

MMIC Array Receivers and Spectrographs Workshop 2

23-24 March 2009



Focal Plane Array Program at

NRAO
P. R. Jewell, K. O'Neil

NRAO

Atacama Large Millimeter/submillimeter Array
Expanded Very Large Array
Robert C. Byrd Green Bank Telescope
Very Long Baseline Array



NRAO Focal Plane Array & Camera Program

- Long-Term Vision:
 - Technical Development for Square Kilometer Array
 - Beam Forming Arrays (Low Frequency through $\sim\lambda$ 1 cm)
 - Increase FOV for interferometers – ALMA, EVLA
- Near-Term Vision
 - Scientific Enhancement for GBT through Large Format Arrays
 - Bolometer Cameras at 3 mm
 - Spectroscopic Systems at Key Frequency Bands
 - 3 mm (W-Band)
 - 1.3 cm (K-Band)
 - Others (e.g, 7 mm)
 - Testbed for SKA work – beam forming arrays



Status & Goals

- GBT Pathfinder Instruments
 - MUSTANG Bolometer Camera (64-pixels) – Early Science
 - 7-pixel K-Band (1.3 cm) Array under construction (for spectroscopy)
- Beam Forming R&D work
 - Collaborative program with BYU
 - MRI grant in progress
 - Instrumentation tests on 20 Meter telescope at GB
- MRI proposal to expand MUSTANG to 256 pixels submitted
- Proposal for 1000-element bolometer camera developed
- Concept for large format (e.g. 100-element) 3 mm array for GBT under development



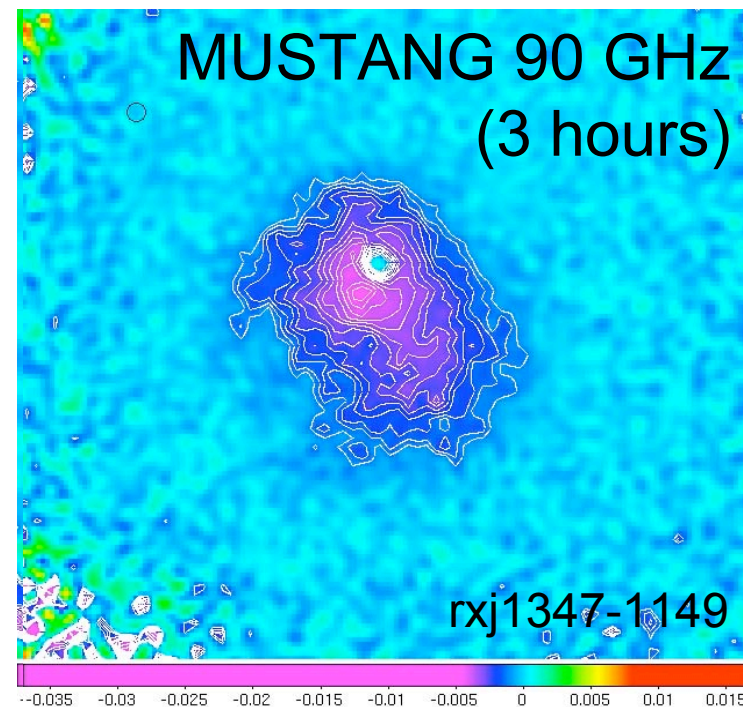
MUSTANG (64 pixel bolometer array)

Resolution	9" (FWHM)
Beam spacing	4"
N-pixel	8x8
Current sensitivity	$T_{\text{sys}}=140\text{K}$
Target sensitivity	$T_{\text{sys}}=28\text{K}$
Bandwidth	18 GHz

