



The QUIJOTE CMB Experiment

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on behalf of the QUIJOTE Collaboration



The QUIOTE Experiment

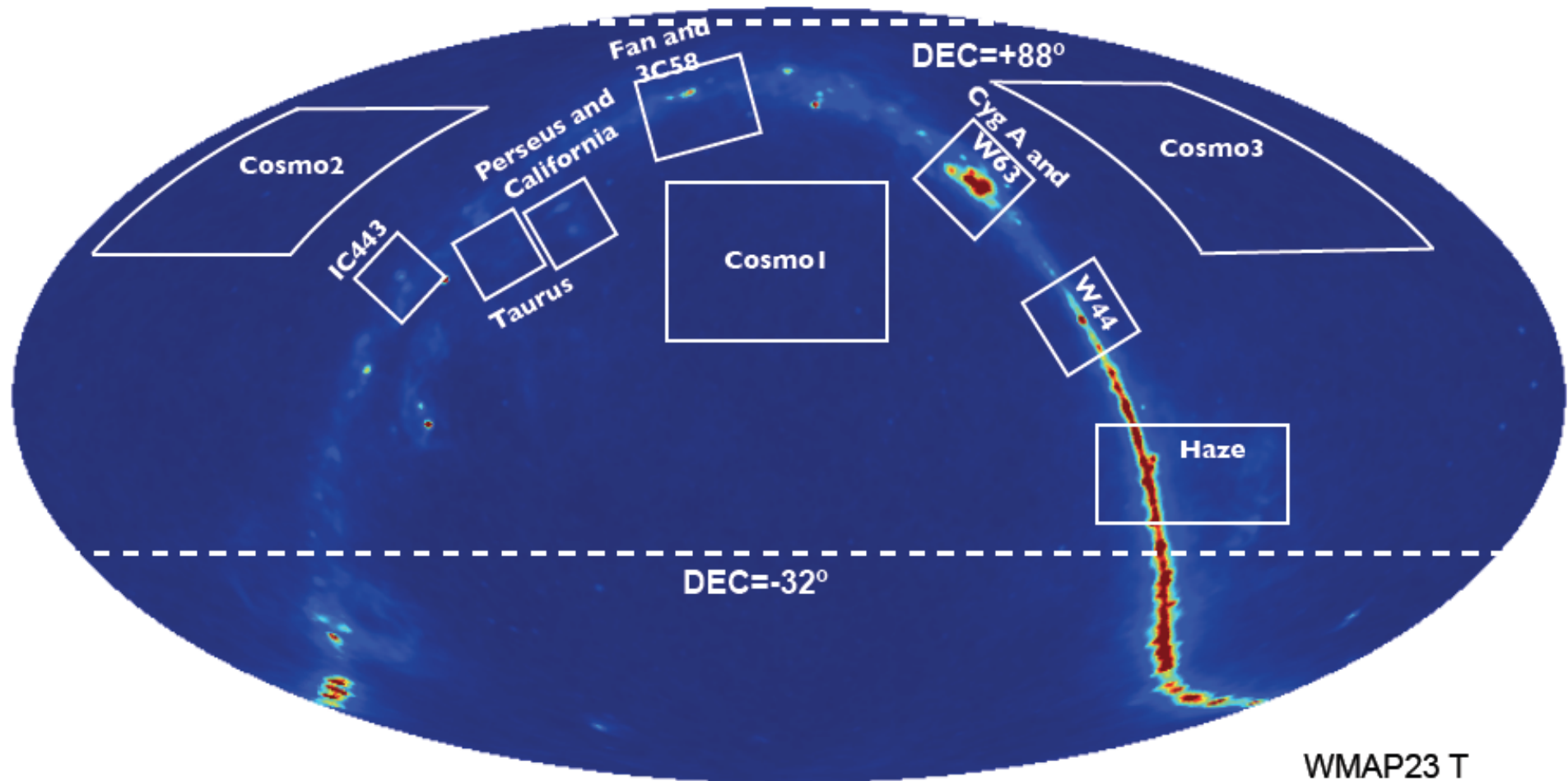
- **Site:** Teide Observatory (altitude 2400 m, 28.3° N, 16.5 W)
- **Frequencies:** 11, 13, 17, 19, 30 and 40 GHz.
- **Angular resolution:** 0.92° to 0.26°
- **Sky coverage:** $-32^\circ < \text{Dec.} < 88^\circ$ (fsky=0.65).
- **2 telescopes and 3 instruments:**
 - Two telescopes installed (2012 and 2014)
 - **Multi-Frequency Instrument (MFI)** with 4 polarimeters at 10-20 GHz. In operation since Nov 2012
 - **Second Instrument (TGI)** with 31 polarimeters @ 30 GHz. First light in May 2016
 - **Third instrument (FGI)** at 42 GHz (31 polarimeters).
 - TGI and FGI in joint commissioning phase
- **Observing strategy:** Deep observations in selected areas plus wide survey
- **Point source follow-up observations** with VLA to correct for polarised sources selected from PLANCK maps. Observations in different epochs are being performed to study variability



