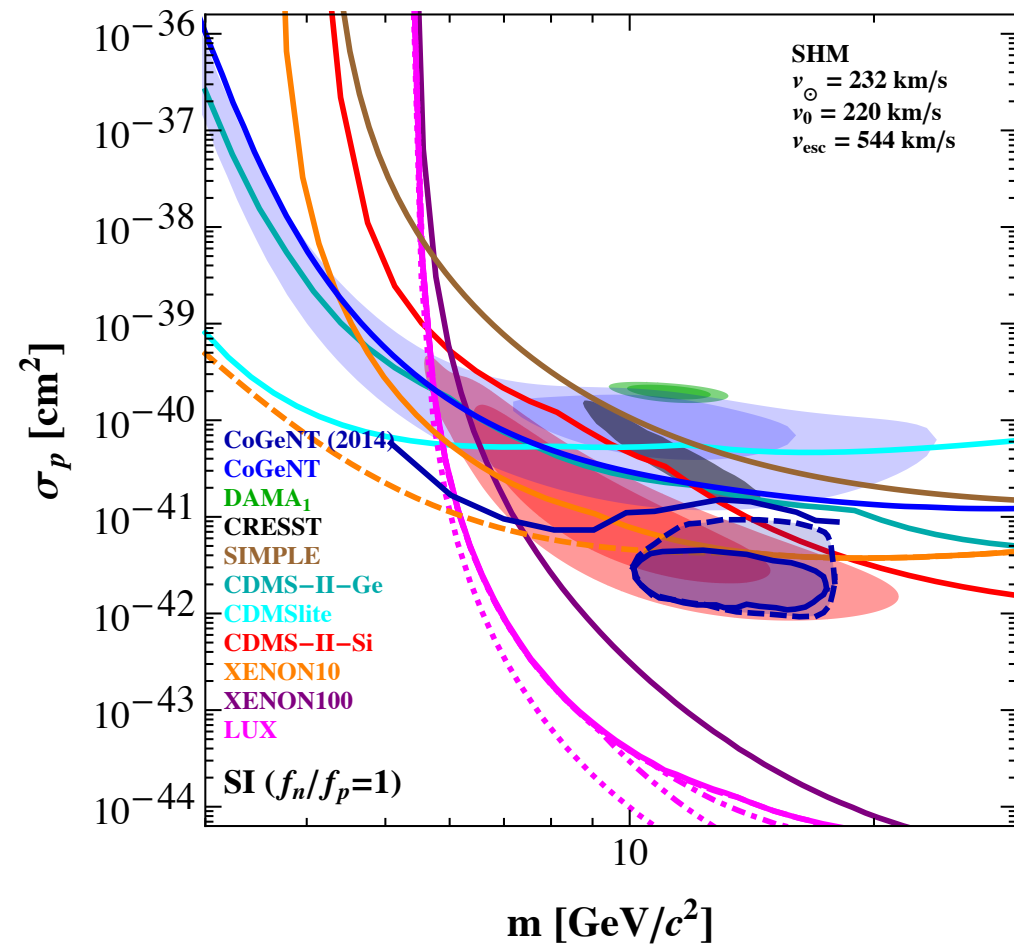
Back to this plot



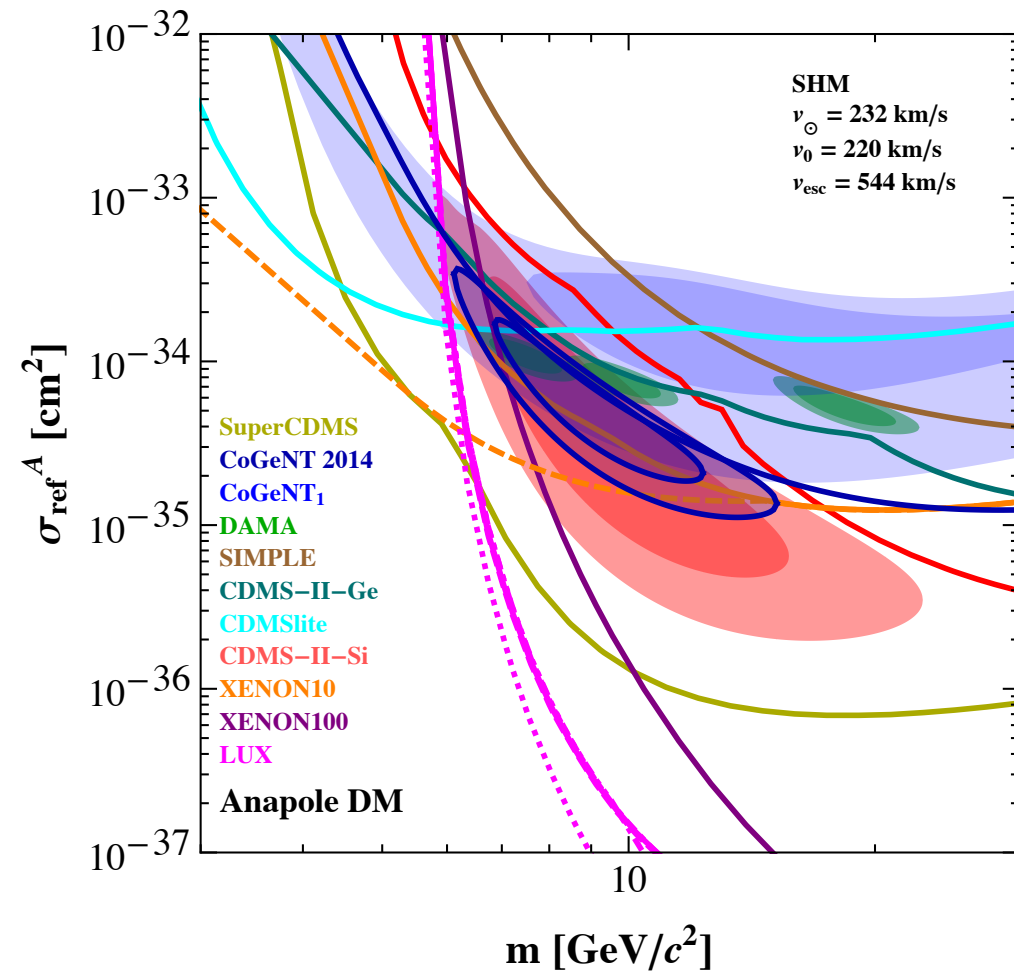
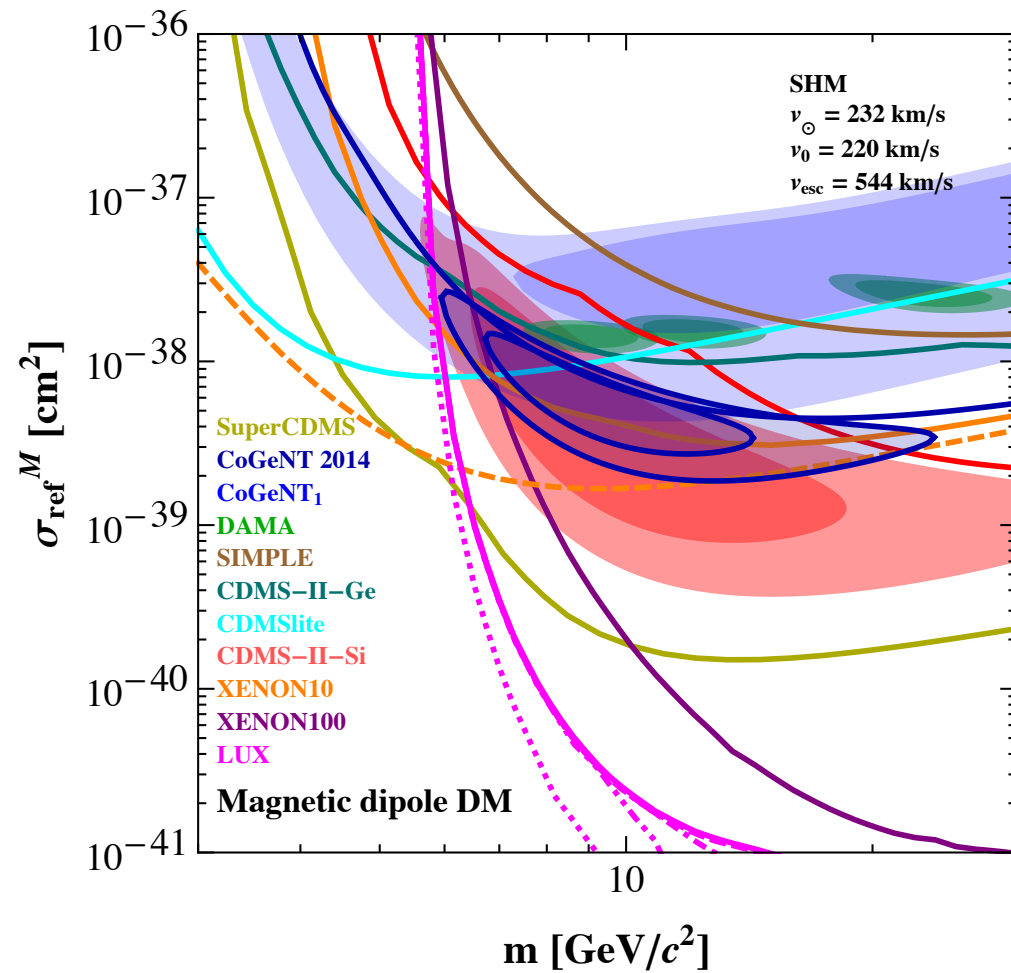
Assumptions behind this plot

- ◎ Spin-independent interaction (with isospin-symmetric coupling)
- ◎ Truncated Maxwell-Boltzmann velocity distribution (Standard Halo Model)

