

Gamma ray excess in the Galactic Center (for the panel discussion)

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TOPICS

- 1 Remarks on WIMP model
- 2 Prompt Emission only analysis ??
- 3 On future detector

Remarks on WIMP model

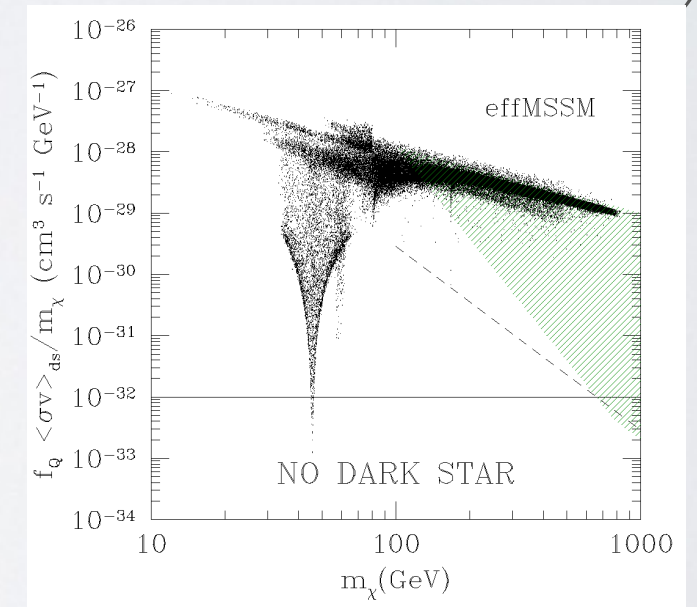
- Characterized by two parameters $(m_\chi, \langle \sigma_{\text{ann}} v \rangle)$ for a given “2-body” final state ($b\bar{b}$, $\tau\bar{\tau}$, W^+W^- , etc)
- Good ; model independence
- Bad ; presence of other possibilities
- Model dependent Internal Bremsstrahlung(IB) (irrelevant for $\sim 10\text{GeV}$ WIMP)
- Models with light neutral particle
(NMSSM+ RN, hidden U(1) gauge boson, ...)

$$\tilde{N}\tilde{N} \rightarrow AA \rightarrow b\bar{b}b\bar{b}$$

- Freeze-out cross section

$$\langle \sigma v \rangle_{\text{present}} \neq \langle \sigma v \rangle_{\text{decouple}} = 3 \times 10^{-26} \text{cm}^3/\text{s}$$

- Mostly $b\bar{b}$ final state \longrightarrow Quite generic situation in Majorana WIMP model
Chiral suppression (P-wave suppression) ; $\langle \sigma v \rangle_{\text{present}} \sim \left(\frac{m_f}{m_\chi} \right)^2$



Prompt Emission Only Analysis ??

- Gamma-ray signal from WIMP

1) Prompt emission (FSR+VIB)

2) Inverse Compton (IC)