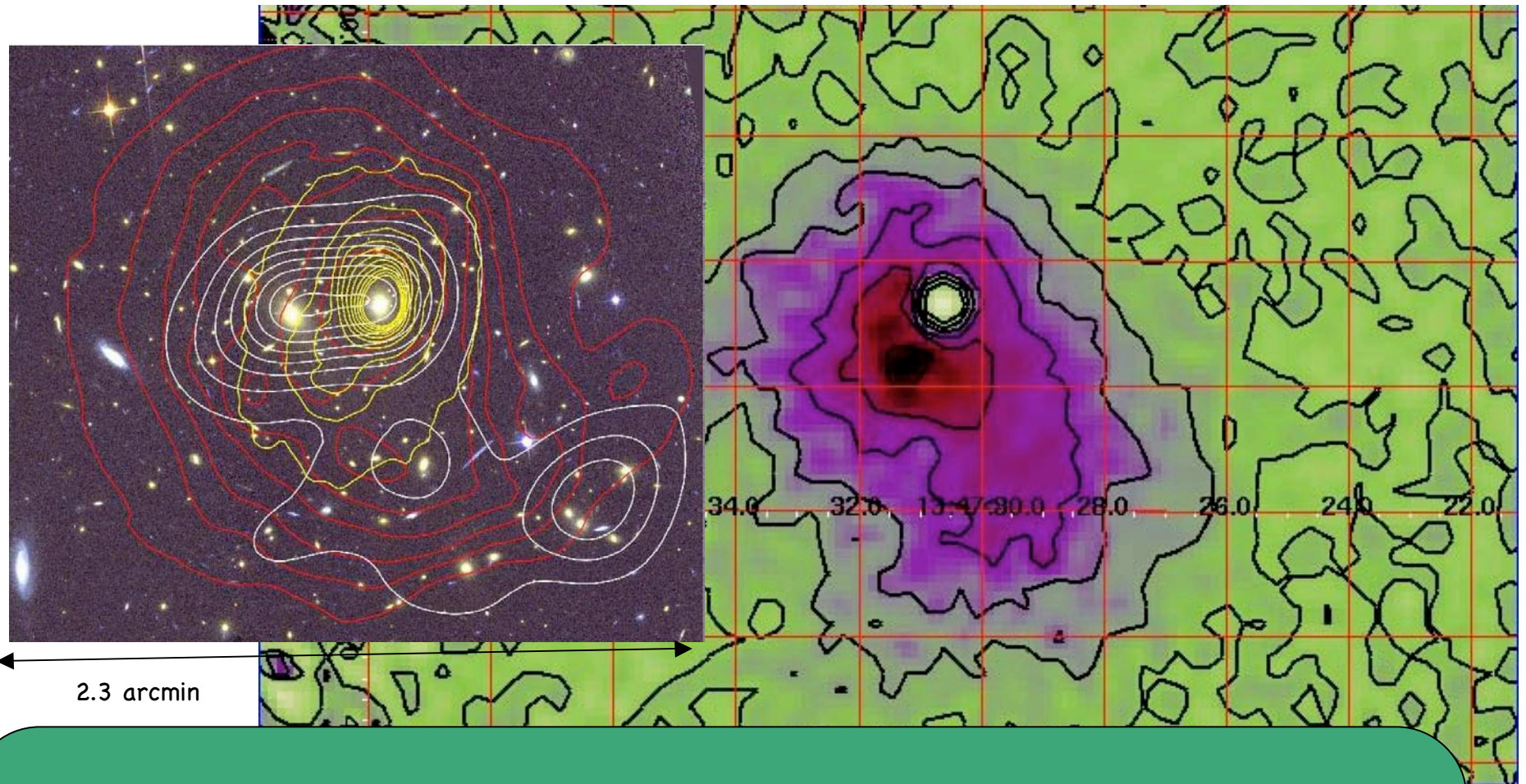
- Currently 5x from photon noise
- Arrays 1x photon noise exist
- 4x increase in Npix ~\$500k
- 16x increase in Npix~\$3M



- 120 detectors on GBT ~ ALMA for 3mm continuum mapping speed



- First MUSTANG+GBT Detection of SZ!
Bodes well for use of SZ as a detailed probe of ICM physics [Mason et al.] 4



Lensing Mass Map
 Galaxy Density
 Xray Surface Brightness

Optical + Xray (Left)

Bradac et al 2008

MUSTANG SZ (Right)