Turning the interaction knob



Turning the interaction knob

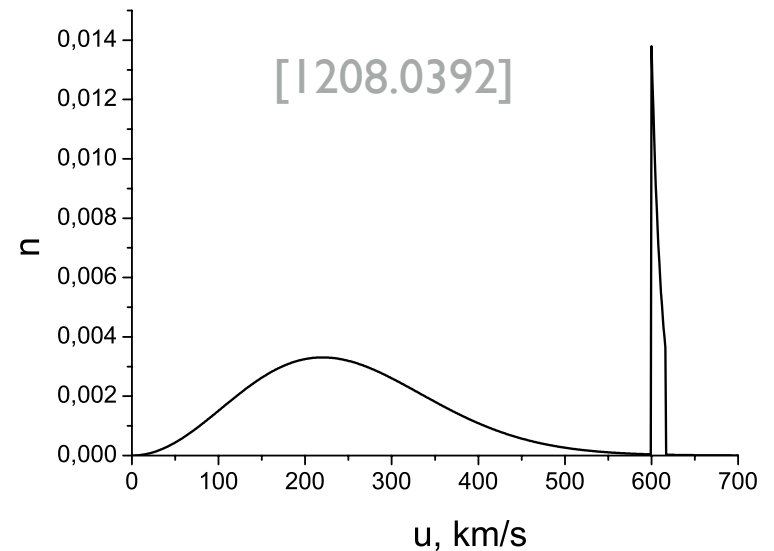
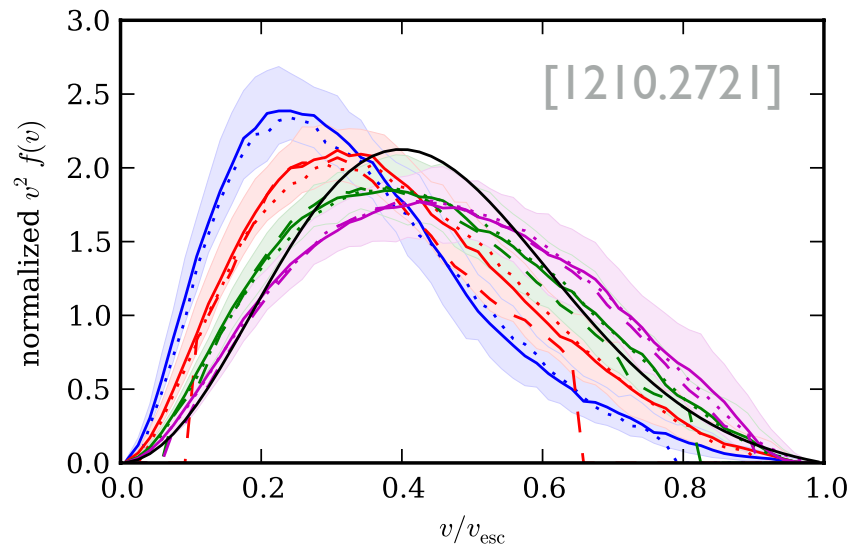


Astrophysical uncertainties

$$R_{[E'_1, E'_2]}(t) = \sum_T \int_{E'_1}^{E'_2} dE' \epsilon_1(E') \int_0^\infty dE_R \epsilon_2(E_R) \underbrace{G_T(E', Q_T E_R)}_{\text{Experiment}} \frac{dR_T}{dE_R}$$

$$\frac{dR_T}{dE_R} = \frac{\xi_T}{m_T} \frac{\rho}{m_{\text{DM}}} \int_{v_{\min}(E_R)}^{v_{\text{esc}}} d^3v \underbrace{f(\vec{v}, t)}_{\text{Astrophysics}} v \frac{d\sigma_T}{dE_R}(E_R, \vec{v})$$

$$\frac{d\sigma_T}{dE_R} \sim |\langle N | \mathcal{L}_{q,g} | N \rangle|^2 \mathcal{A}_T^2 F_T^2(E_R) \quad \text{Particle physics}$$



Three approaches

- Marginalize over astrophysical uncertainties
see e.g. Arina [1310.5718] and references therein
- Try to find alternative halo models, either driven by physical arguments or by fitting simulations or observations
see e.g. references in Freese, Lisanti, Savage [1209.3339]
- Try to factor astrophysics out of your problem as much as you can
Fox, Liu, Weiner [1011.1915], Frandsen, Kahlhoefer, McCabe, Sarkar, Schmidt-Hoberg [1111.0292][1304.6066], Gondolo, Gelmini [1202.6359] + Del Nobile, Huh [1304.6183][1306.5273][1311.4247][1401.4508], Herrero-Garcia, Schwetz, Zupan [1112.1627][1205.0134] + Bozorgnia [1305.3575], Feldstein, Kahlhoefer [1403.4606], Fox, Kahn, McCullough [1403.6830]

Tweaking the rate

$$R_{[E'_1, E'_2]}(t) = \int_0^\infty dv_{\min} \tilde{\eta}(v_{\min}, t) \mathcal{R}_{[E'_1, E'_2]}(v_{\min})$$