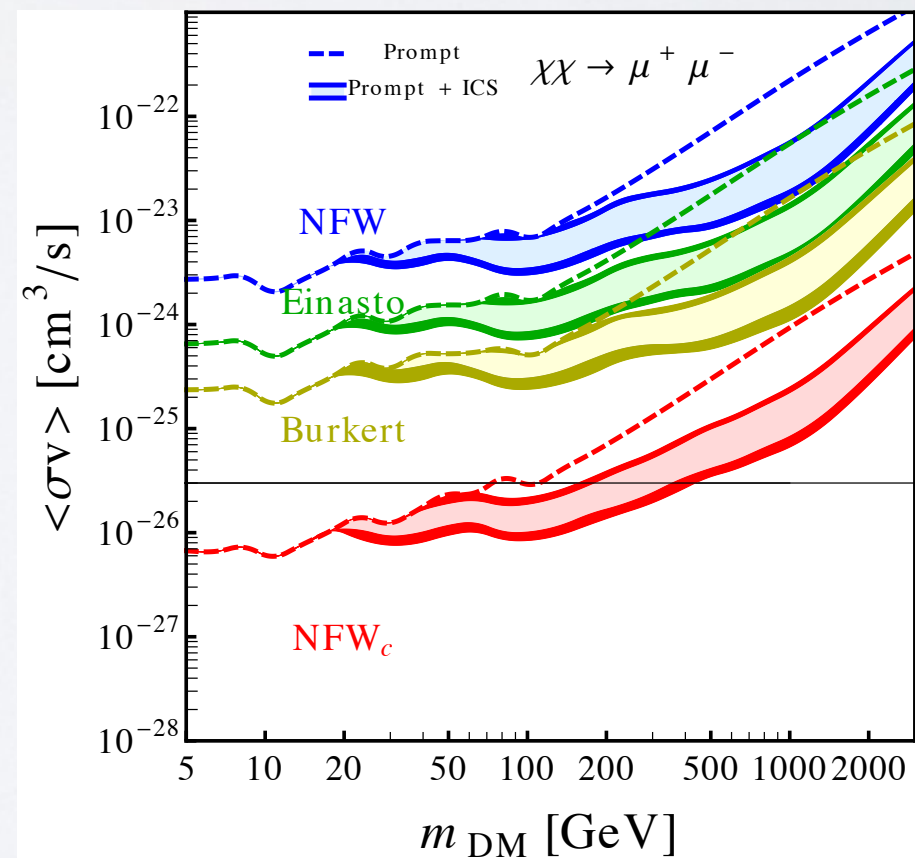
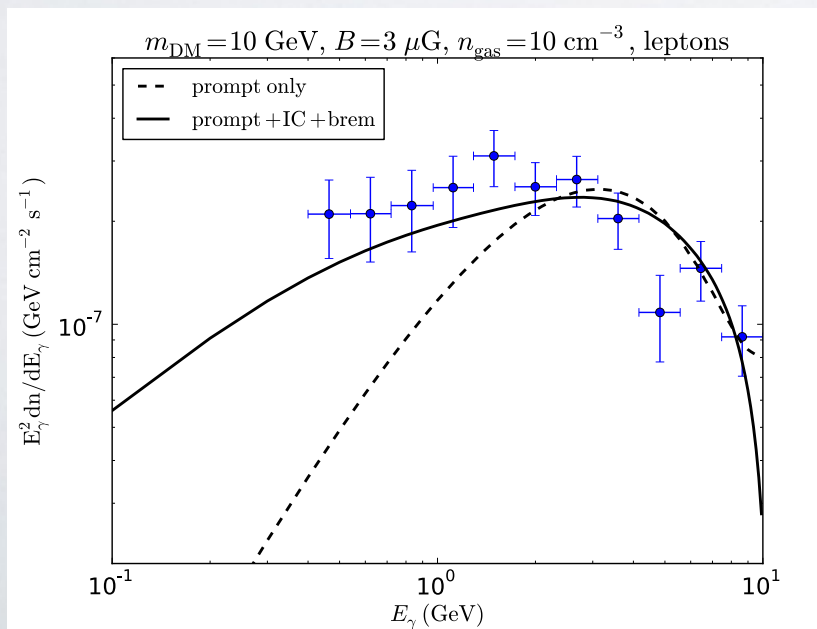
$\chi\chi \rightarrow e^-e^+ \rightarrow \text{propagation} \rightarrow \text{IC}(\text{CMB} + \text{IR} + \text{starlight})$

- Lacroix, Boehm and Silk (1403.1987)

"Fitting the Fermi-LAT GeV excess: on the importance of including the propagation of electrons from dark matter"

- IC component ; sensitive to $|\vec{B}|$

$$N_{\gamma}^{\text{IC}} \sim \frac{b_{\text{IC}}}{b_{\text{IC}} + b_{\text{syn}}} n_e^2$$



On the future detector

- Three ways to improve
 - 1) Large effective Area
 - 2) Good energy resolution
 - 3) Good angular resolution
- Of course, “Large effective Area” cannot be sacrificed.
(Success of Fermi Satellite)
- Energy resolution VS Angular resolution

Energy resolution

- DM mass \sim few TeV
(prominent spectral feature from IB)
- line spectrum ($\gamma\gamma, \gamma Z, \dots$)

Angular resolution

- DM mass \sim few 10 GeV
- Continuum spectrum