



# Optical and Infrared Remote Sensing – from orbit and in situ

Wendy Calvin, Keck Workshop

Unlocking the Climate Record in Mars' Polar Layered Deposits

August 9, 2017

- Color and/or BW Imagery
  - Mariner 9, Viking
  - MGS- MOC (WA, NA)
  - Odyssey - THEMIS
  - MRO- HiRISE, MARCI, CTX
- Spectroscopy
  - MGS- TES, MRO-MCS
  - MRO-CRISM & Mars Express- OMEGA



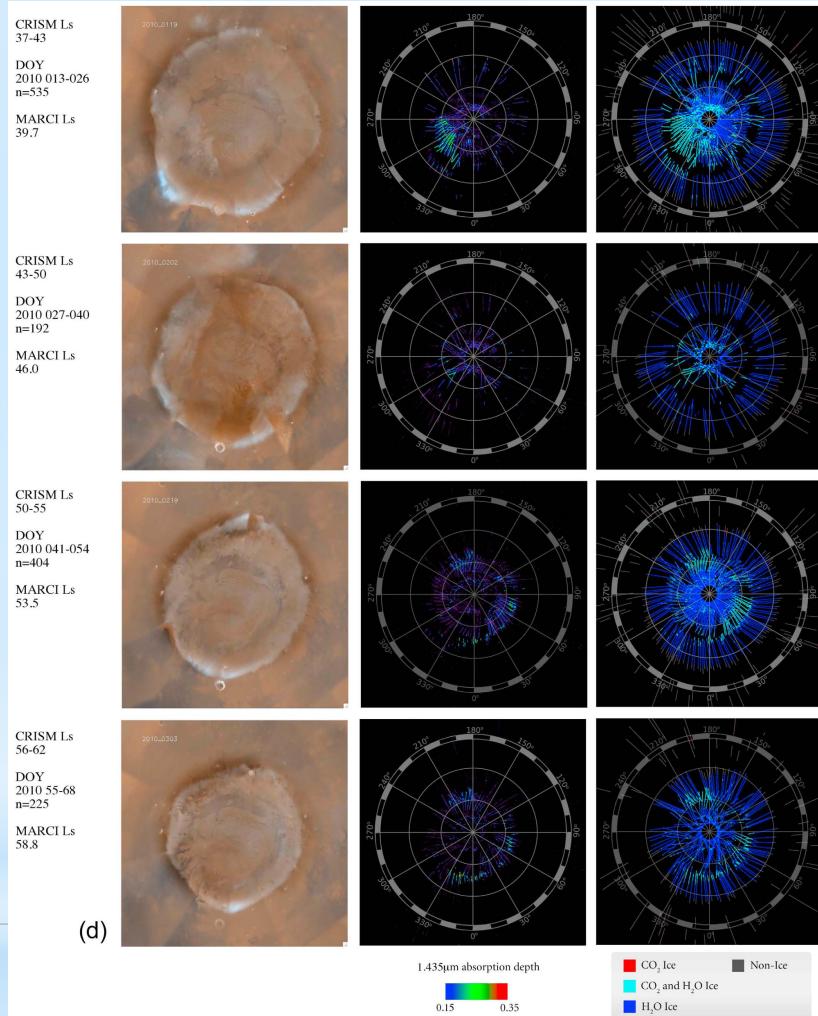
## Types of observations

- Annual Processes
  - Seasonal cap advance and retreat
- Interannual Variability
  - Mass wasting, accumulation changes in RSPC
  - Mobility of high and low albedo deposits
- Composition and Physical state
  - Grain size or optical path length, grain coherence
  - H<sub>2</sub>O vs CO<sub>2</sub> vs non-ice materials
- Stratigraphy (see Shane & Patricio's Short Course Talks)