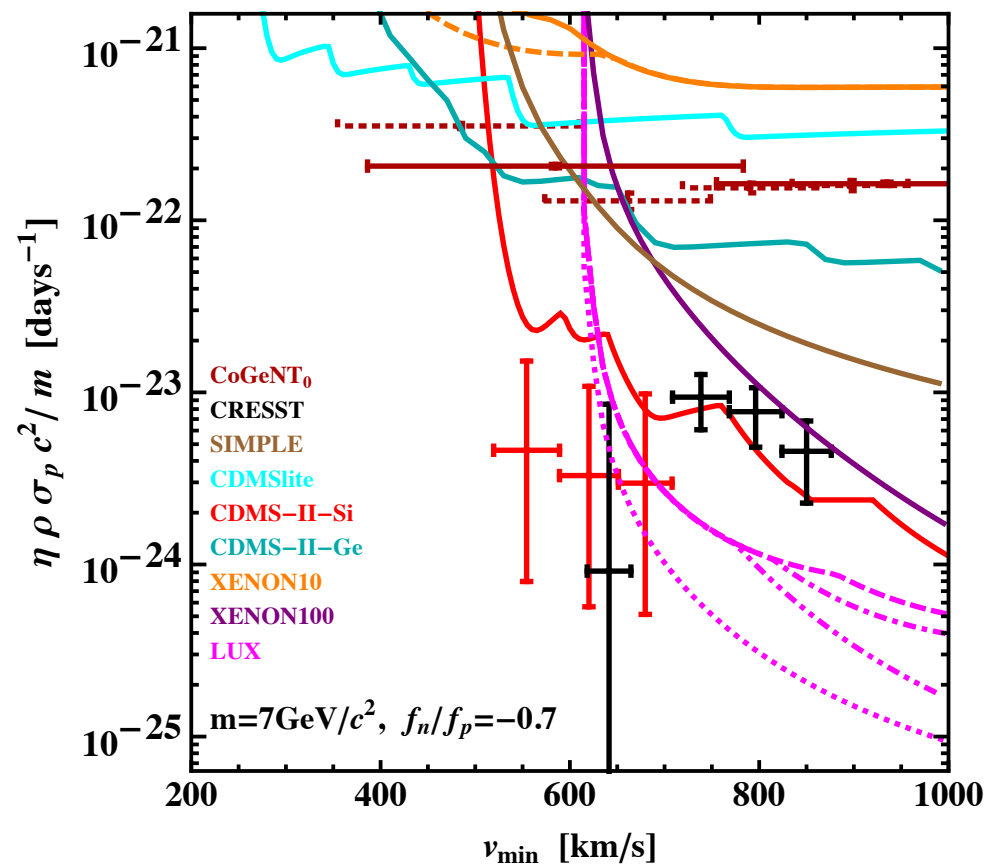
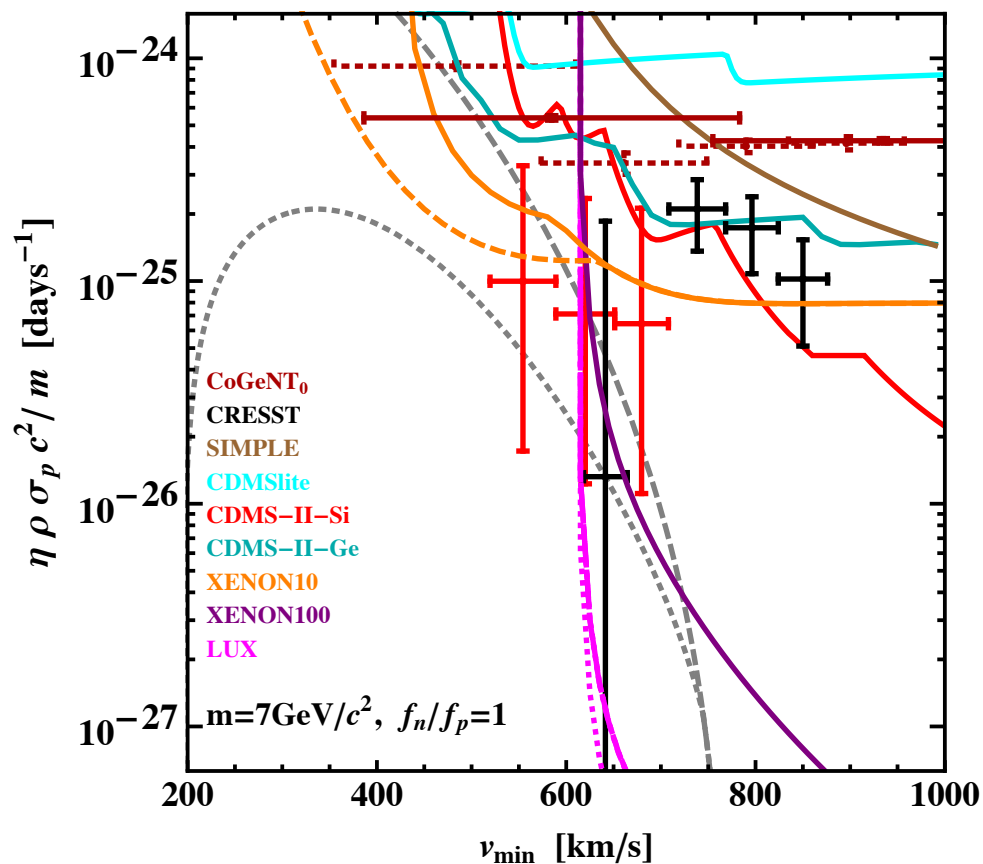
$$\tilde{\eta}(v_{\min}, t) \equiv \frac{\rho \sigma_{\text{ref}}}{m_{\text{DM}}} \int_{v_{\min}}^\infty d^3v \frac{f(\vec{v}, t)}{v}$$

