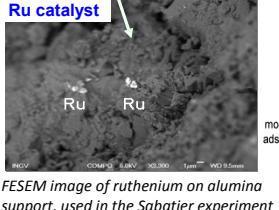
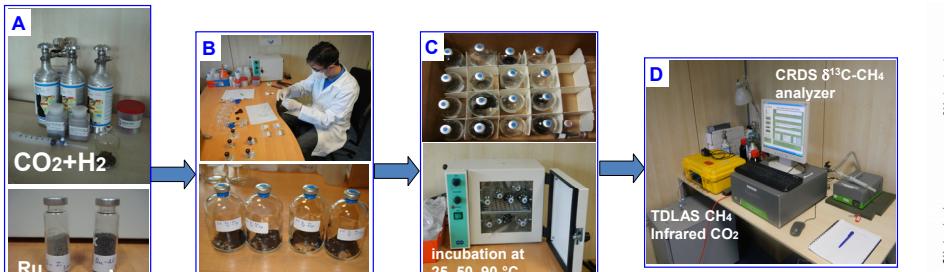


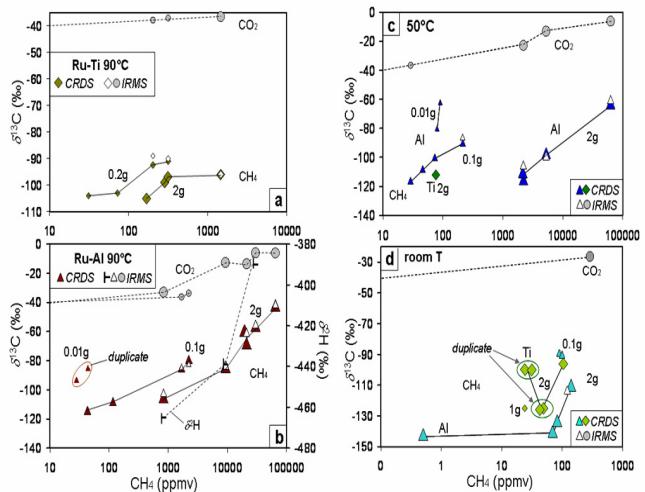
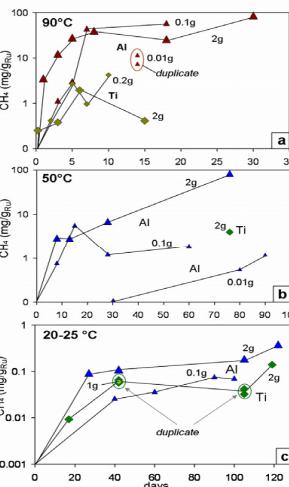
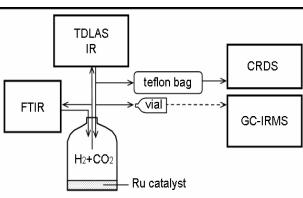
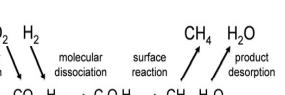
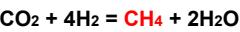
FAST ABIOTIC PRODUCTION OF METHANE AT TEMPERATURES <100°C

Fischer-Tropsch Type (FTT) reactions, e.g., the Sabatier synthesis between H₂ and CO₂, are considered a main source of abiotic methane (CH₄) on Earth and other planets. Several laboratory FTT experiments demonstrated abiotic CH₄ production at T > 200°C, by using Fe, Ni or Cr catalysts, simulating hydrothermal conditions in peridotite-hosted submarine systems. Nevertheless, Fe-Ni-Cr catalysts cannot support CH₄ generation at T < 100-150°C, which are those of continental serpentinization systems.

In 2014 we reported rapid production of considerable amounts of CH₄ (>800 ppmv in 155 mL bottles after 1 day, up to 6.5 vol% after 1 month) via Sabatier reaction at 90, 50 and 25°C, using concentrations of non-pretreated ruthenium equivalent to those occurring in chromitites in ophiolites or continental igneous complexes (Etiope and Ionescu, 2015; *Geofluids*, 15, 438-452).



SABATIER REACTION



In 2015 we repeated the experiments by using ¹³C-enriched CO₂ and we confirm fast production of CH₄ at percentage levels.

The experiments so far performed show that:

1. large CH₄ amounts can be produced in dry conditions below 100°C using extremely small quantities of ruthenium
2. under the same experimental conditions (<100°C), Fe, Ni and Cr oxides do not produce methane
3. low T FTT reactions can produce CH₄ with a large C isotope fractionation between CO₂ and CH₄, leading to relatively "light" (¹³C-depleted) CH₄, resembling microbial gas
4. CO₂-CH₄ isotope separation decreases over time and by increasing the temperature
5. minor amounts of C₂-C₆ hydrocarbons are also generated

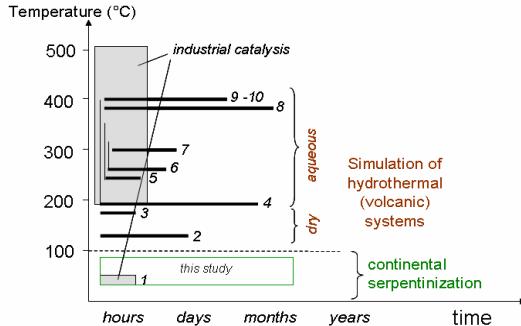
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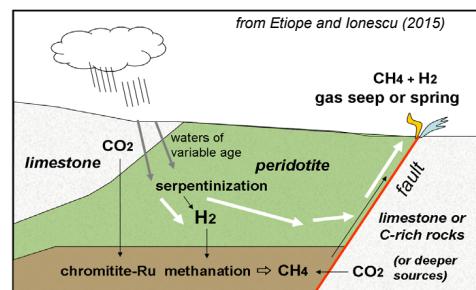
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LABORATORY FTT EXPERIMENTS - INDUSTRY vs GEOLOGY



Schematic plot of temperature and time scales considered in FTT experiments in geology and industrial catalysis. All geologic experiments were conducted at T > 100°C. Our study is the first to report CH₄ generation and its isotopic composition below 100°C, which is the T of continental serpentinization.

1. Thampi et al. (1987); Jacquemin et al. (2010); 2. Lancet et al. (1970); 3. Medina et al. (2000); 4. Horita and Berndt (1999); 5. Taran et al. (2010); 6. Ni and Yin (2011); 7. Taran et al. (2007); 8. Foustoukos and Seyfried (2004); 9. Fu et al. (2007); 10. Zhang et al. (2013). The "industrial catalysis" field is based on Wang et al. (2011) and references therein.



Schematic representation of CH₄ production at a land-based ultramafic serpentinization site. Methane-rich seeps and hyperalkaline springs are typically found in correspondence with faults, often in tectonic contacts between peridotites and limestones or carbon-rich metasedimentary rocks.

Our laboratory data are compatible with the isotopic patterns of naturally occurring CH₄ in land-based seeps and hyperalkaline springs. Our experiments suggest that Ru-enriched chromitites could potentially generate methane at low T on Earth and other planets. For example, since ruthenium was reported in Martian meteorites, low T Sabatier reaction may produce abiotic CH₄ in Martian Ru-bearing rocks (Etiope, Ehmann and Schoell, 2013, *Icarus*, 224, 276-285).