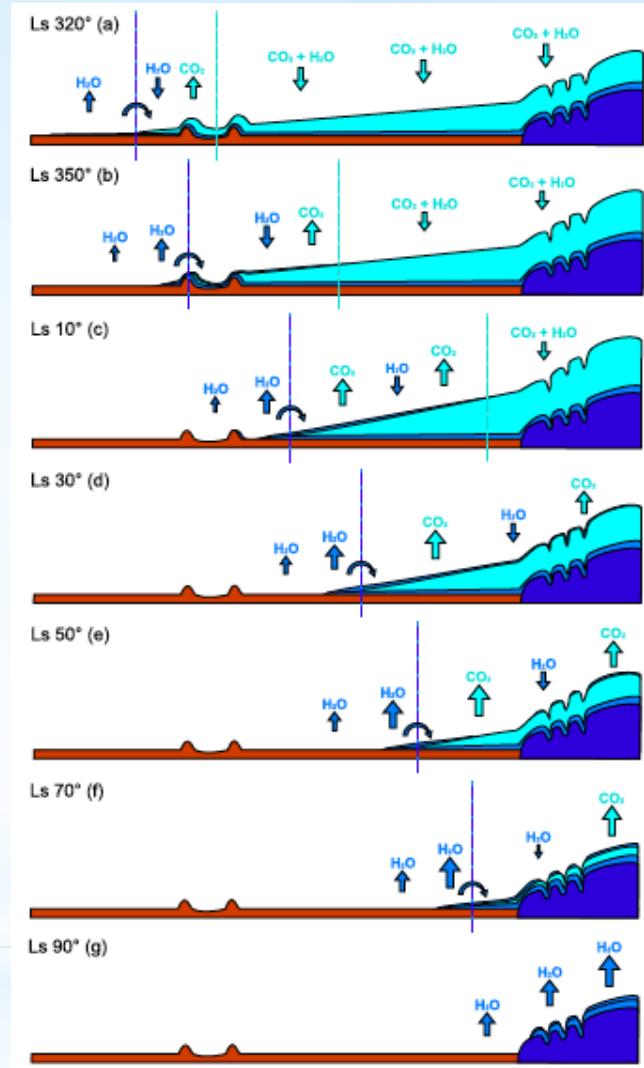


# Primary Variables Observed

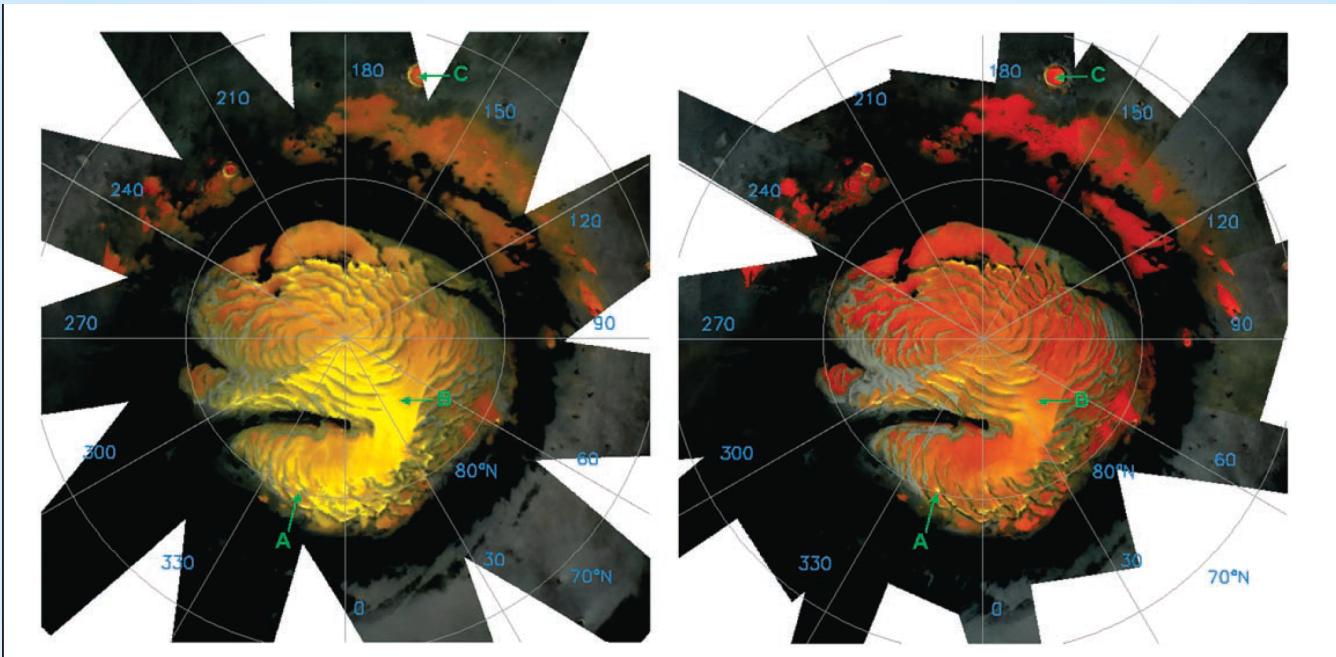
# North Cap Recession



Brown et al. 2012



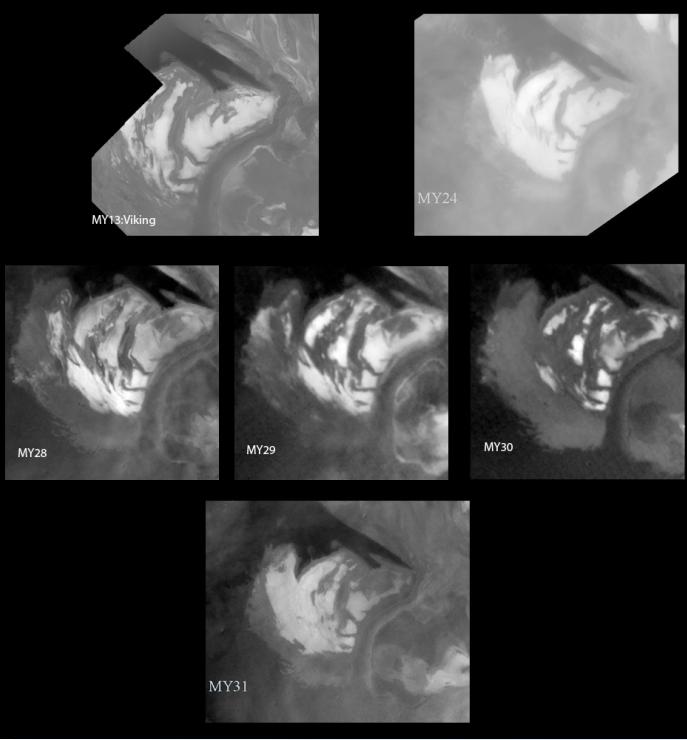
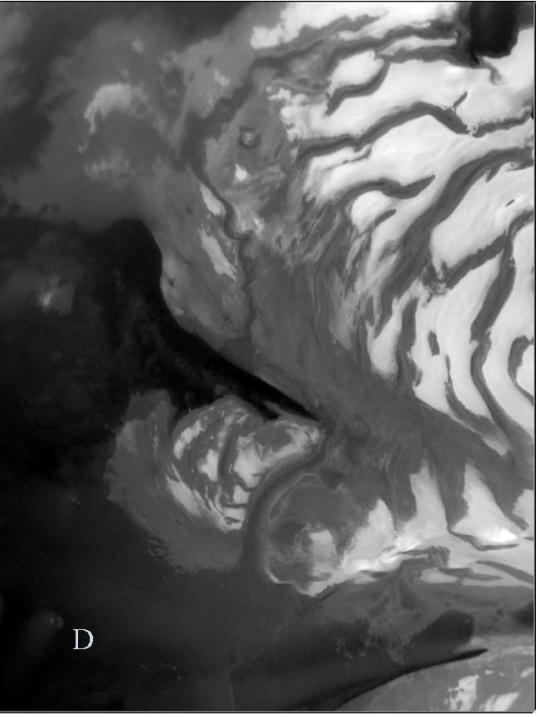