detector-independent,
unknown function of $v_{\min}$
to be determined by data

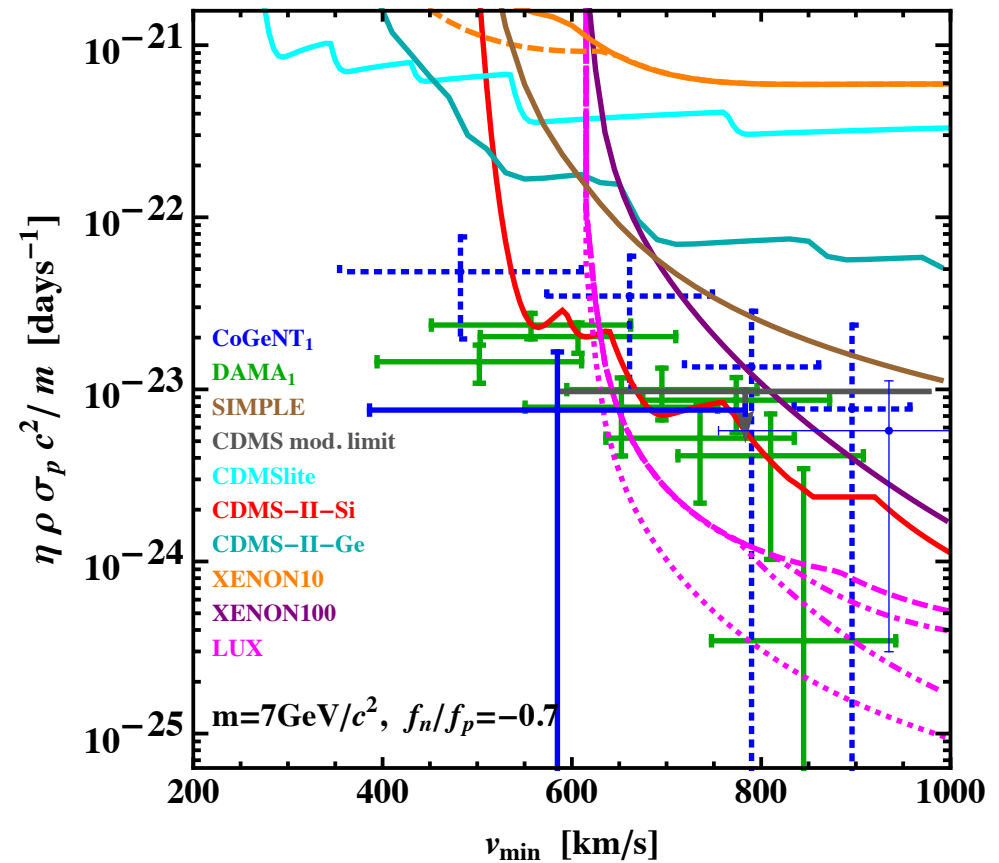
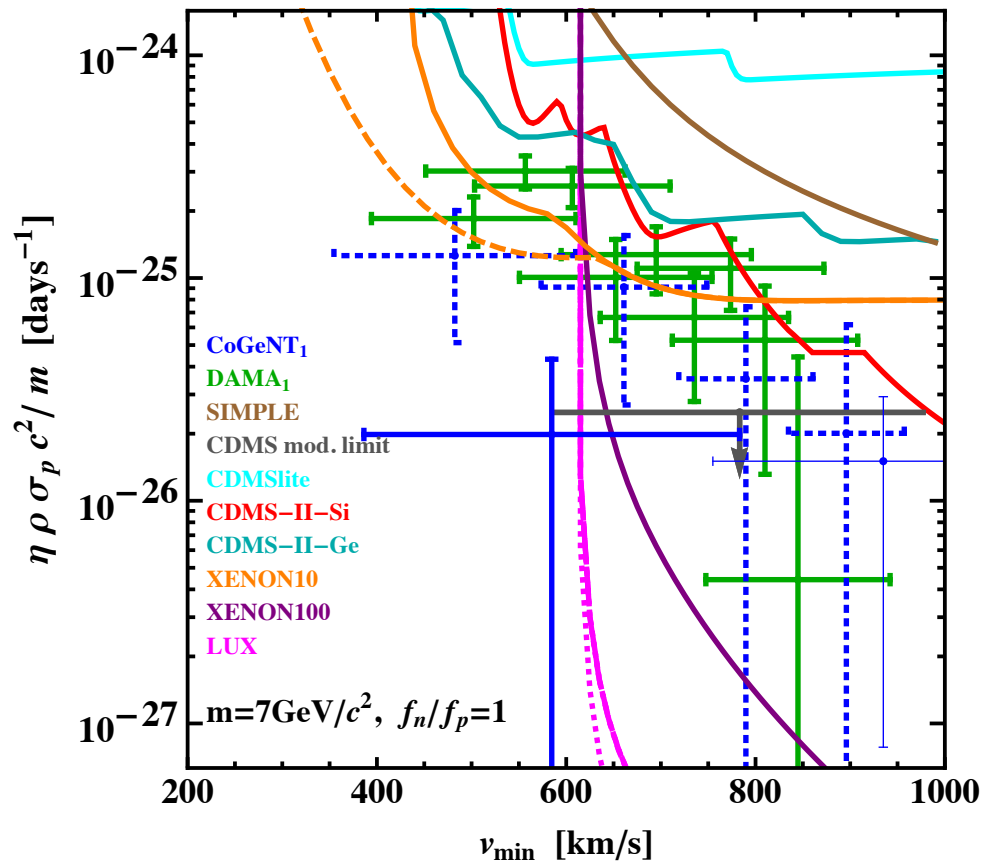
known function
determined by:

- detector
- interaction
- DM mass

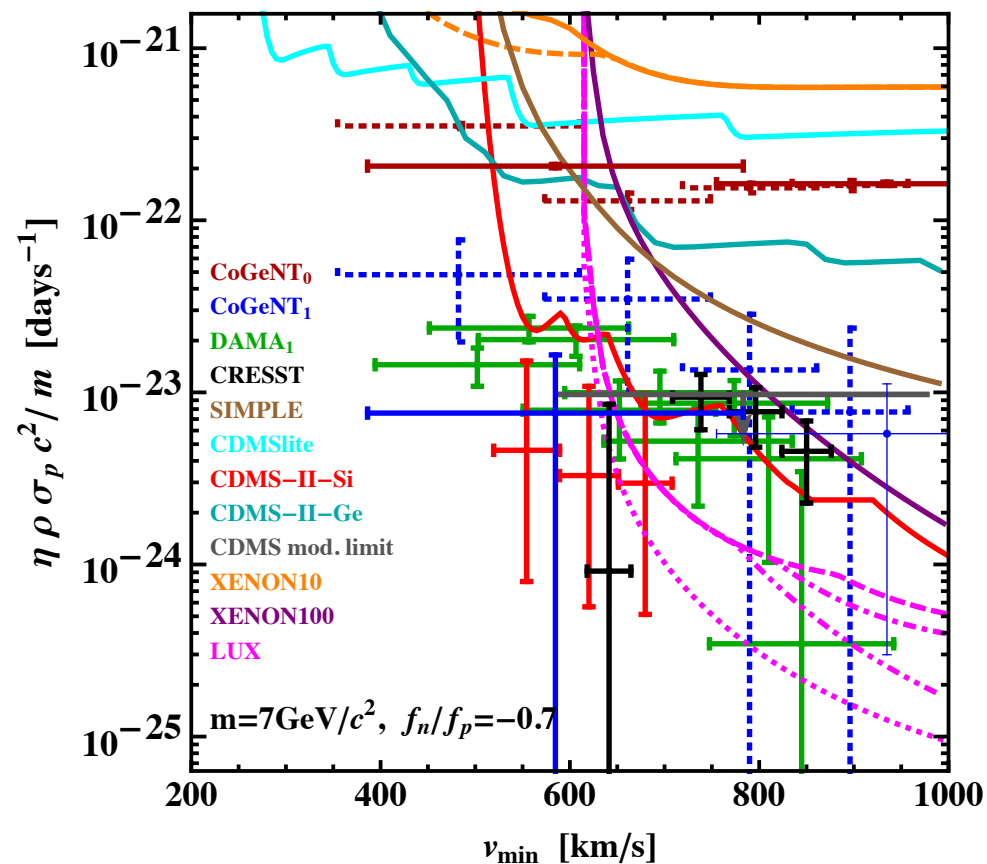
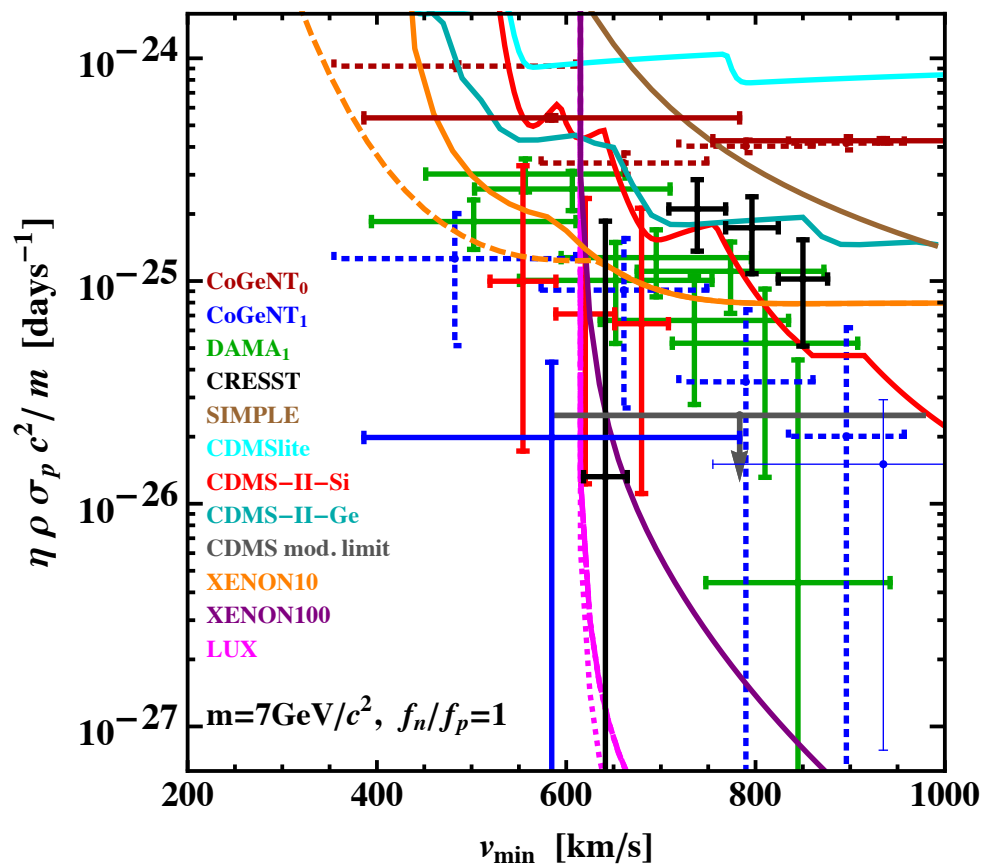
Halo-independent plot