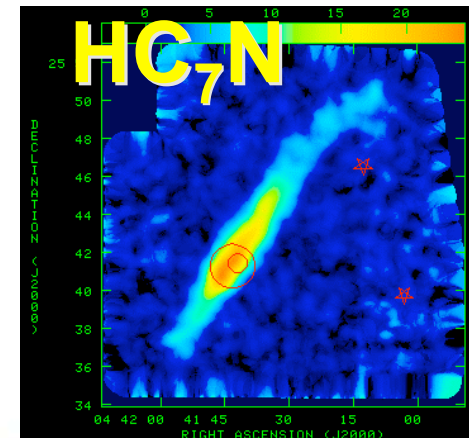
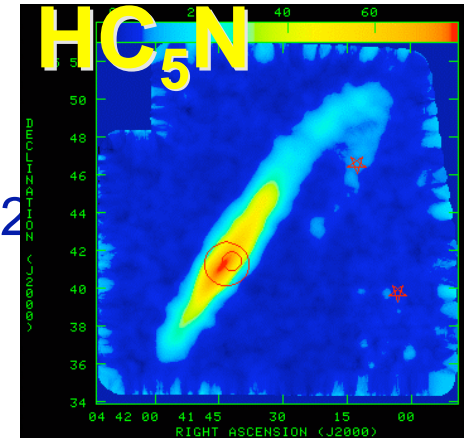
Similar angular scales

GBT Spectral Line FPAs

- 7-pixel prototype K-band Array to be released in late 2012
 - First element tested on the sky – results excellent



- Next step involves increasing the number of pixels
 - 100+ pixel W-band (3 mm) array
 - 61-pixel K-band (1.3 cm) array
- Signal transmission and processing infrastructure required
 - Need increased data transmission capability
 - Need new signal processing system



Large Format FPAs

- Building a large FPA has no technological blocks, but.....
- Full cost of first large array is very expensive:
 - R&D into:
 - Packaging the system
 - Calibration algorithms to take advantage of the multiple beams
 - Data displays both for quick look (while observing) and data reduction
 - Need digitized data transmission system
 - Current system is \$30k/line; Need R&D into something less expensive
 - New backend- FPGA technology would be straightforward to use
 - Estimated cost for first, large-format FPA is \$25-\$30M infrastructure
Want to reduce this cost through R&D (w/ partners)
Subsequent arrays would be ~1/4 cost.
NRAO is seeking partners for such projects.