Scientific goals

- To provide polarization maps at 6 frequencies in the range 10 - 40 GHz with sufficient sensitivity to correct radio foreground emission (synchrotron and AME) and to constrain the imprint of B-modes down to $r=0.05$
- To characterise radio foregrounds in a unique frequency range [10-20 GHz] which is not covered by other experiments so far
- Observational strategy
 - Wide survey
Covering 20,000 deg^2
 $\approx 15 \mu\text{K}/(\text{beam } 1^\circ)$ with the MFI @ 11, 13, 17 and 19 GHz
 - Deep cosmological survey
It will cover around 3,000 deg^2 in three separated fields. The scientific goal is to reach $r=0.05$ after 3 years of operations of the TGI+FGI
 $10 \mu\text{K}/(\text{beam } 1^\circ)$ after 1 year with the MFI @ 11, 13, 17 and 19 GHz
 $\approx 1 \mu\text{K}/(\text{beam } 1^\circ)$ after 1 year with the TGI and FGI @ 30 and 40 GHz
 - Other Galactic regions
Covering few hundred deg^2 . To understand radio foregrounds
 $\approx 30\text{-}40 \mu\text{K}/(\text{beam } 1^\circ)$ with the MFI @ 11, 13, 17 and 19 GHz

QUIJOTE cosmological and Galactic fields

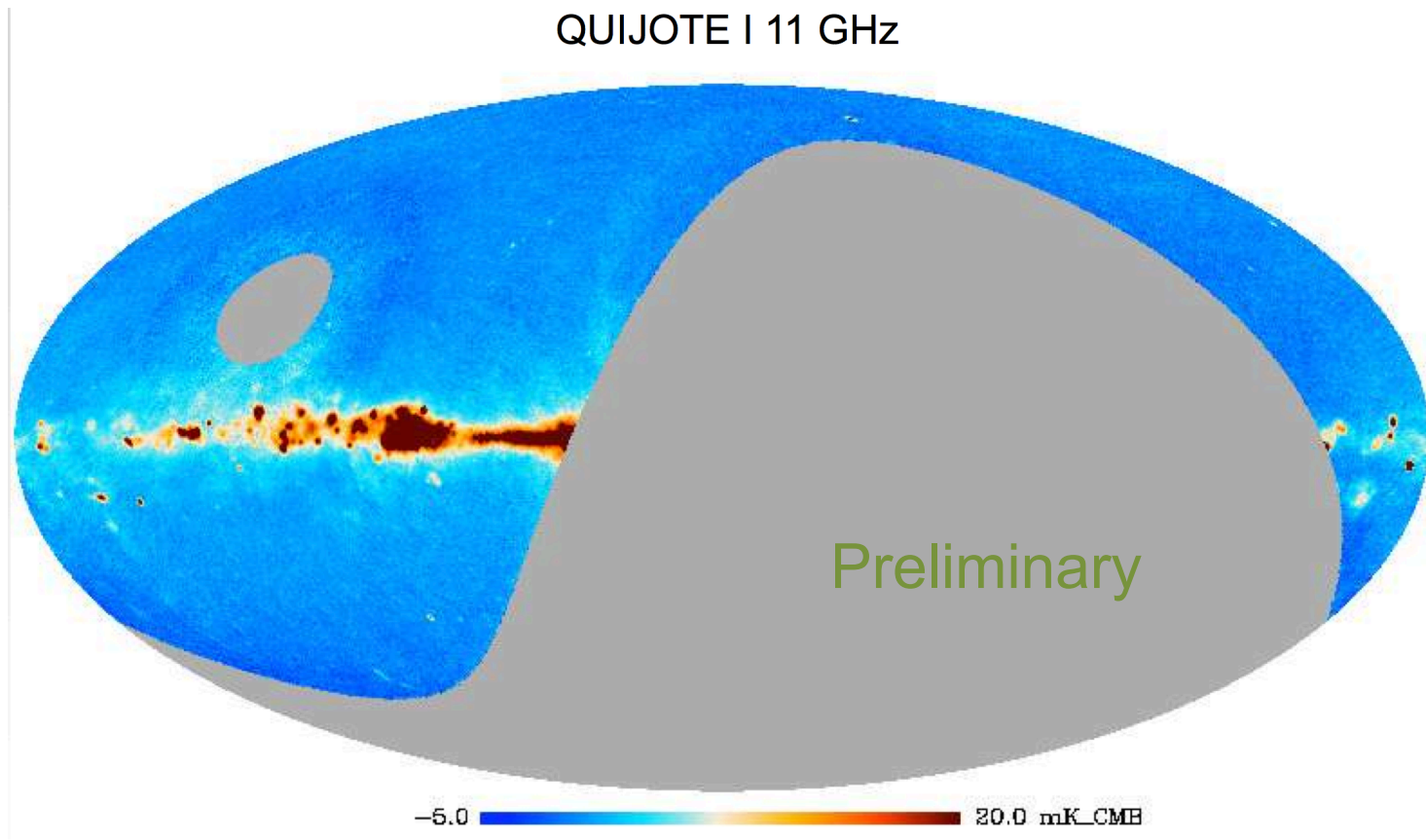


Observation time: 21.000 hours (2.4 years)