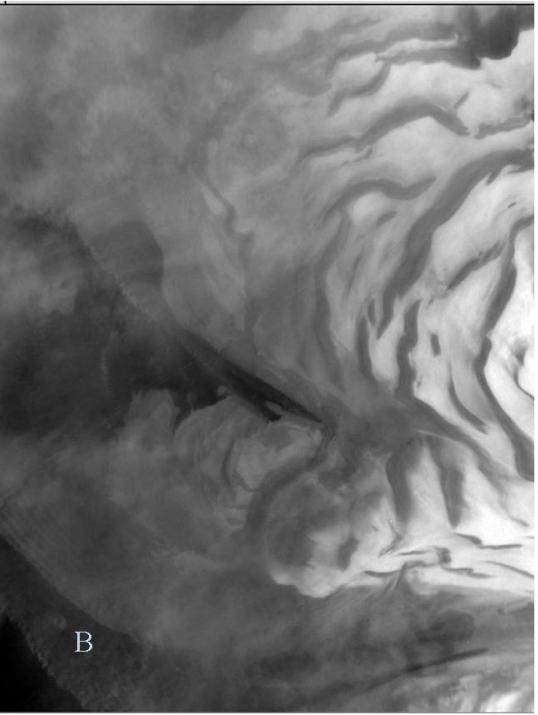
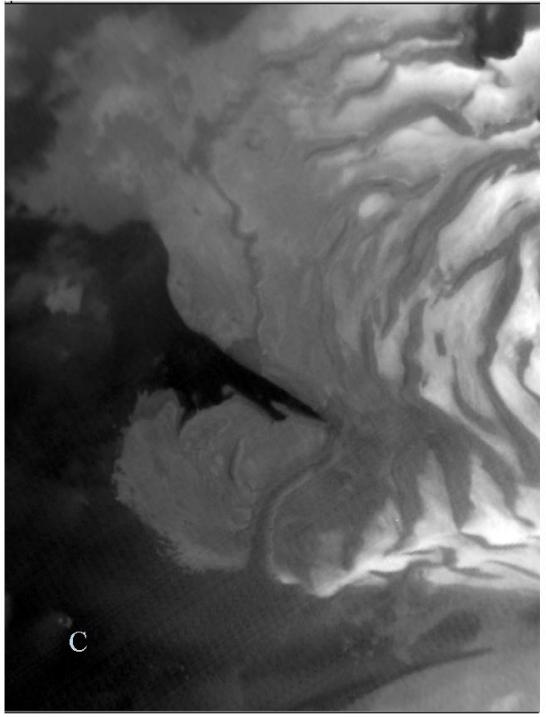
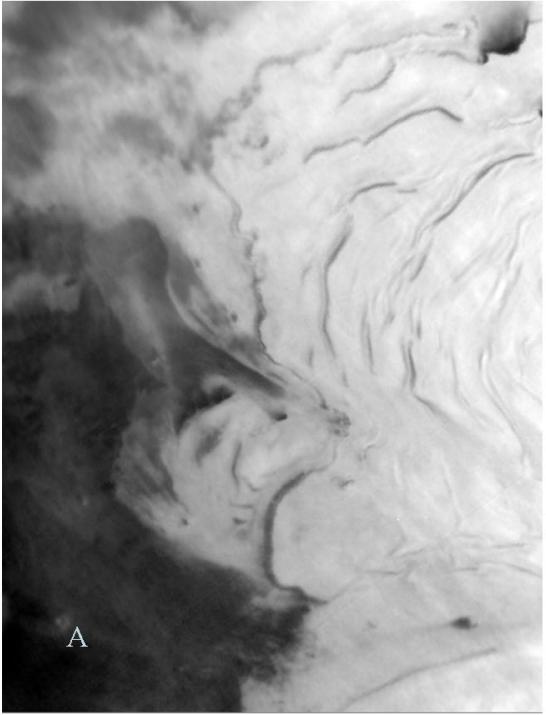
In contrast to retreating south cap, water ice signature appears throughout seasonal cap retreat in the north.



