



Exoplanet High Contrast Imaging: Space

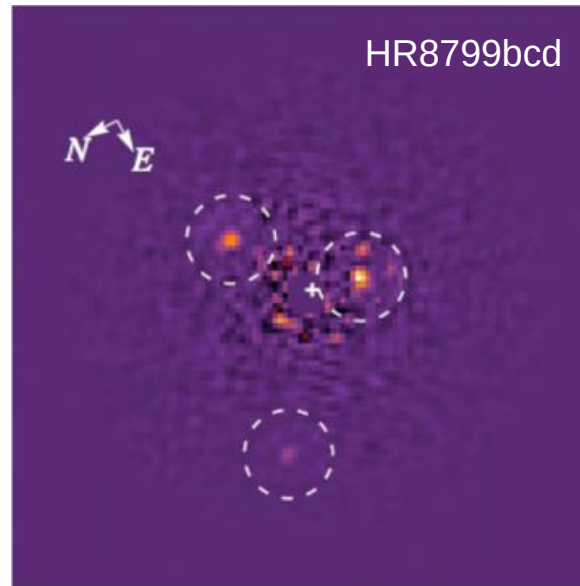
Eric Cady

Jet Propulsion Laboratory, California Institute of Technology

**KISS Short Course: Mastering the Wave - The Whys and Hows of
Exoplanet Imaging**

August 22, 2016

Overview



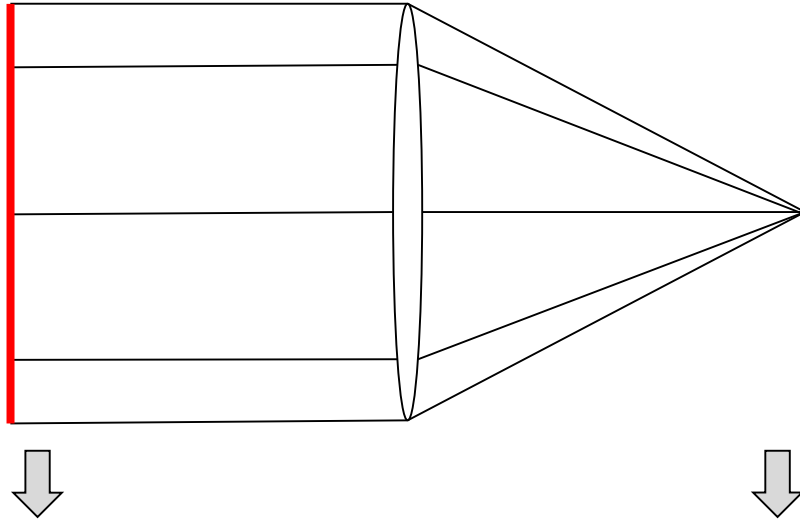
Soummer *et al.* 2011

Outline:

- A brief discussion of diffraction
- The Lyot coronagraph
- Imaging planets from space today
- Coronagraph technology development
- Starshades
- Imaging planets from space in the near future

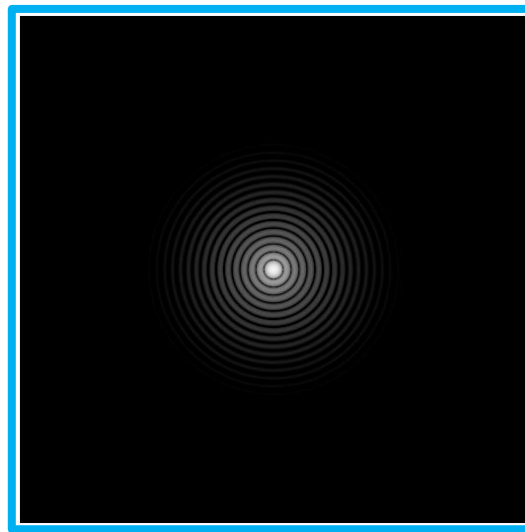
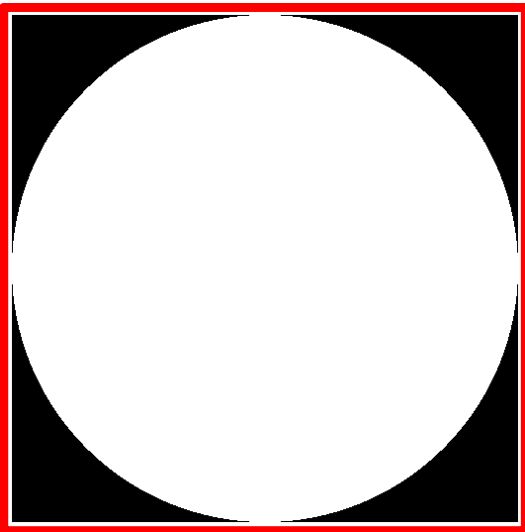


Diffraction in action

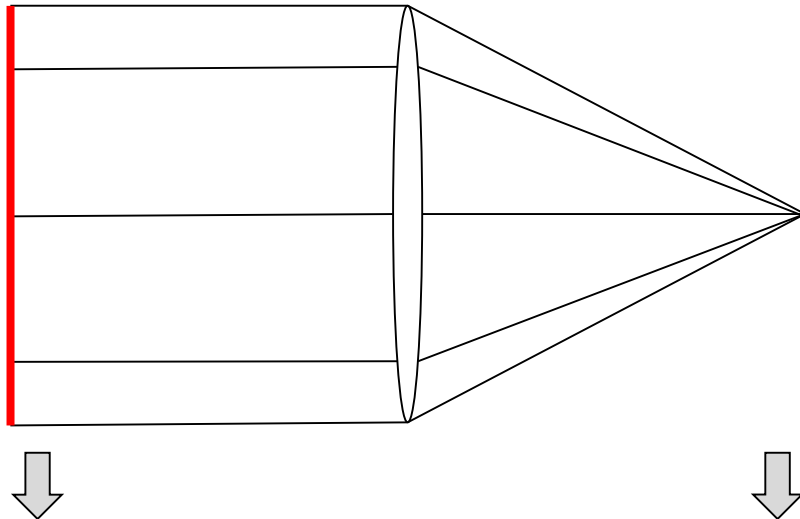


A star imaged by a telescope *never* focuses to a single point.

The shape of the star's image ("point spread function" or "PSF") can be directly related to the shape of the telescope aperture.

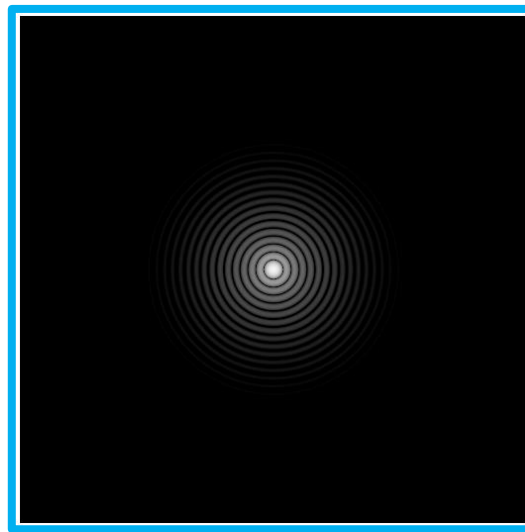
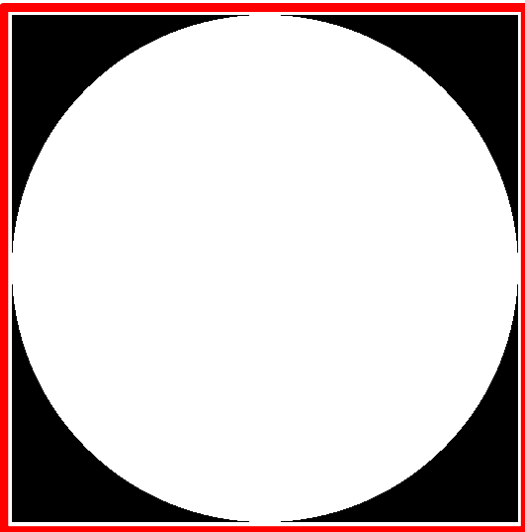


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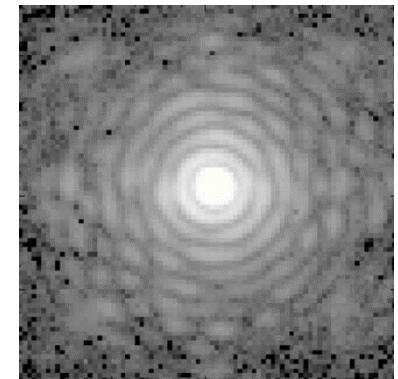