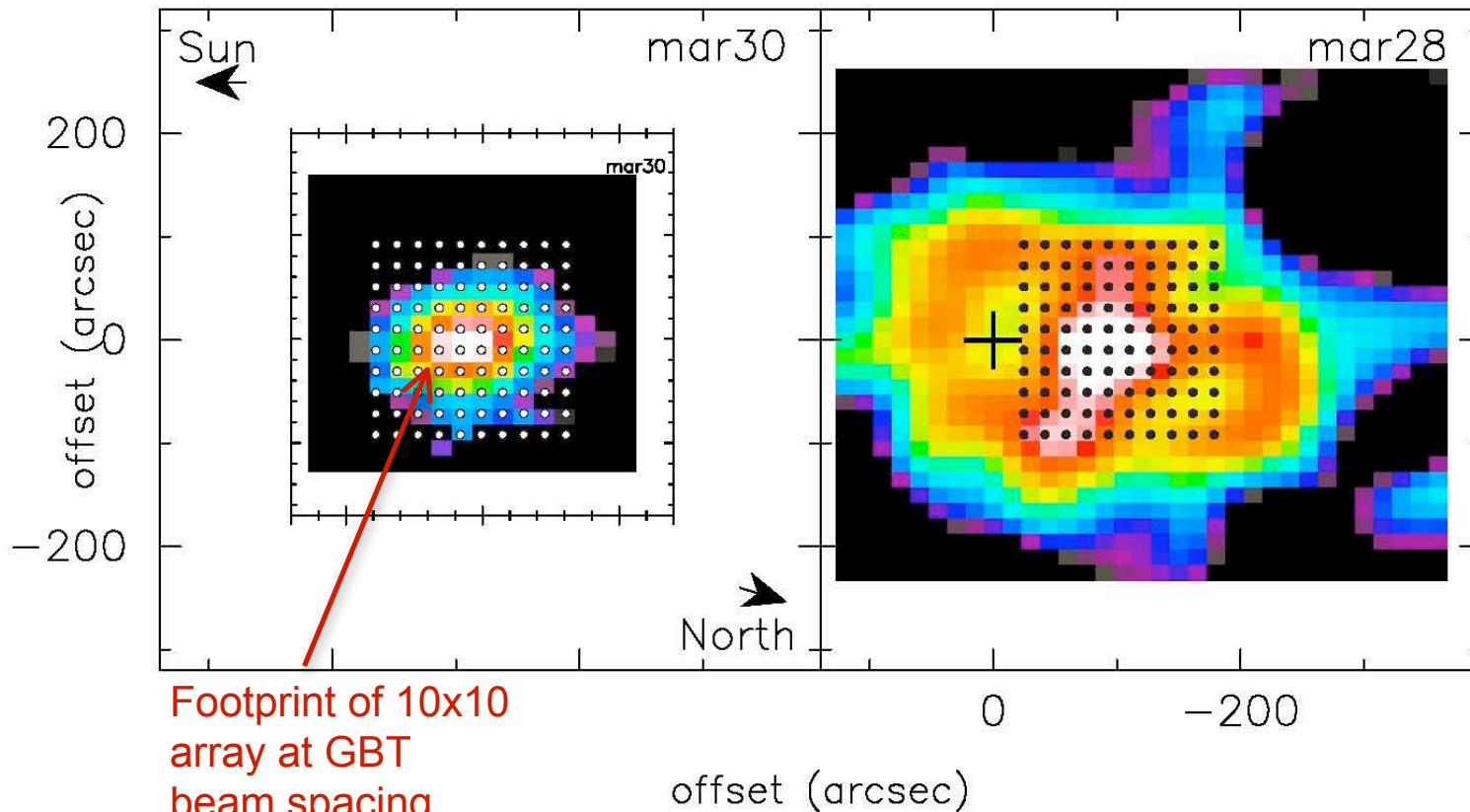


Science Application – Molecular Line Imaging of Comets

Cometary line emission is time-variable, with low surface brightness (large, filled aperture will be optimal). One molecular species gives rise to another (e.g. HCN \rightarrow HCO+), so rapid, sensitive imaging is key.

HCN J=1-0

HCO+ J=1-0



Footprint of 10x10 array at GBT beam spacing (image resolution will be x7 higher)

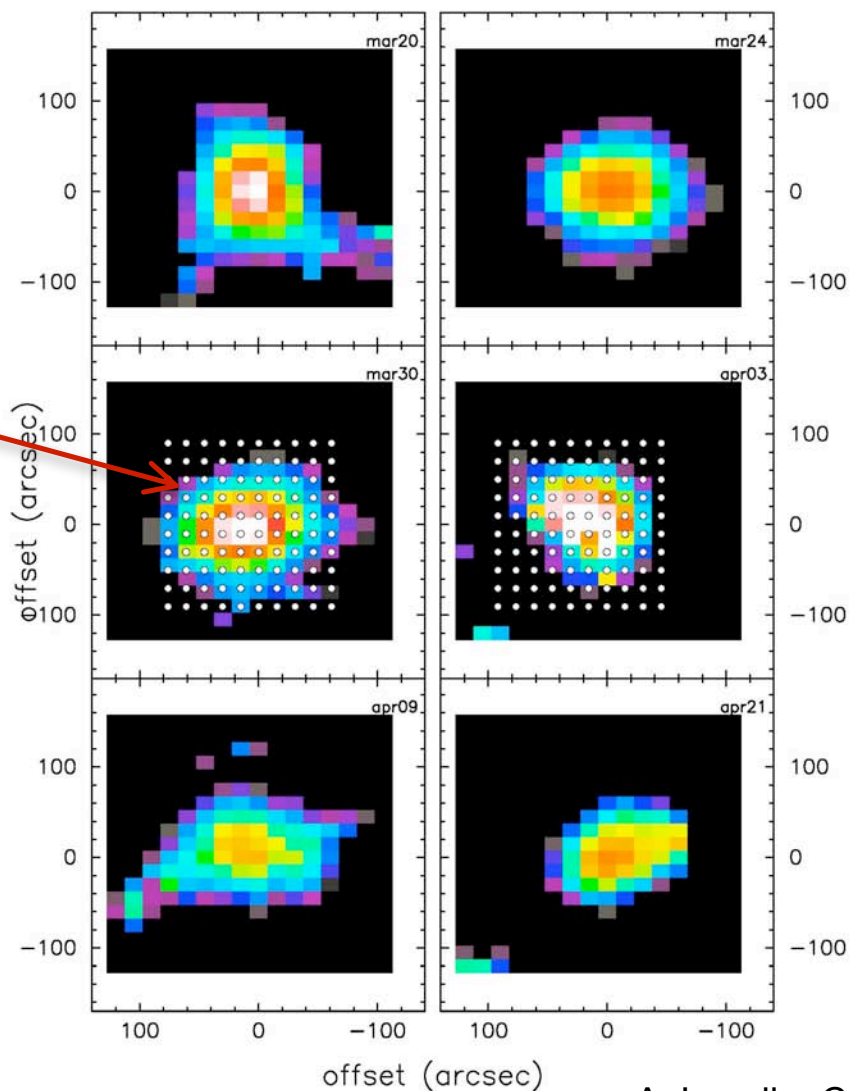
A. Lovell – Comet Hale-Bopp / FCRAO



Science Application – Molecular Line Imaging of Comets

HCN J=1–0 maps

Footprint of 10x10 array at GBT beam spacing (image resolution will be x7 higher)