## ◆ Seasonal Change in Grain Size

Composite is 3 wavelengths in the water bands 1, 1.2 1.5  $\mu\text{m}$   
Shift from yellow to red  $\sim 100 \mu\text{m}$  to 1 mm

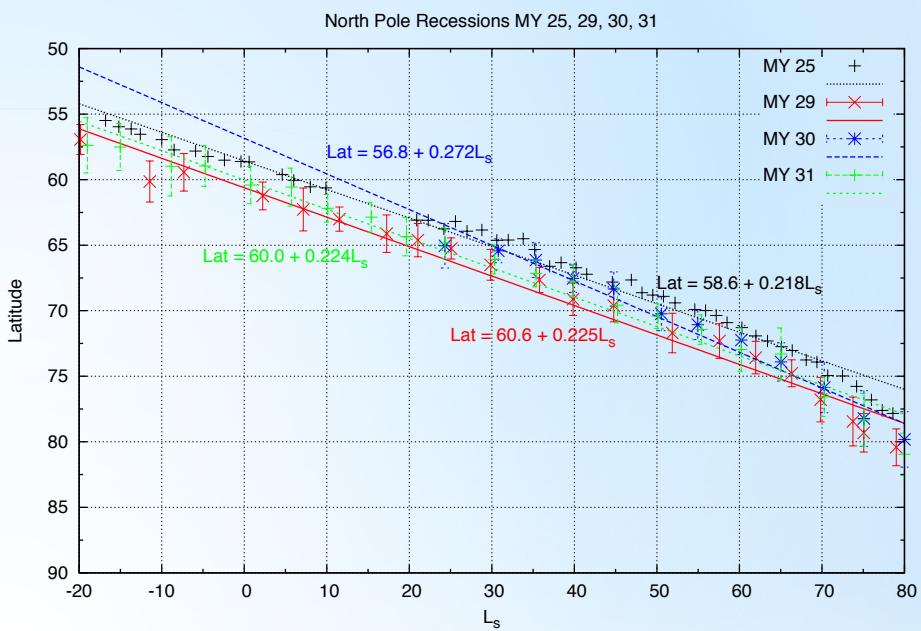
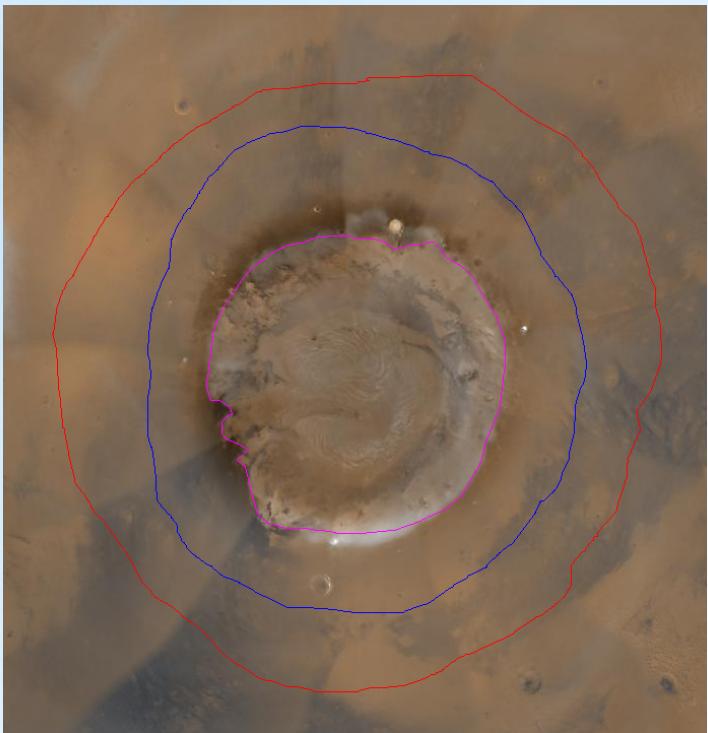
Langevin et al. 2005, Ls 96 vs 108