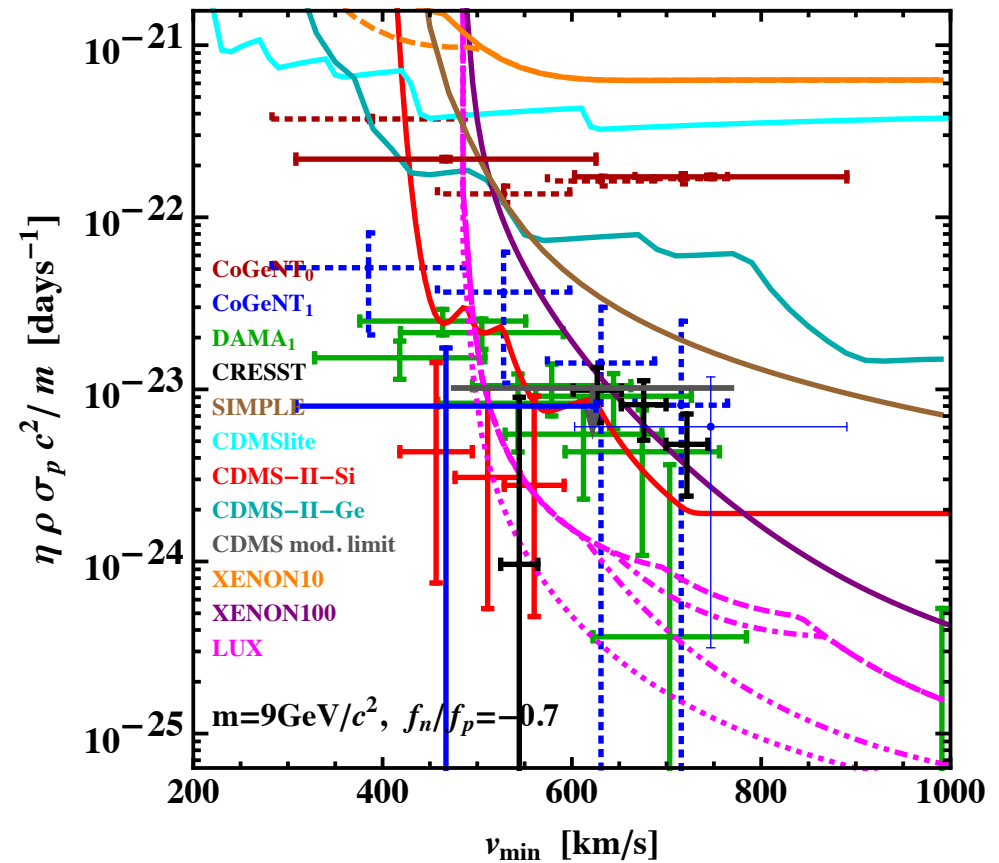
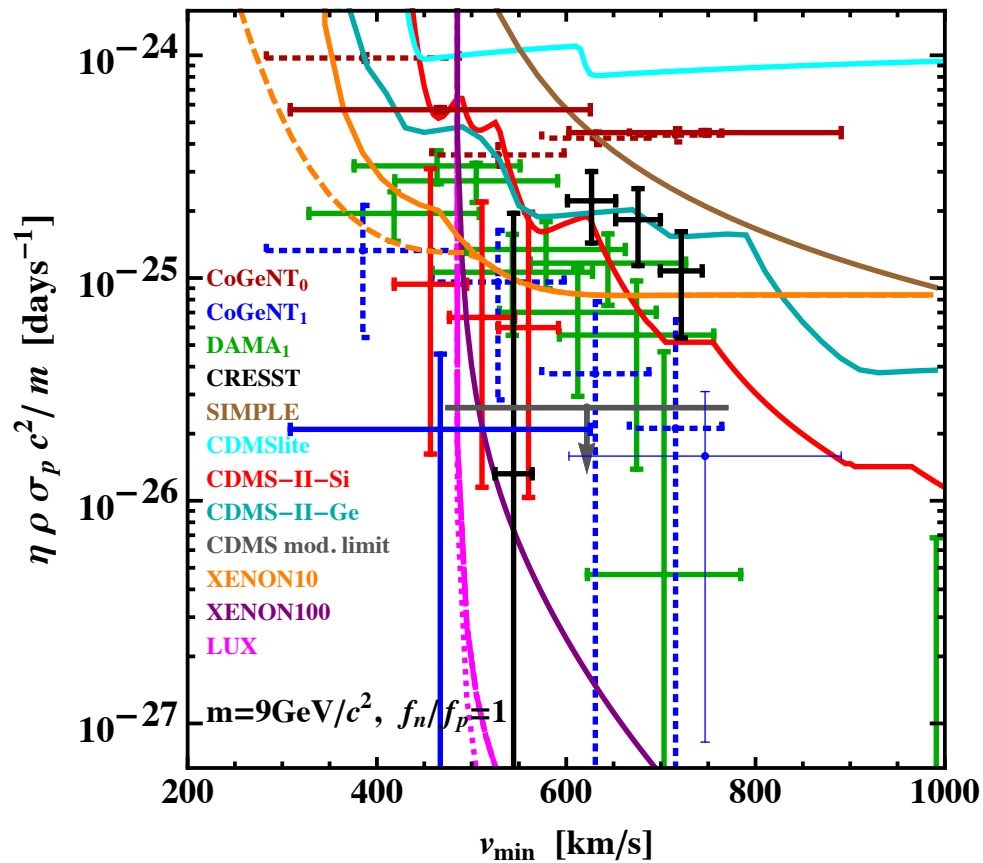
Halo-independent plot