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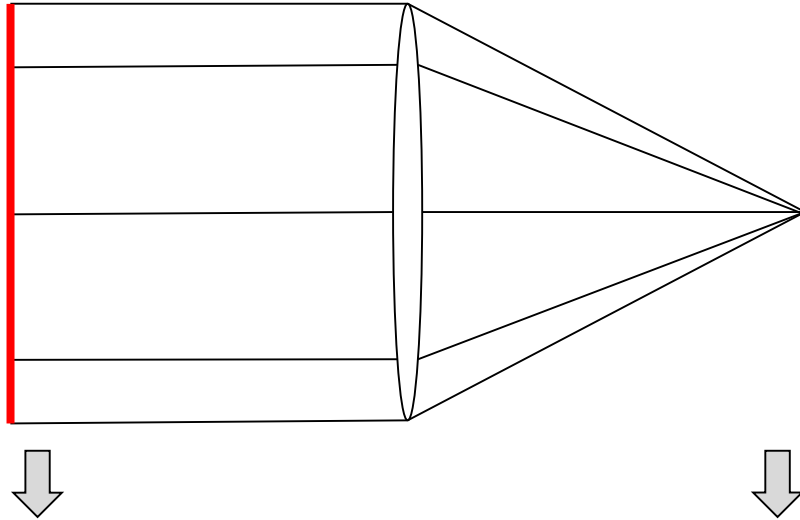


"Airy pattern"



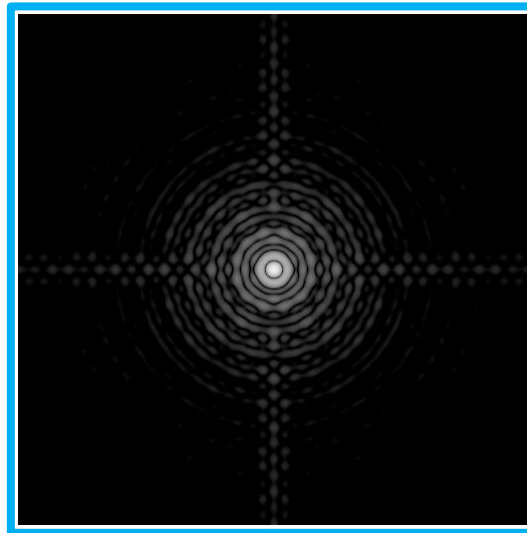
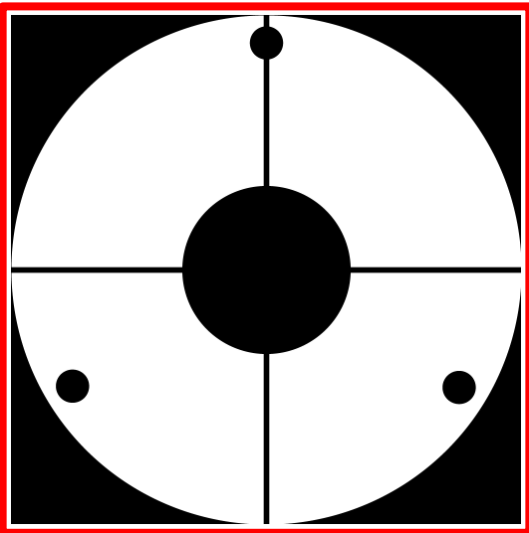
HD121107 with Palomar WCS,
Serabyn *et al.* 2007

Diffraction in action



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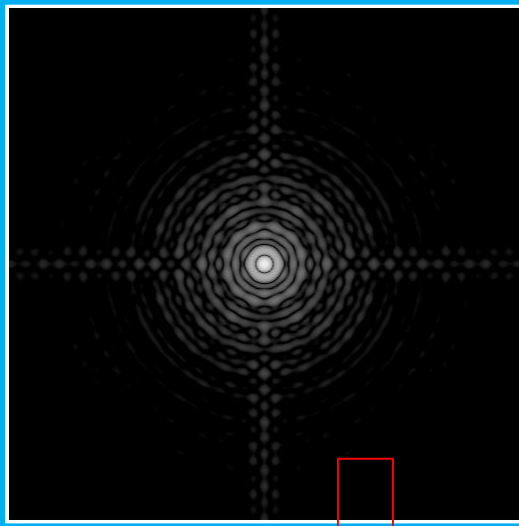
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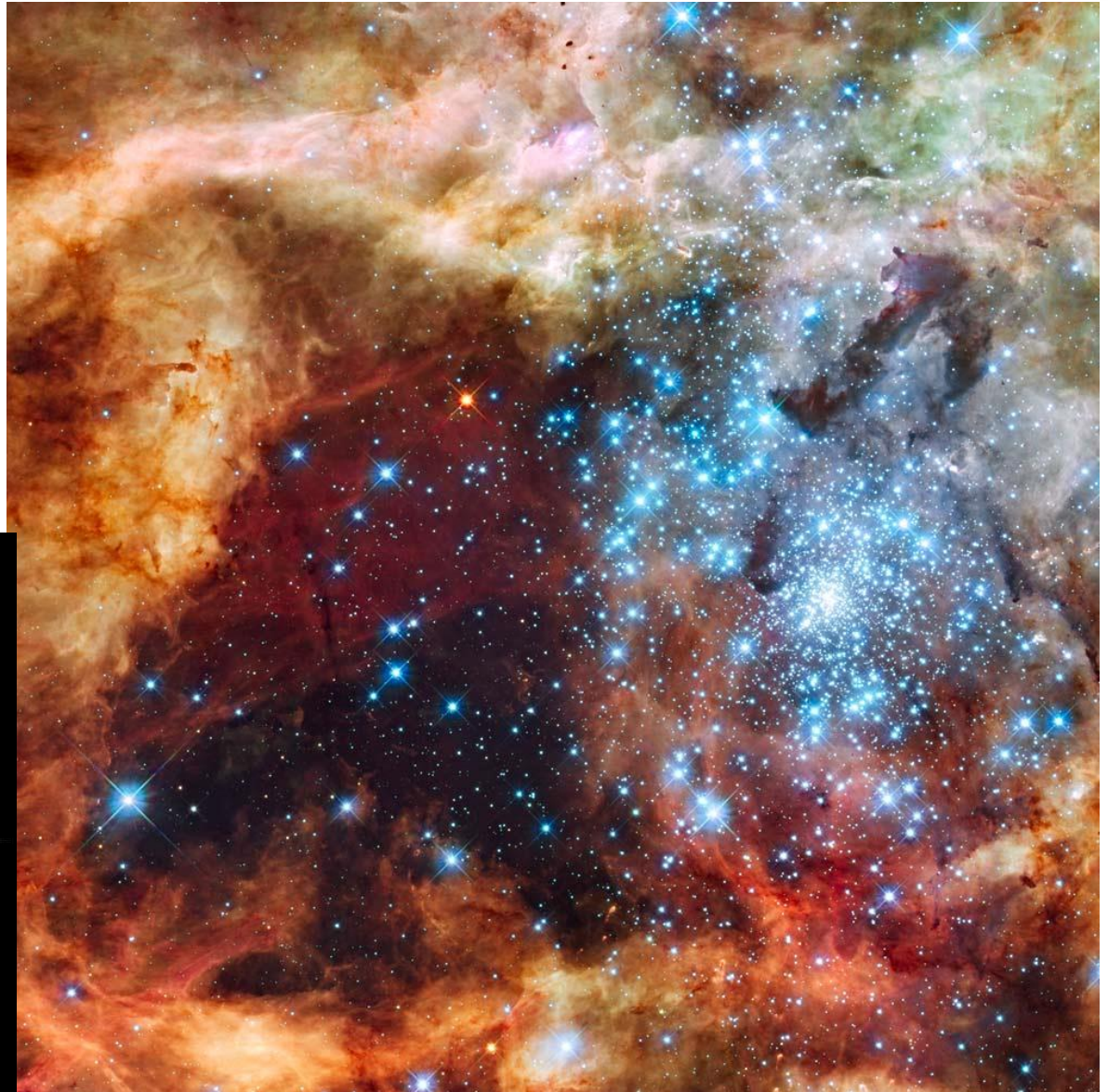
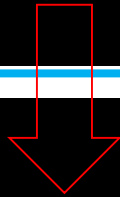
Hubble PSF