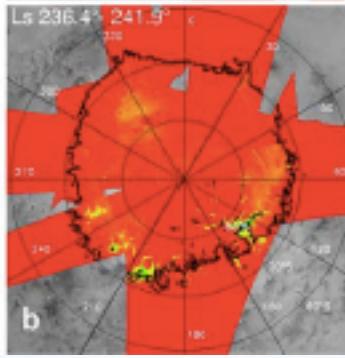
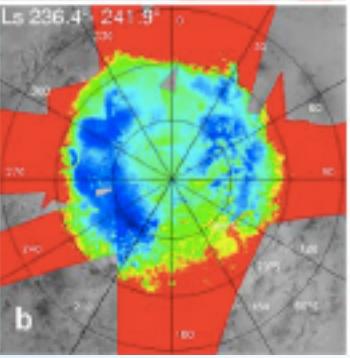
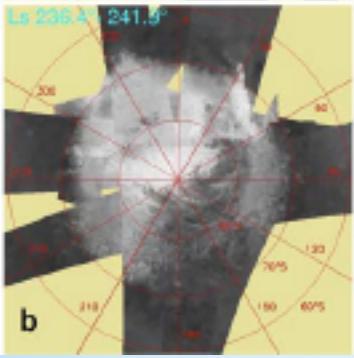
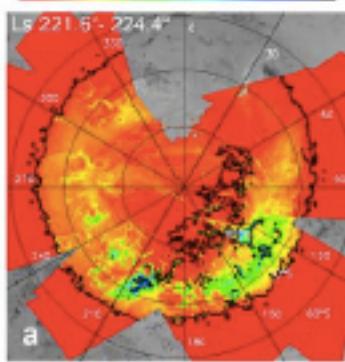
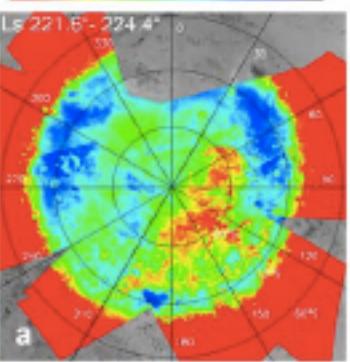
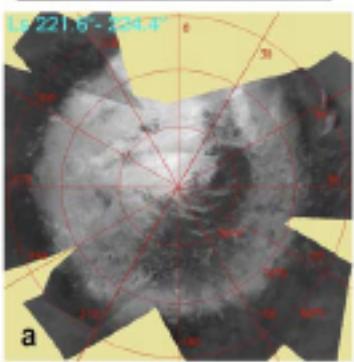
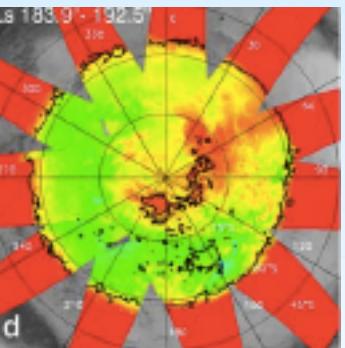
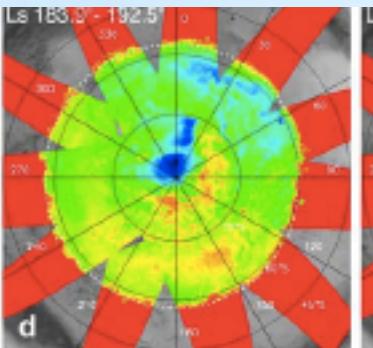
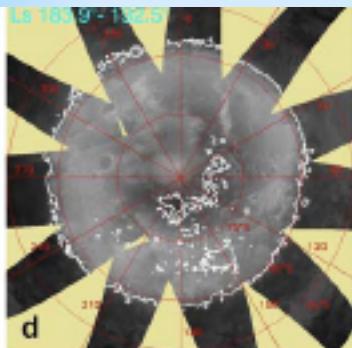
North retreat: Last of the CO<sub>2</sub> (Ls 73), last of the H<sub>2</sub>O (87), minimum of high albedo (95), return of high albedo (118).

Abalos Mensae over many MY at Ls 137.

Dust and frost mobility.  
Ice annealing?



# ◆ North Recession Curves



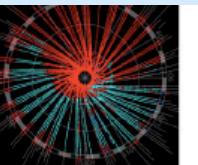
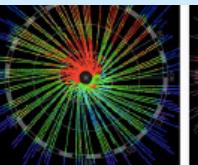
1400 nm absorption depth



GRISM Ls  
191.5-199.2

DOY059-072  
 $n=463$

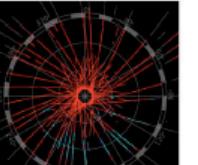
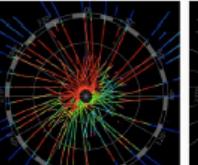
MARCI Ls  
195.6