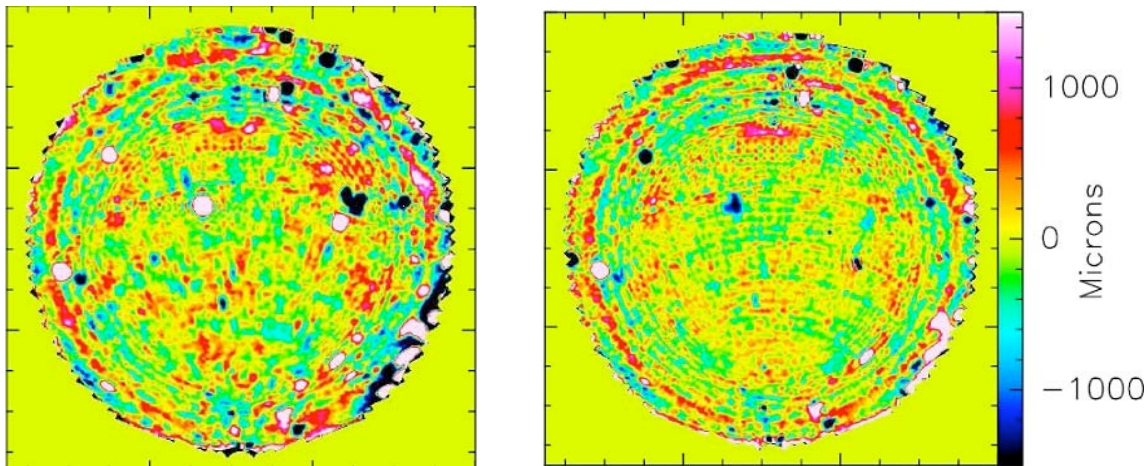


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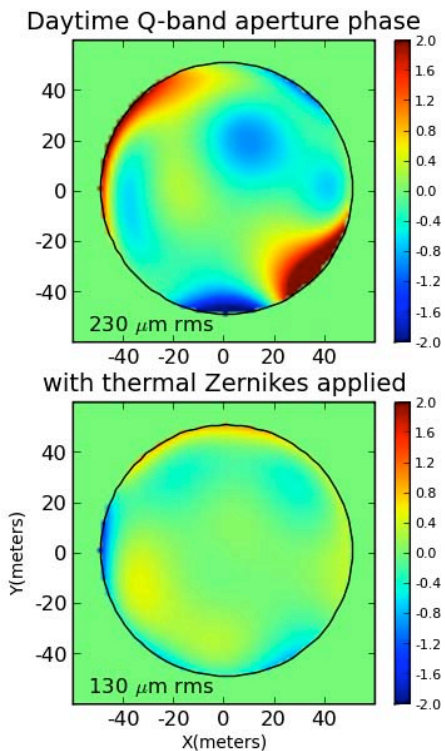


Optimizing the GBT for 3 mm Operation

- Holography project has reached 15% aperture efficiency at 90GHz
- Goal of 35% (20%) aperture efficiency at 90 (115) GHz



Phase-coherent Holography



Out-of-Focus (OOF) Holography

Optimizing the GBT for 3 mm – Continued: Dynamic Scheduling System (DSS)

- $\geq 20\%$ of Green Bank's weather is usable at high frequencies
 - GBT high frequency time is highly oversubscribed
 - Need to ensure the best possible use of a limited resource
 - GBT DSS will maximize science output at high frequencies
- GBT DSS will be released Fall 2010 in prototype form with enhanced versions to follow



Next Steps

- NRAO is writing “Requests for Information” papers for the Decadal Survey (due April 1) that will emphasize the FPA program
- Developing science and technology drivers for a 3 mm (W-Band) Array
- Actively seeking partners!
 - Contact Karen O’Neil in Green Bank