Diffraction spikes in Hubble



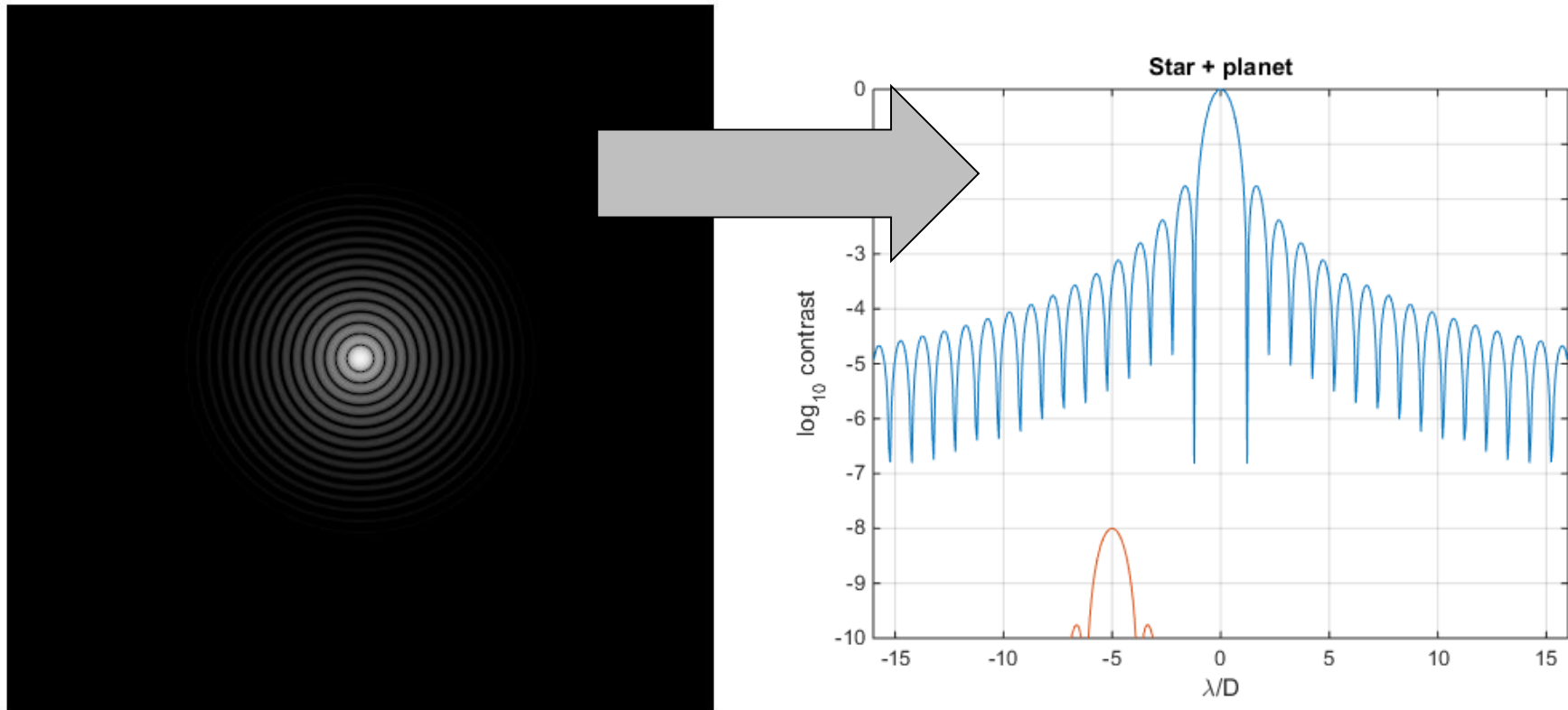
zoom out:



30 Doradus with Hubble WFC3, Credit: NASA, ESA, F. Paresce (INAF-IASF, Bologna, Italy), R. O'Connell (University of Virginia, Charlottesville), and the Wide Field Camera 3 Science Oversight Committee

The challenge: faint and close

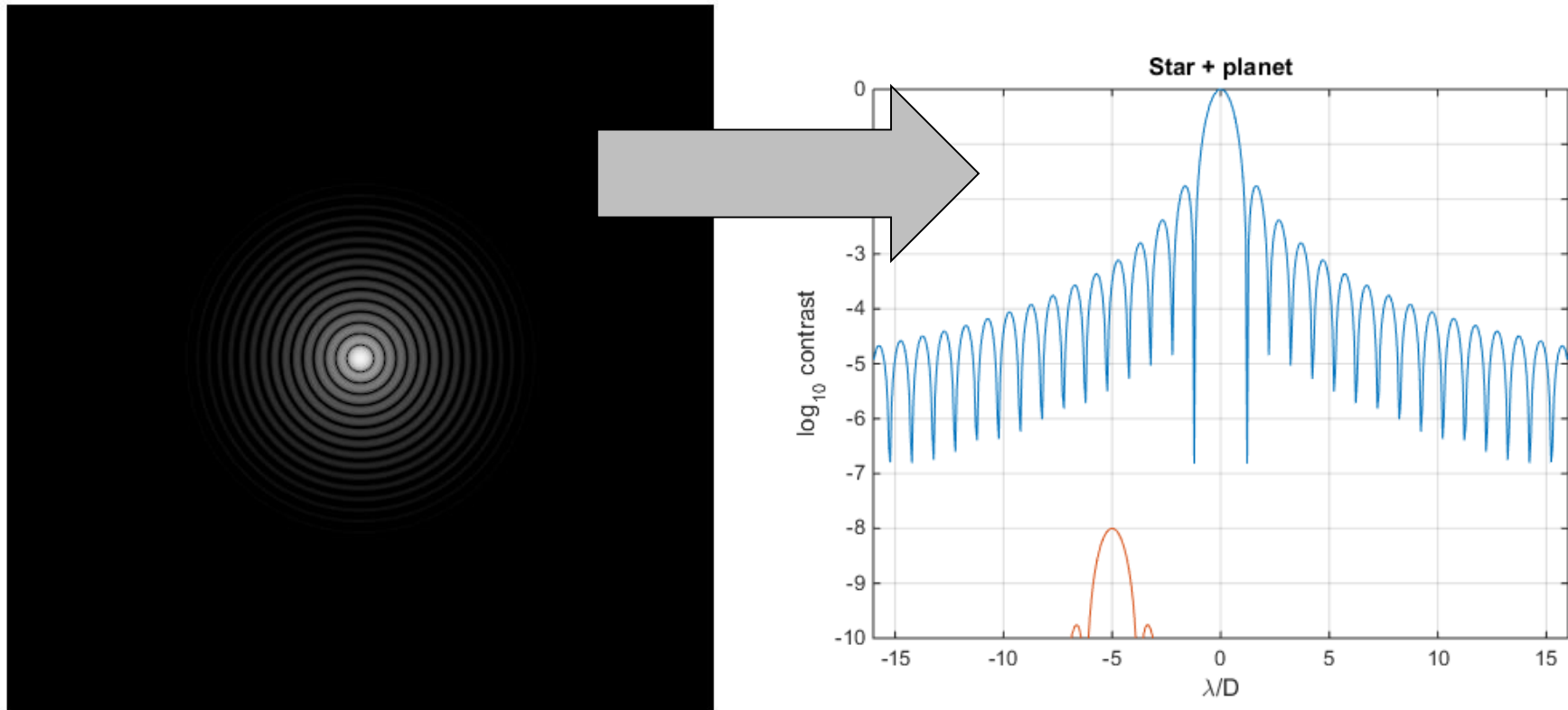
Unfortunately, most planets are buried under the “wings” of a telescope PSF.



The challenge of high-contrast imaging: *get rid of the diffracted starlight to see those faint nearby objects.*

The challenge: faint and close

Unfortunately, most planets are buried under the “wings” of a telescope PSF.



The challenge of high-contrast imaging: *get rid of the diffracted starlight to see those faint nearby objects.*

PSF depends on aperture → different telescopes may need different solutions!

Enter the coronagraph

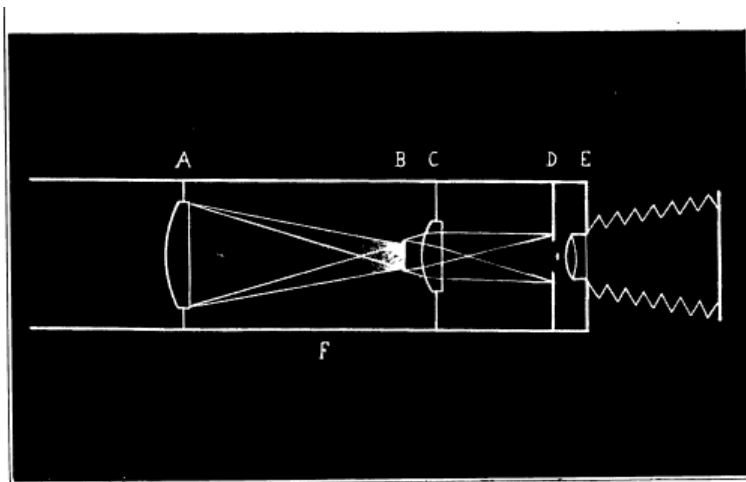
One method for doing this: the coronagraph

A class of instruments that:

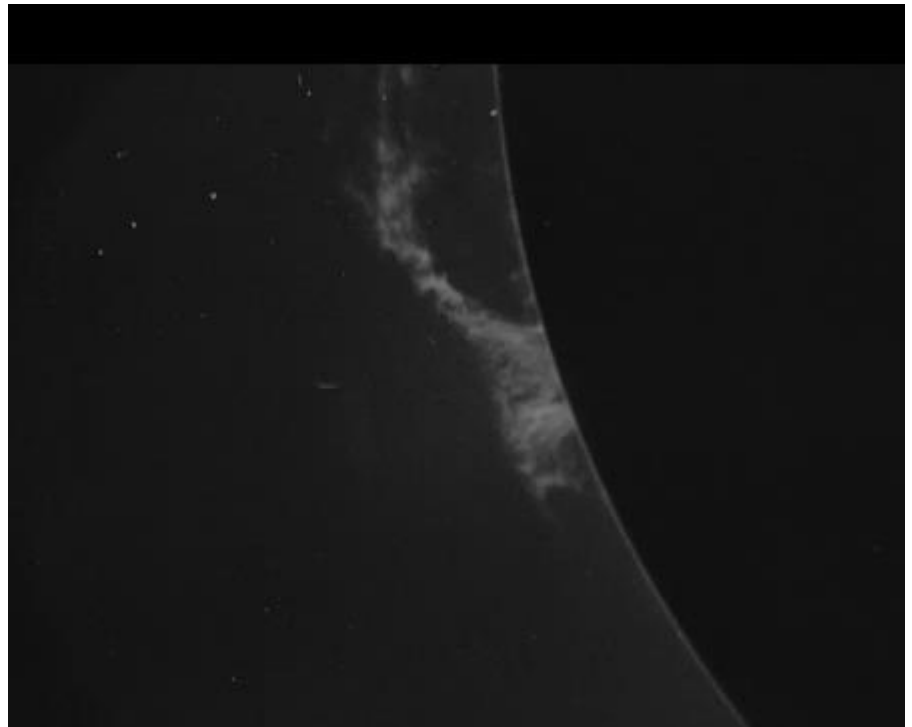
- use diffraction to block, redirect and cancel the starlight inside the telescope
- have a minimal effect on the planet light

First invented by Bernard Lyot to study the solar corona in 1930s

- another faint astrophysical feature right next to a bright one

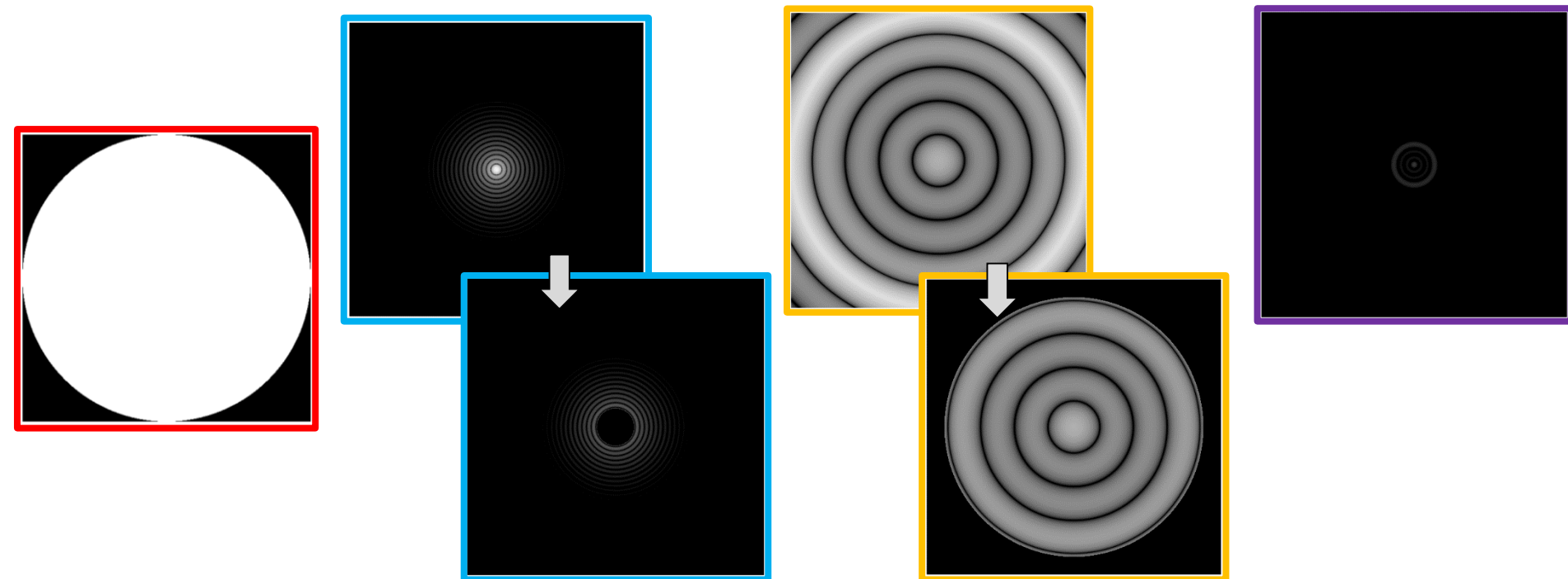
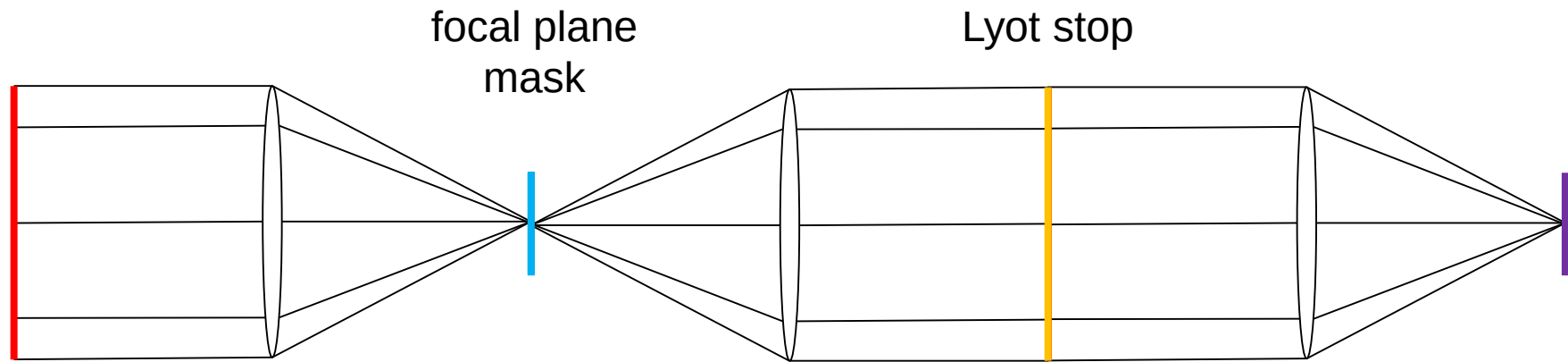


Lyot 1939



Lyot 1935

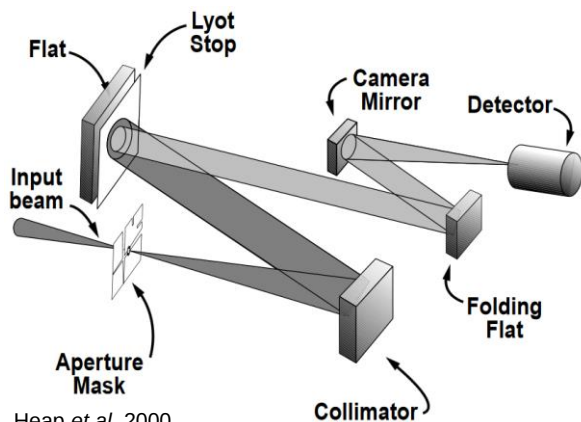
The “classical” Lyot coronagraph



Space coronagraphy, today

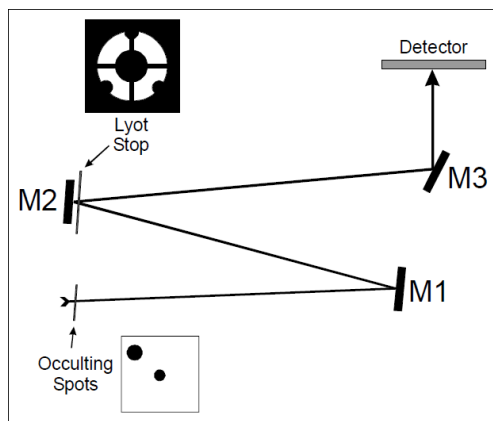
Hubble has had three Lyot coronagraphs used in its instruments to look at planets:

STIS



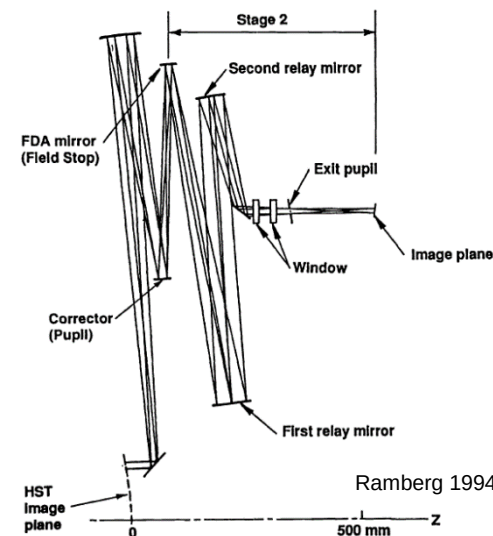
Heap *et al.* 2000

ACS/HRC

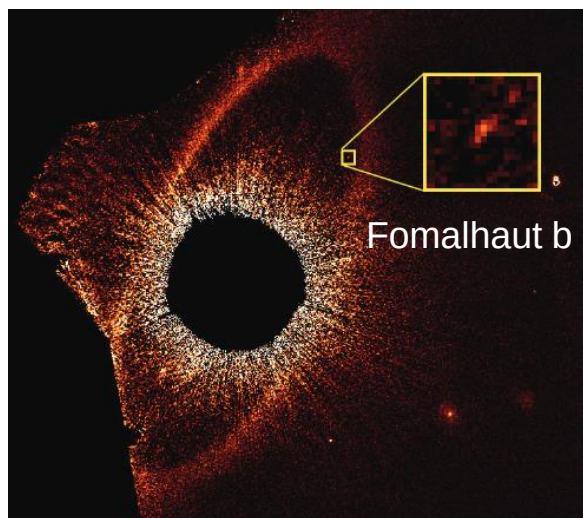


Krist *et al.* 2003

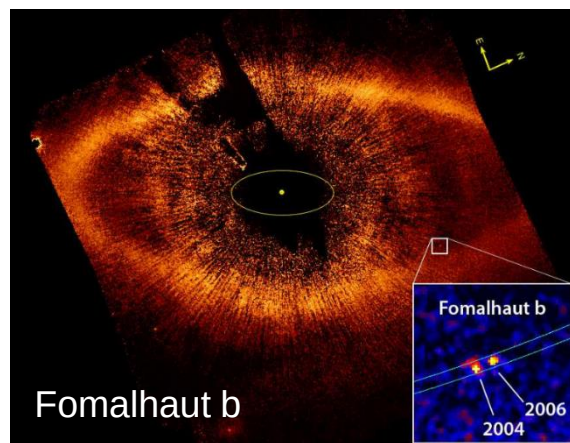
NICMOS



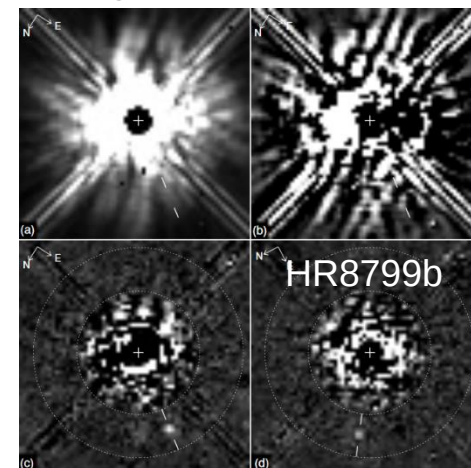
Ramberg 1994



Kalas *et al.* 2013



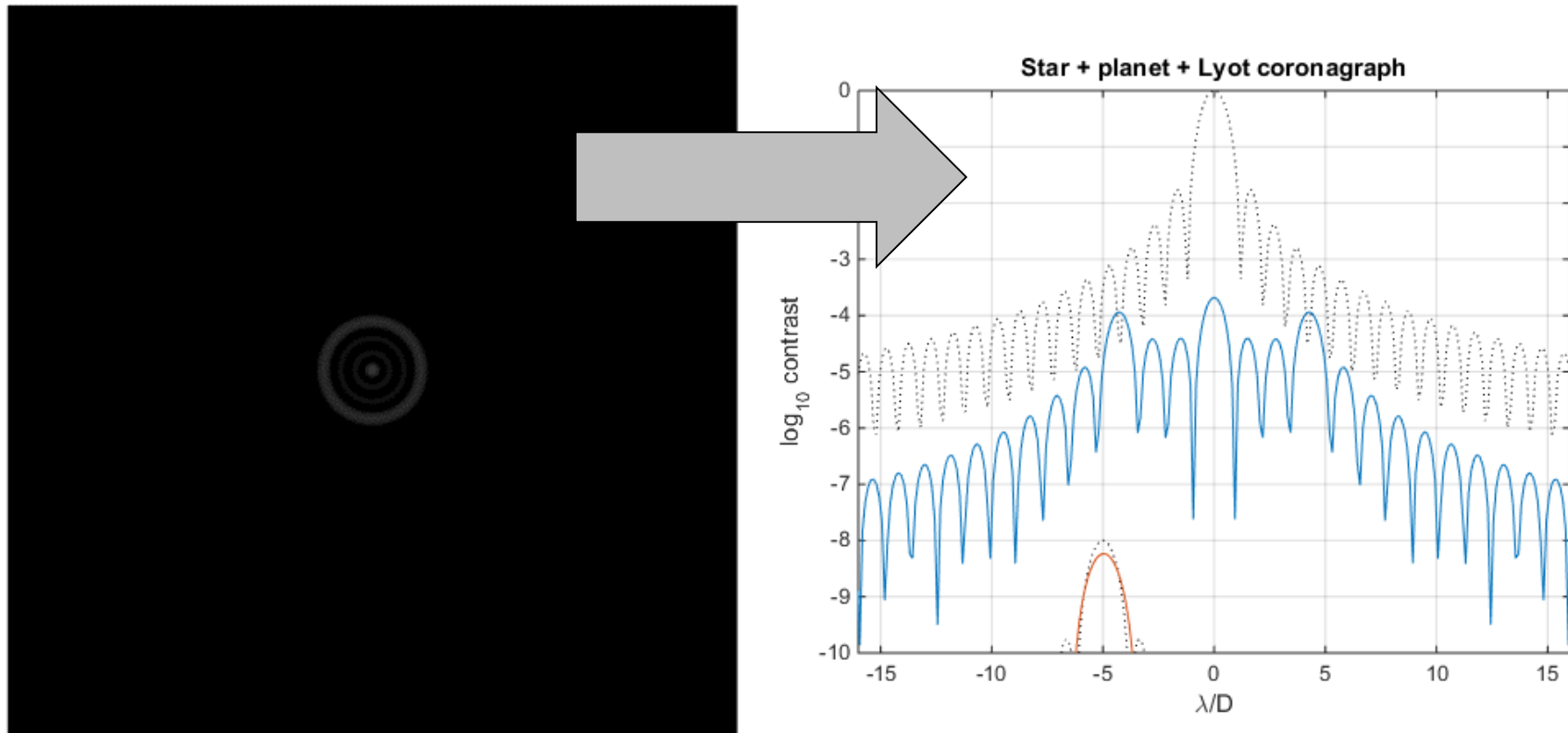
Kalas *et al.* 2008



Lafrènière *et al.* 2009

A slice through the PSF

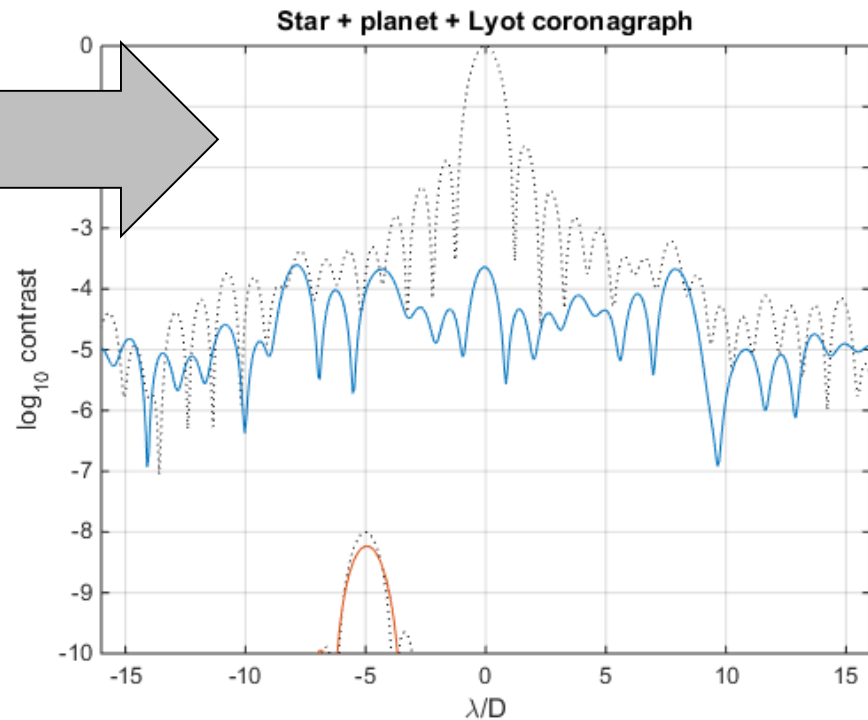
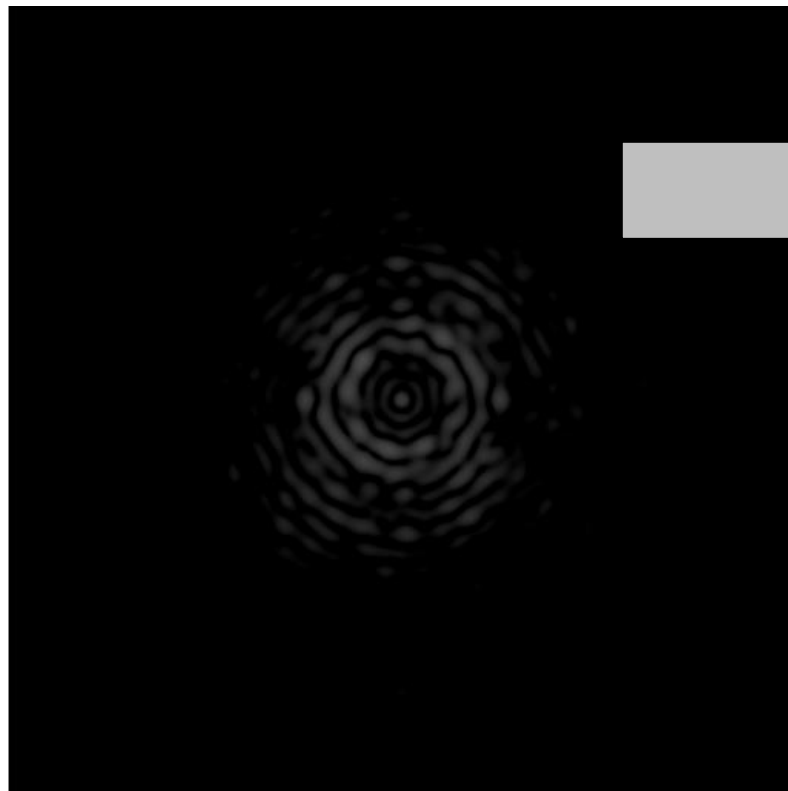
Let's look at a slice through that post-coronagraph PSF:



1. Wings of PSF are suppressed by mask + Lyot stop, not just center
2. Coronagraph affects planet slightly as well, but not as much as star
3. ...even given that, **our planet is still too faint for this coronagraph!**

It gets worse

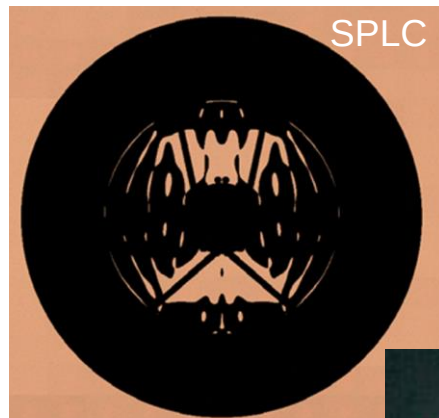