GRISM Ls  
204.6-207.6

DOY081-086  
 $n=201$

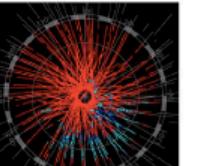
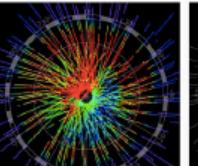
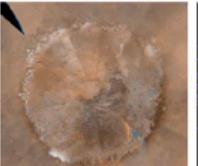
MARCI Ls  
205.2



GRISM Ls  
209.4-216.1

DOY089-100  
 $n=447$

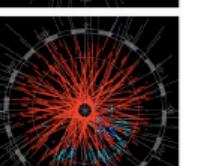
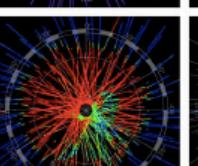
MARCI Ls  
211.8



GRISM Ls  
216.7-224.8

DOY101-114  
 $n=434$

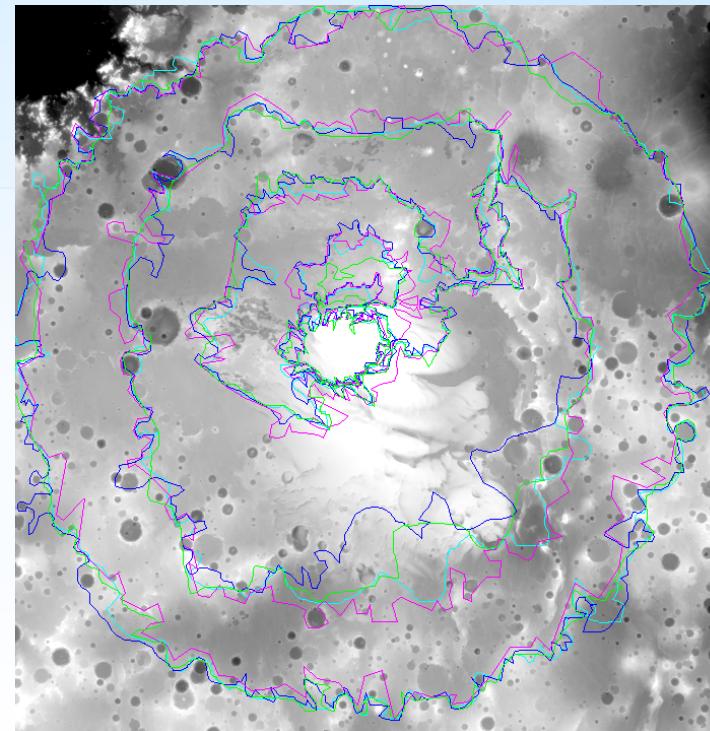
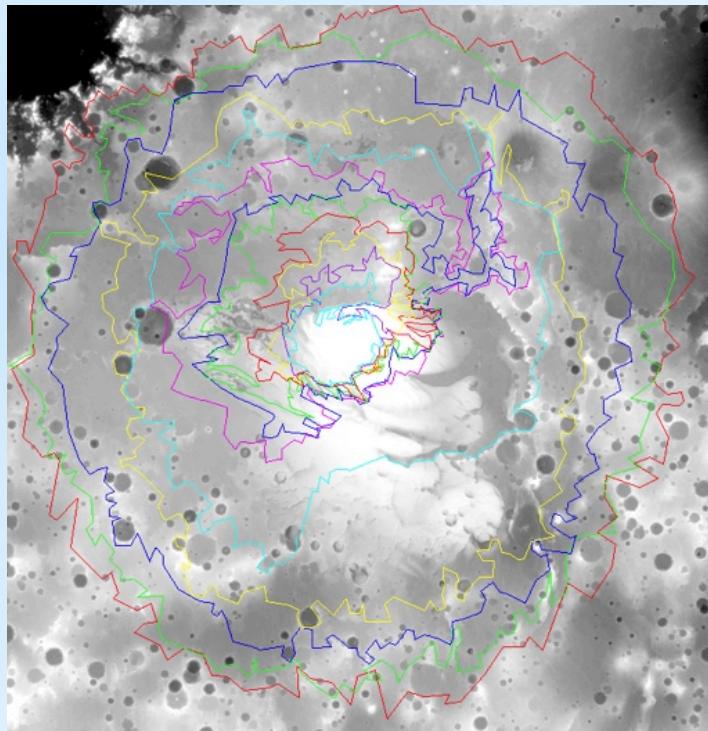
MARCI Ls  
221.6