Wide survey



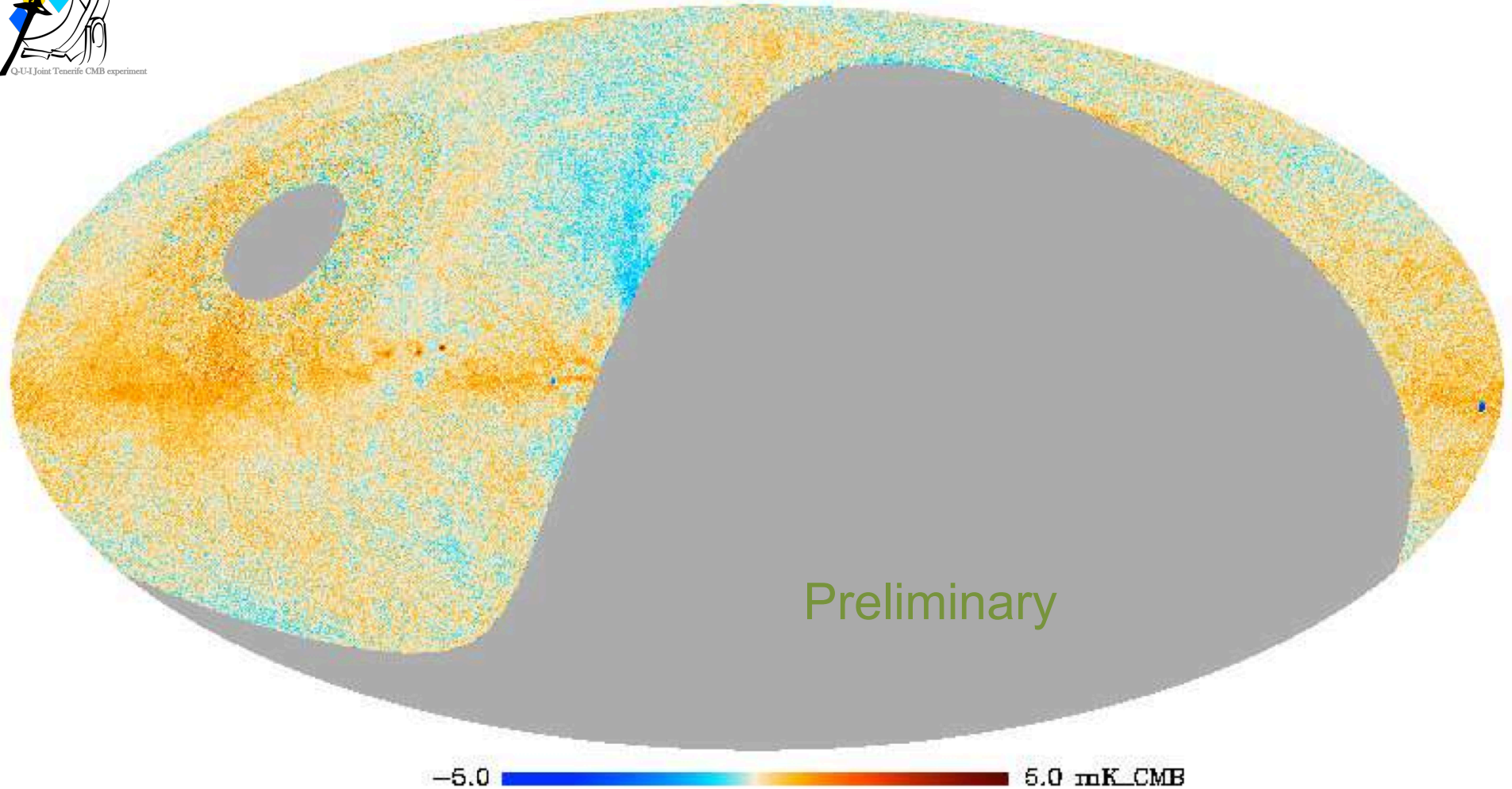
- 8,500 hrs on a region of 20,000 deg² in the northern sky.
- Still on-going (will reach ~10000 hrs).
- Goal: ~15 μ K/beam in polarization