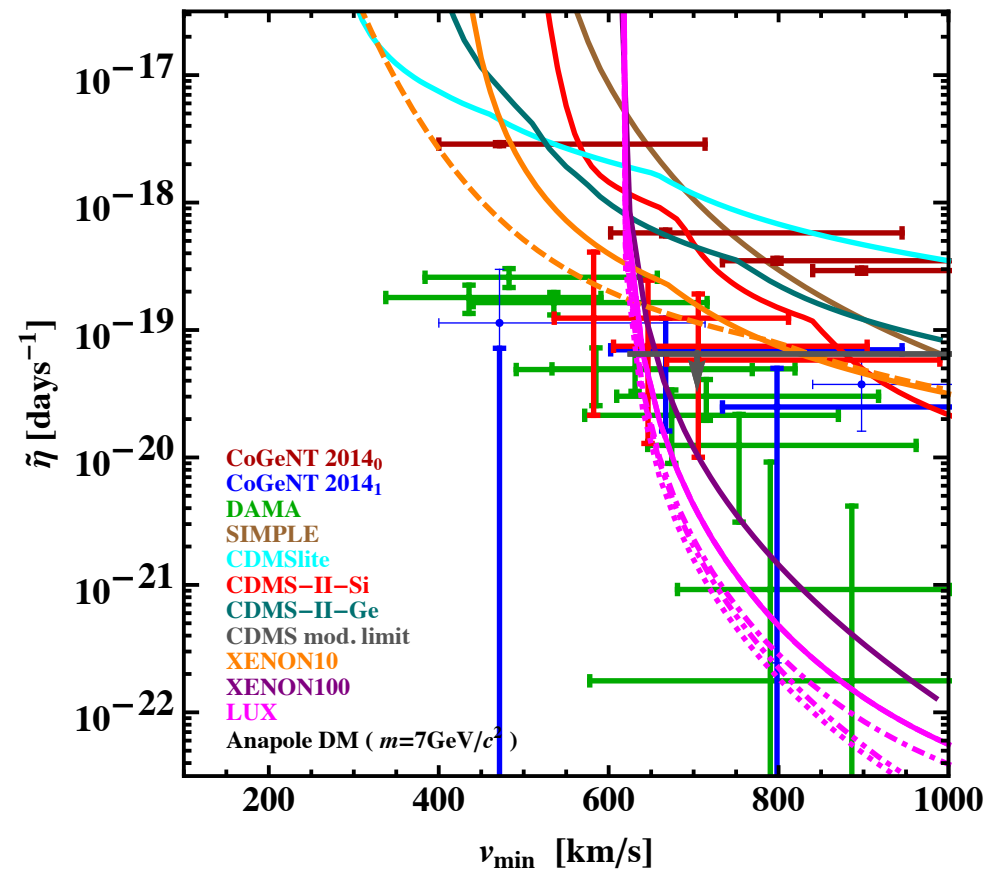
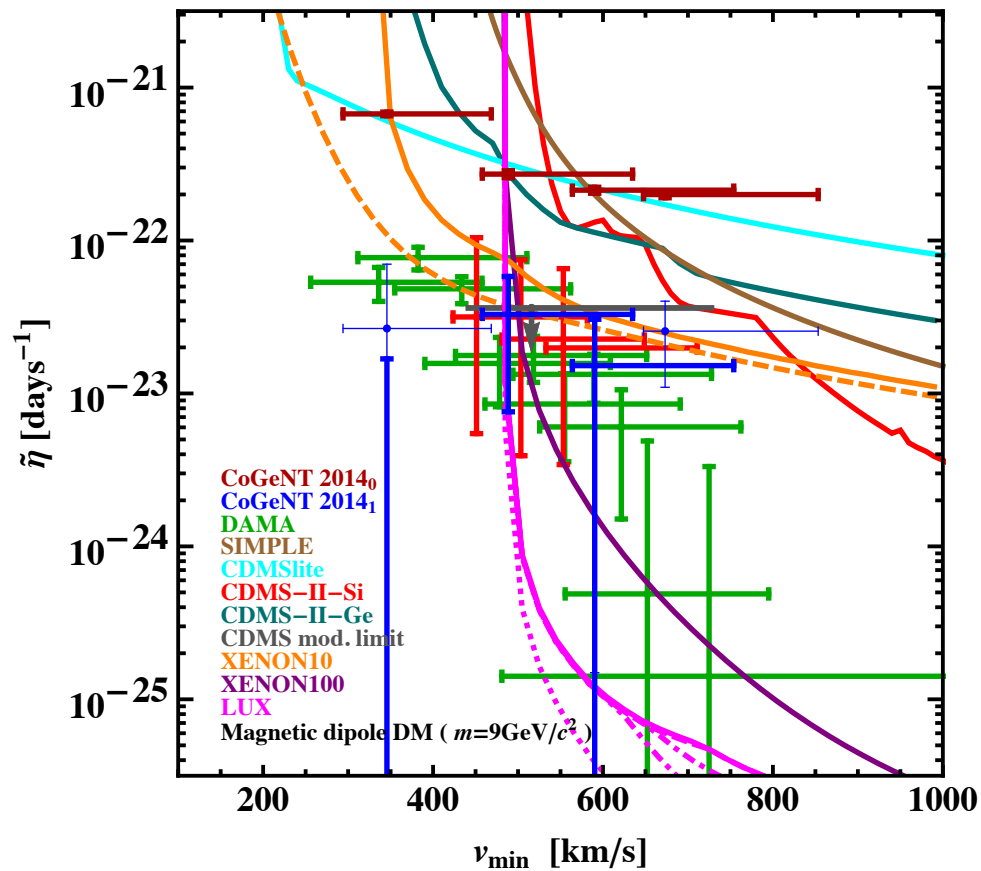
Halo-independent plot



Halo-independent plot



Halo-independent plot



Conclusions

With an eye on trying to explain current direct detection experiments,

- Much work has been done in trying new roads to interpret the data
- Uncertainties are being investigated (and possibly reduced)
- Model-independent methods are being devised
- Possible explanations are being proposed
- Experiments keep being developed

Outlook

Difficult to think what we should be doing differently or better:

- Large community of “dark matter enthusiasts”
- Big ongoing effort to detect the ~~bastard~~ DM
- Many experiments
- Different detection strategies: direct, indirect, collider

A lesson from the LHC

Theorists have long been biased by the mirage of new physics at the Electroweak scale, e.g. Supersymmetry (or WIMPs)

But the Higgs might well be the only new fundamental particle the LHC will ever see

The secret message is:

- be open-minded
- do not exclude possibilities a priori
- look for the darn thing everywhere