Brown et al. 2010

Langevin et al. 2007

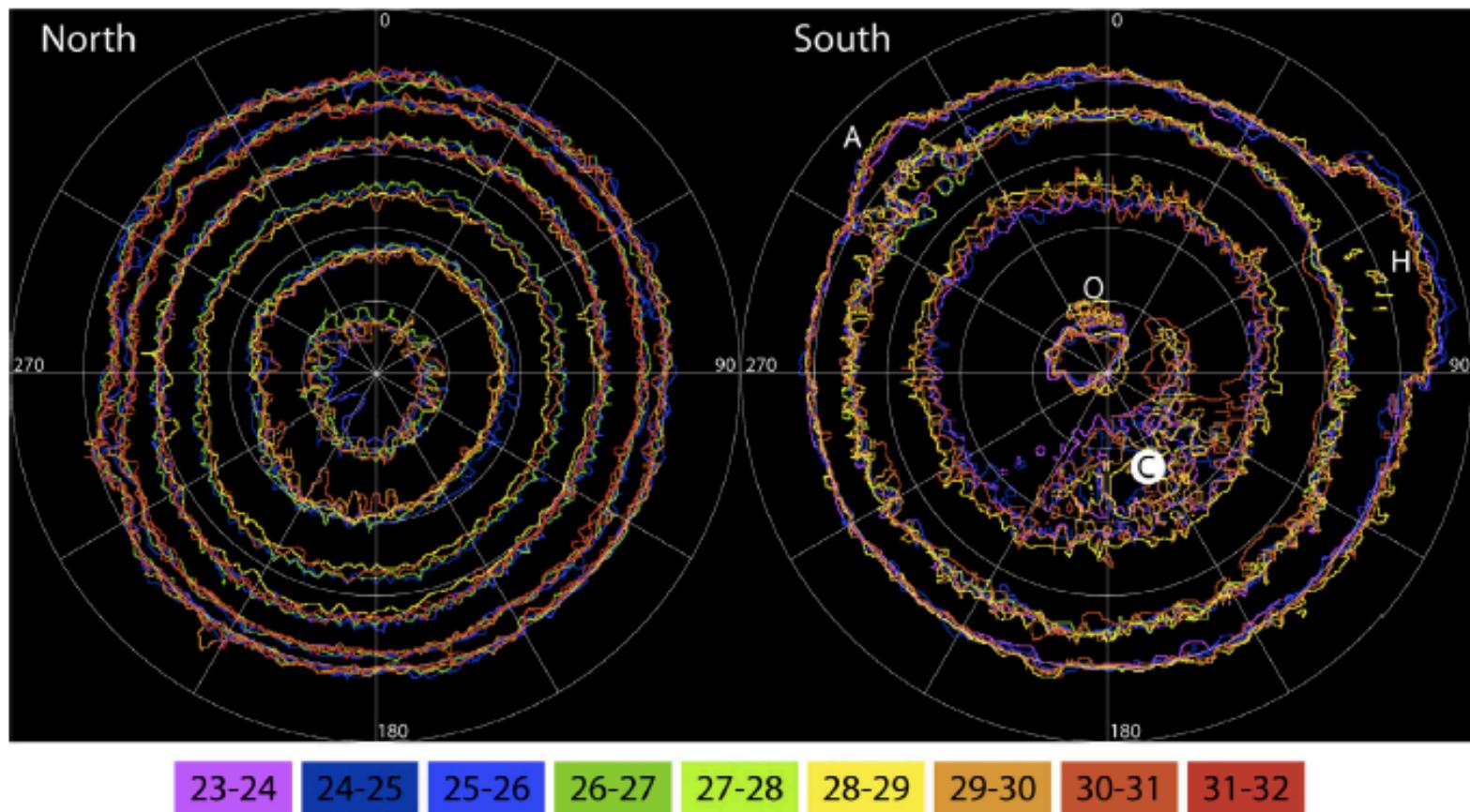
South Cap Retreat



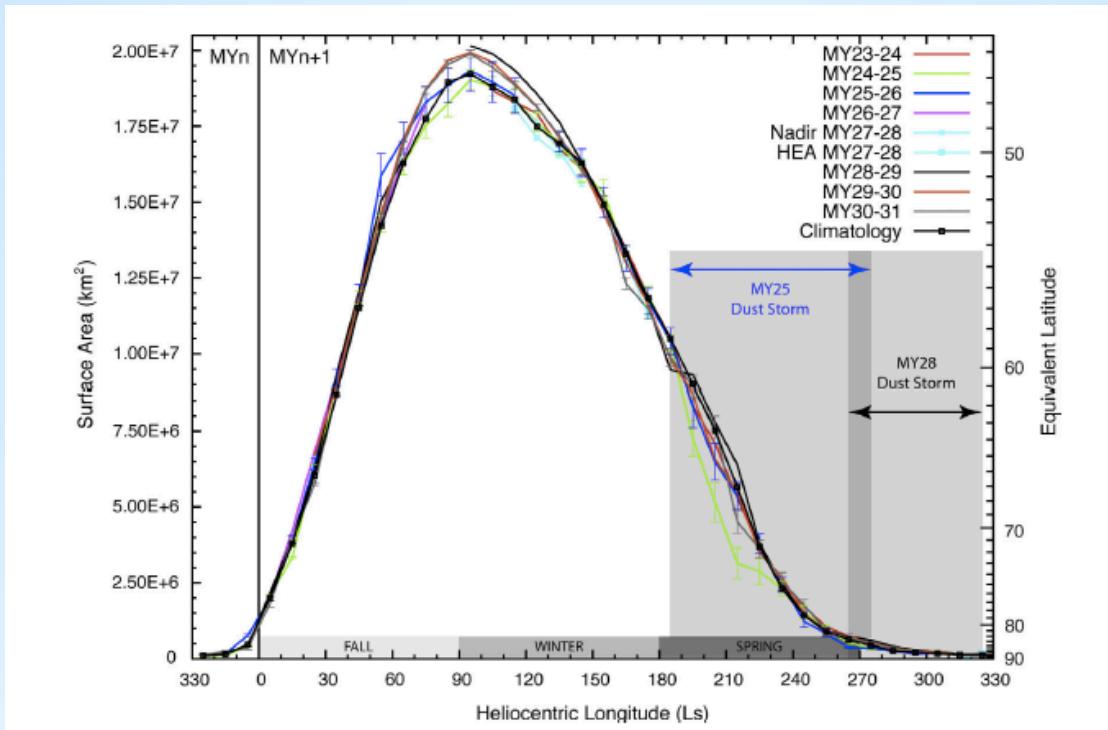
# MARCI Retreat Contours

MY 29  