Wide survey: Q 11 GHz

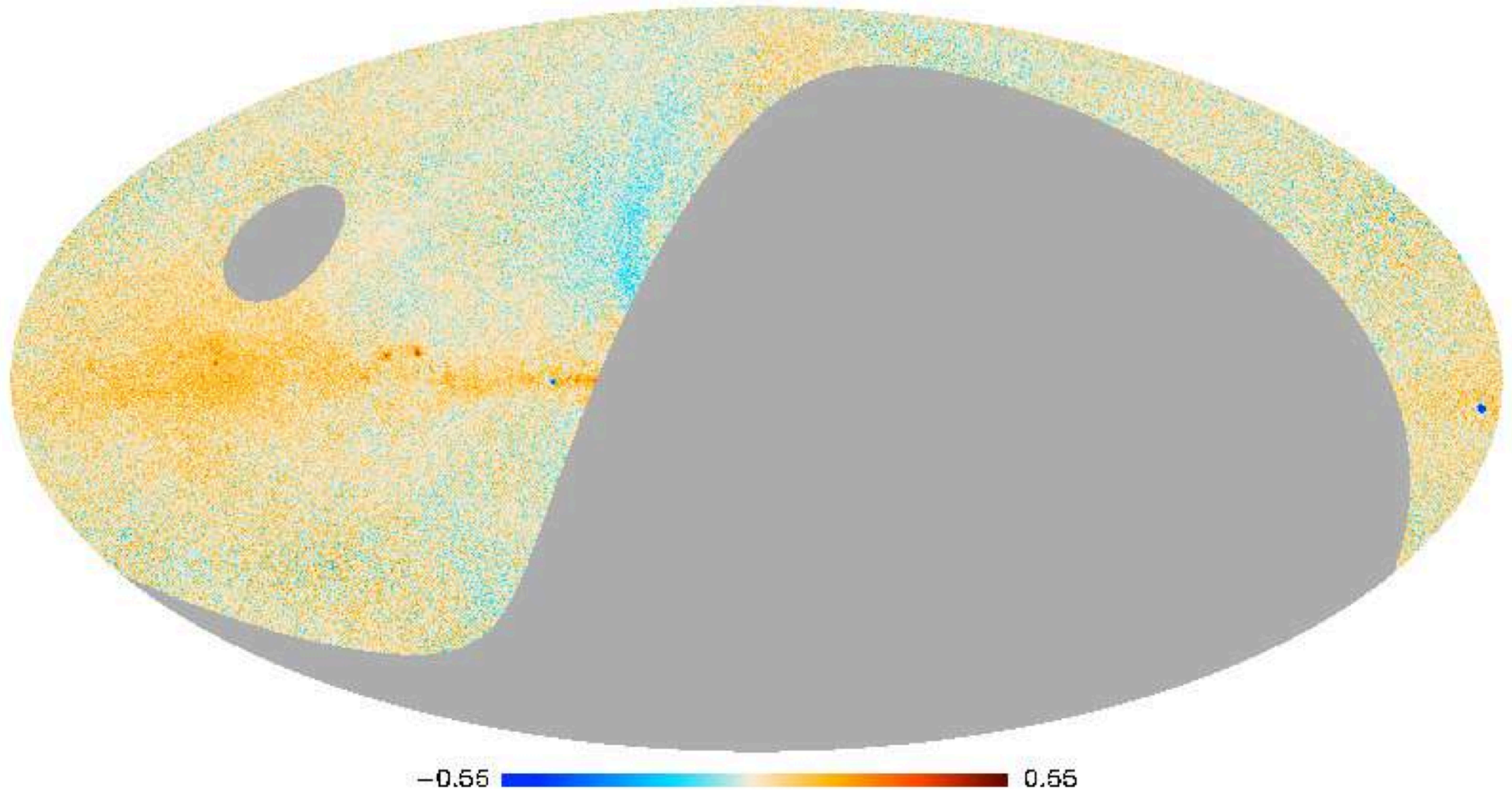


QUIJOTE Q 11 GHz



WMAP: Q 23 GHz

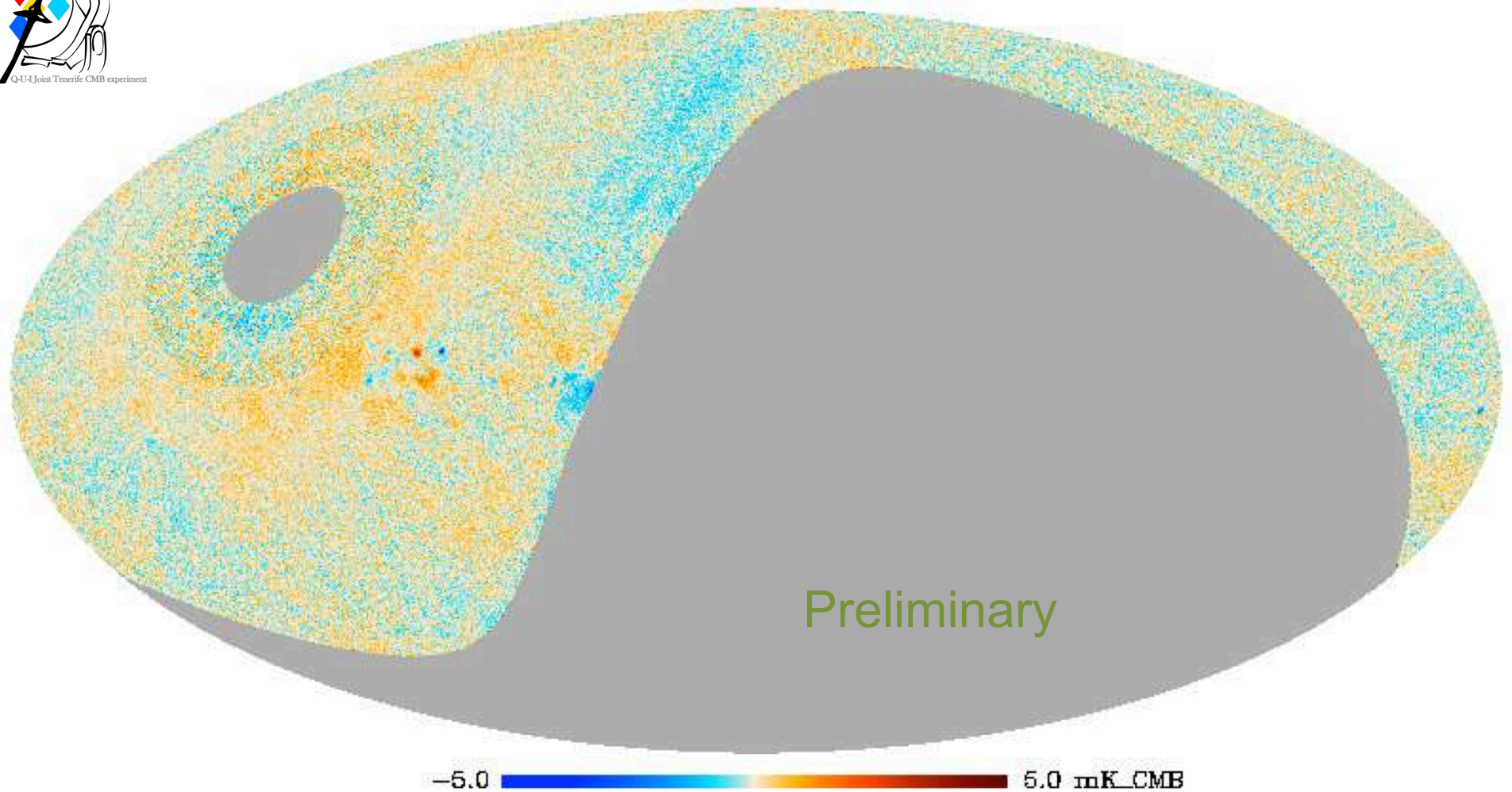
WMAP Q 23 GHz



Wide survey: U 11 GHz

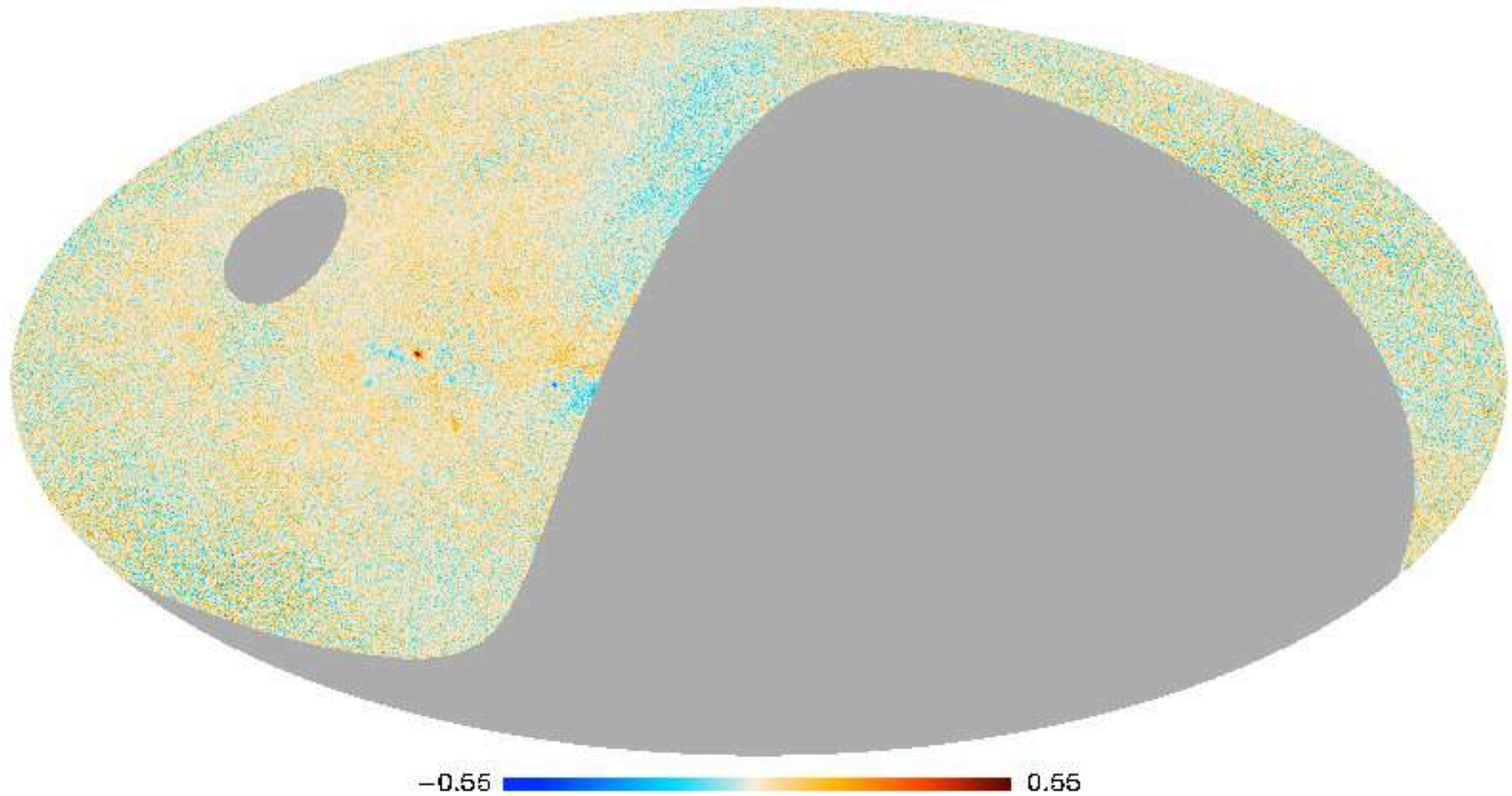


QUIJOTE U 11 GHz



WMAP: U 23 GHz

WMAP U 23 GHz



QUIJOTE: plans

➤ Upgrade of MFI

- Increasing sensitivity by a factor ~ 1.7
- Ready in 2 years (already funded)

➤ Extension of QUIJOTE to the South Hemisphere

- In collaboration with Wits University (South Africa)
- A prototype of an MFI pixel to be constructed (already funded)
- To be installed and tested at the 7.6m telescope at HartRAO \sim in 1.5 years
- Plan to install a complete replica of QUIJOTE if the observations with the prototype are successful (not funded yet)