If our mirrors are not perfect:



- We get features (“speckles”) all over the place
- We don’t have those wings suppressed anymore

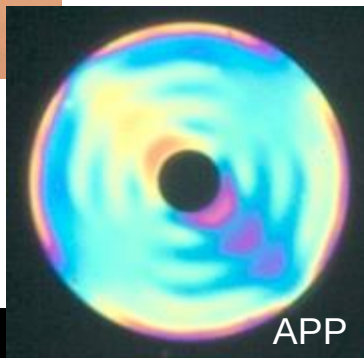
How do we fix this?

1. Better coronagraph designs

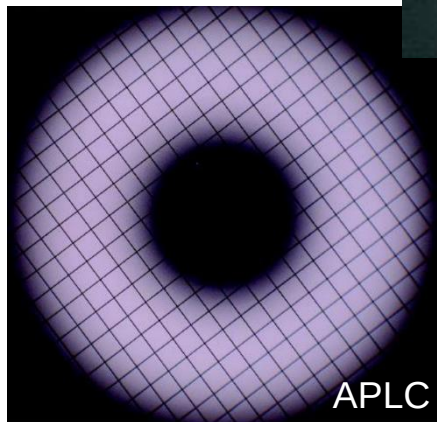


Balasubramanian *et al.* 2016

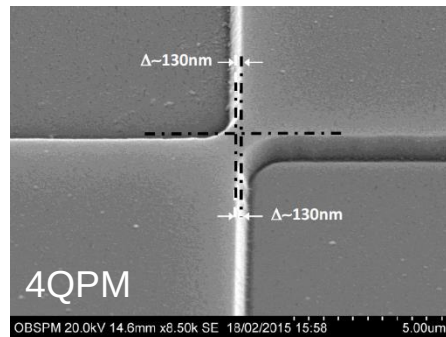
Change the
pupil



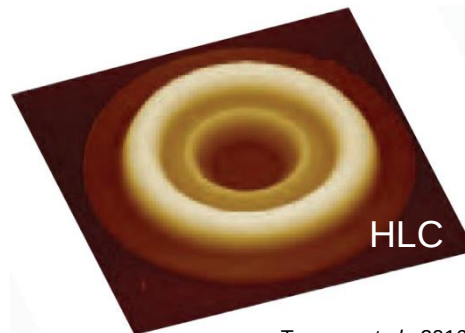
Otten *et al.* 2014



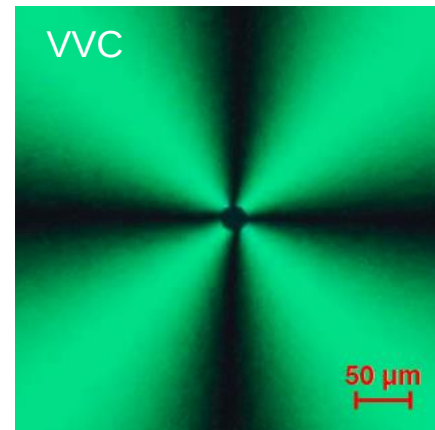
Macintosh *et al.* 2008



Bonafous *et al.* 2016



Trauger *et al.* 2016



Mawet *et al.* 2009

Change the
focal plane
mask



Guyon *et al.* 2012

Introduce
additional
optics

How do we fix this?

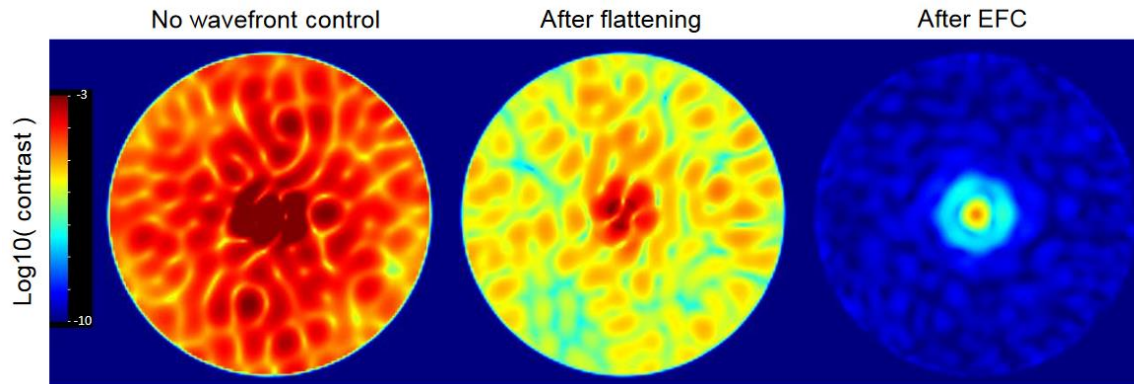
1. Better coronagraph designs
2. Wavefront control

Use wavefront sensors to measure how the starlight was changed (“aberrated”)

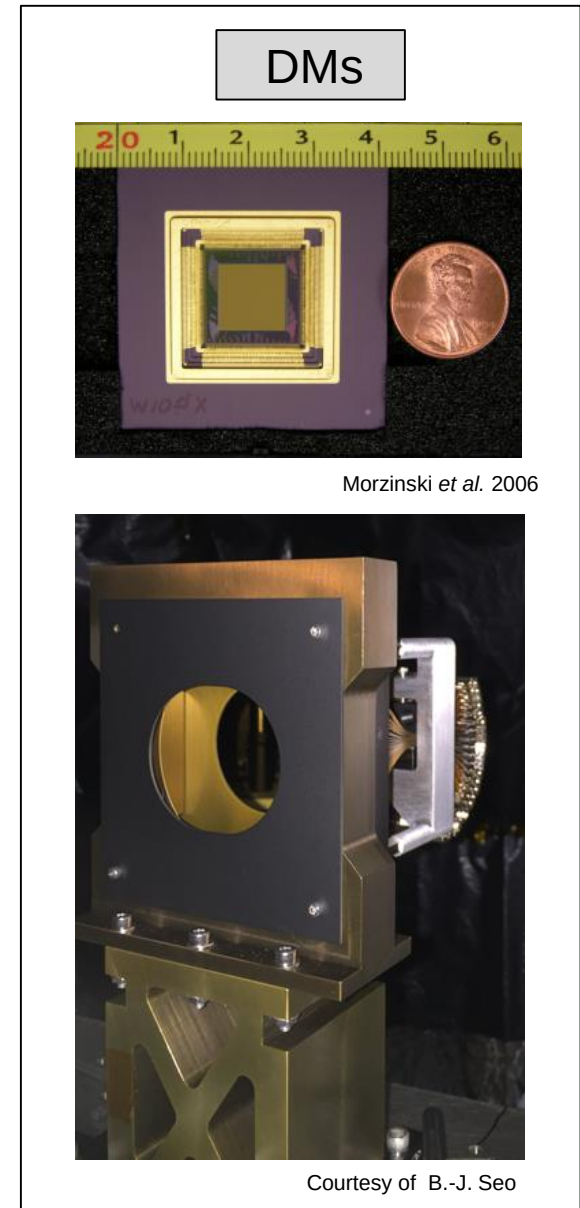
- Science camera is often best for this

Use deformable mirrors (**DMs**) to correct those aberrations

The combination is very effective at reducing speckles within a limited area near the star (“**dark hole**”)

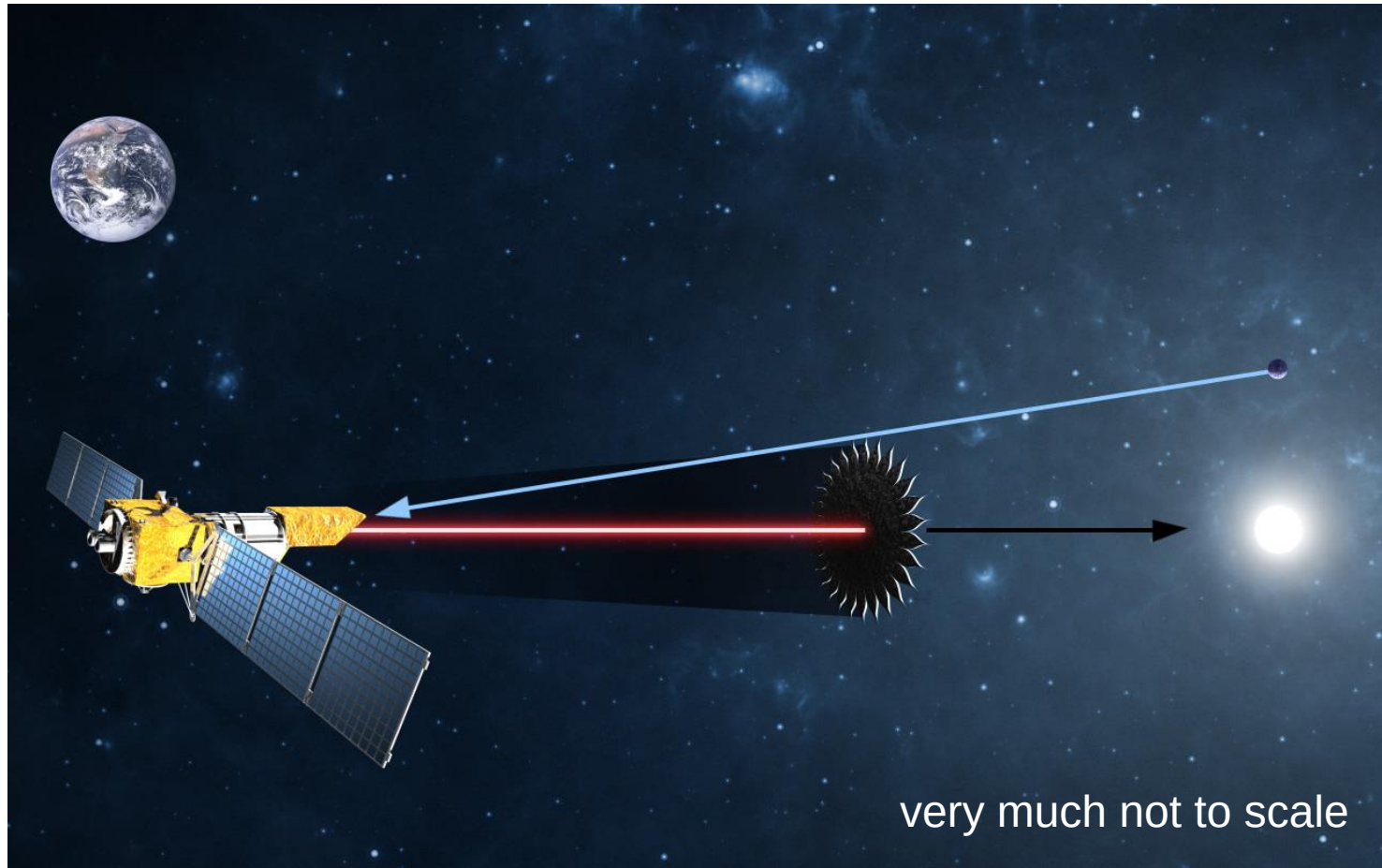


Krist, Nemati, and Mennesson 2015



Take it outside

An alternate approach: use a **starshade** to block the light before it enters the telescope. (“external coronagraph”)



Exo-S final report, 2015

Right size not to block the planet. No wavefront control needed!

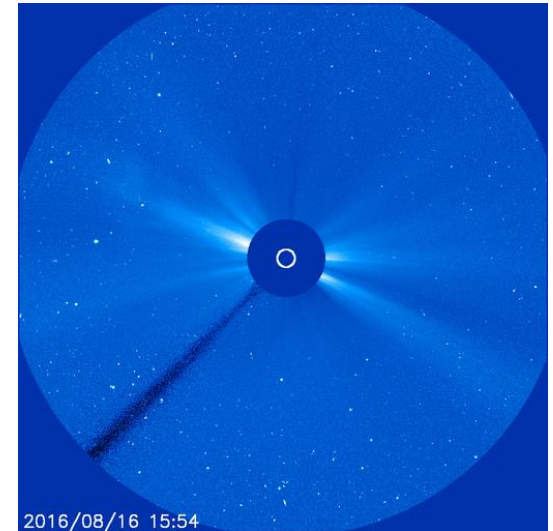
Starshades

Like coronagraphs, the starshade concept was developed for solar observations

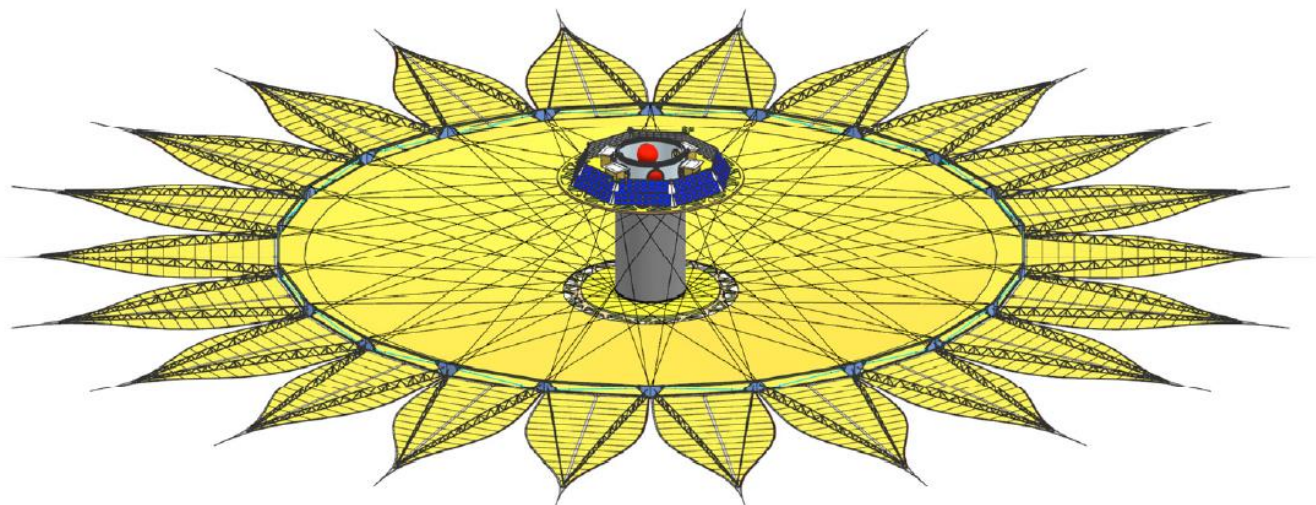
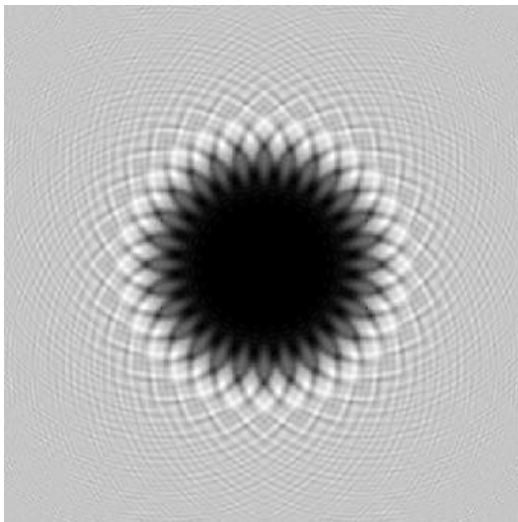
- LASCO on SOHO has two “external coronagraphs” right now looking at sun

Starshade is external to *telescope*

- Flown in formation at distances of tens of thousands of kilometers
- Creates a shadow where telescope must remain
- Edge is shaped to control diffraction



From C3 on LASCO realtime feed



Still a concept, but a promising one.

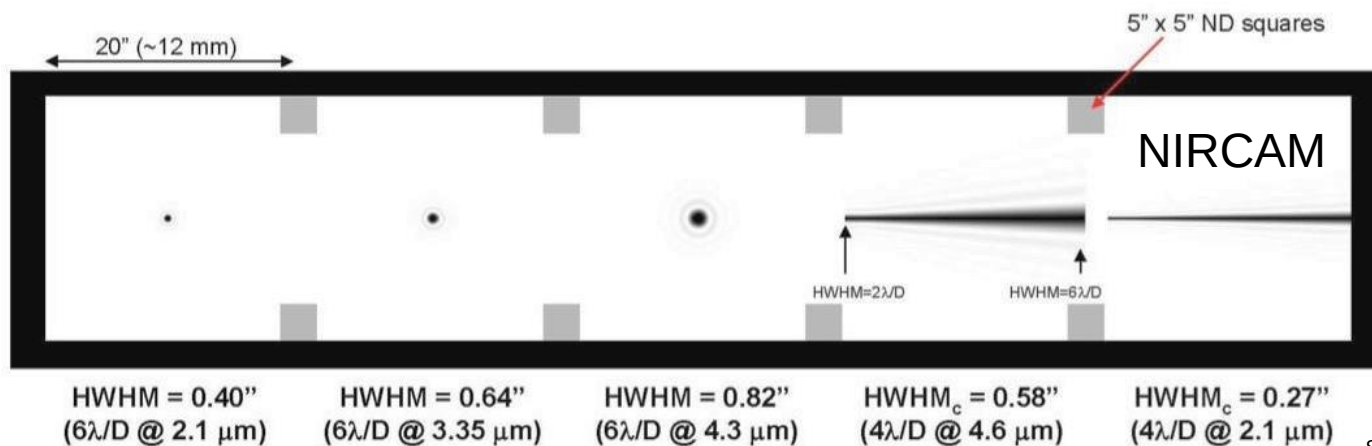
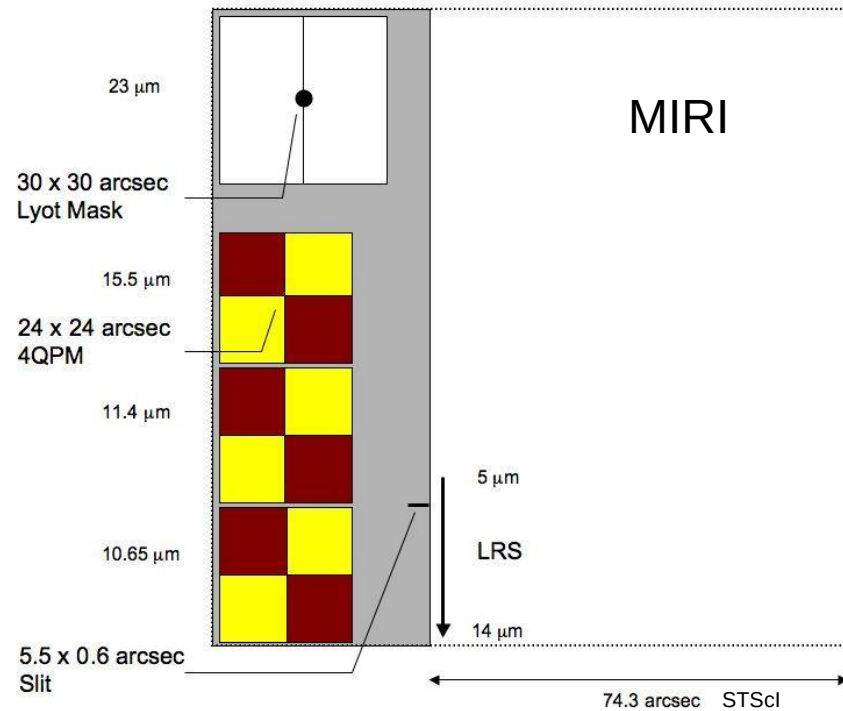


Future of space high-contrast imaging

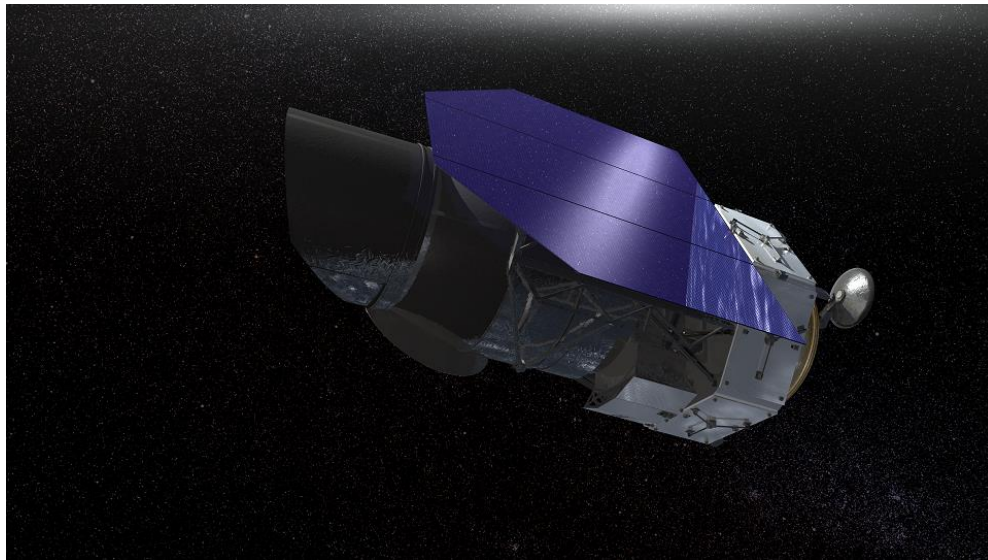


jwst.nasa.gov

JWST:
coronagraphs
in two
instruments

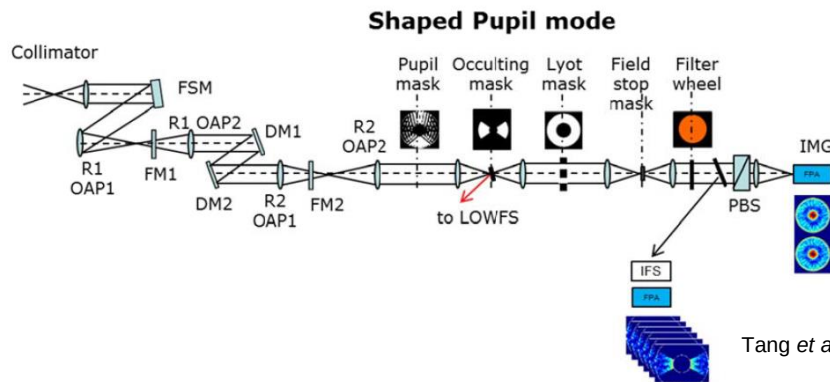
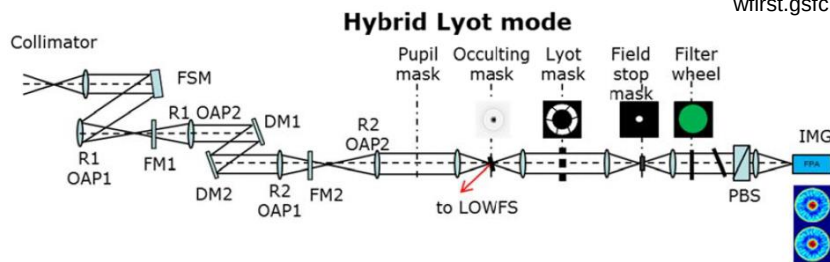


Future of space high-contrast imaging



WFIRST: dedicated coronagraph instrument

wfirst.gsfc.nasa.gov



Tang et al. 2015

Other possible missions:

- HabEx and/or LUVOIR?
- Starshade with WFIRST?

And of course, plenty of imaging to be done on the ground...