Ls 205 to 315, every 10

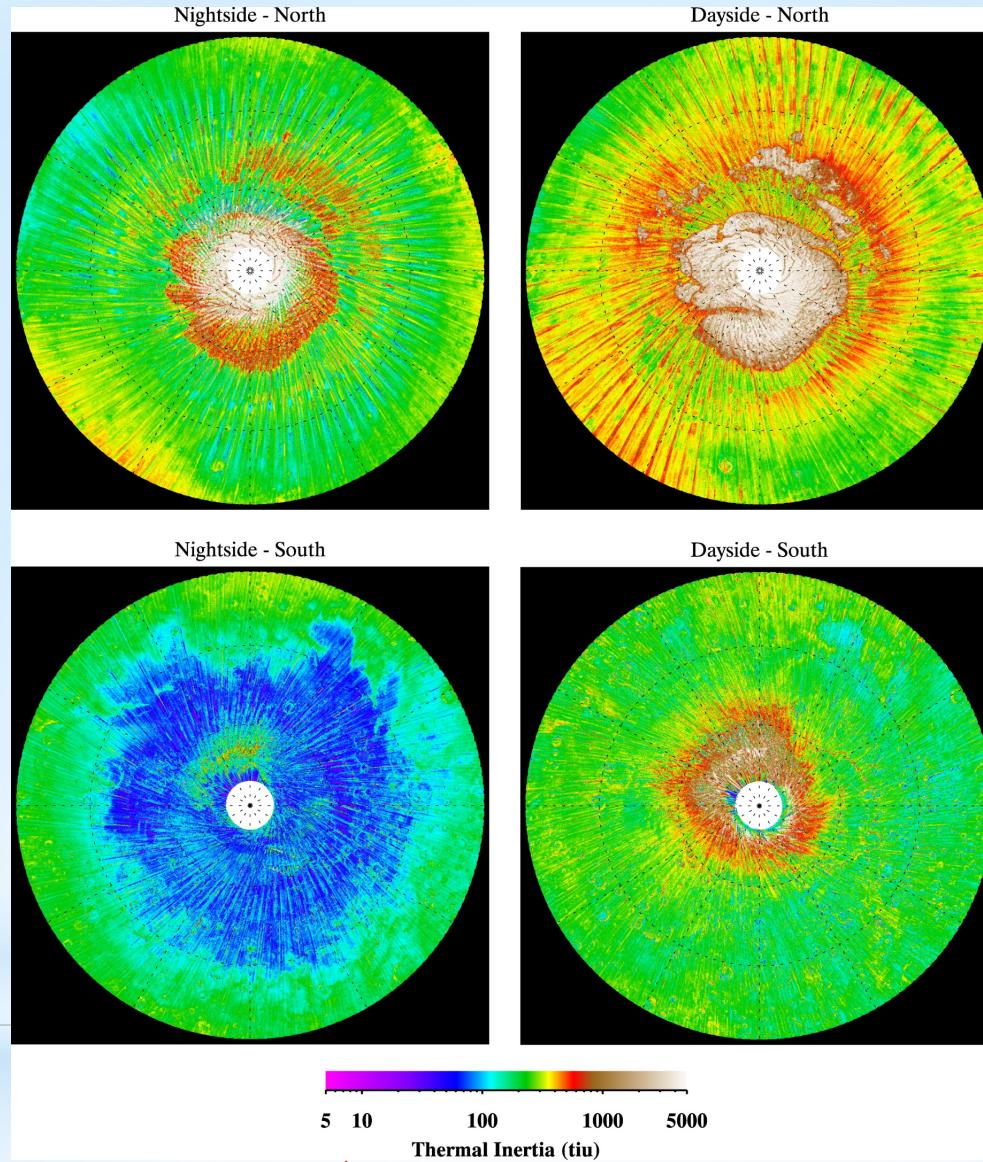
MY 28,29, 30, 31,  
Ls 205, 235, 265, 295, and 315



Piqueux et al 2015. Synthesized 8 years of thermal data from TES & MCS.



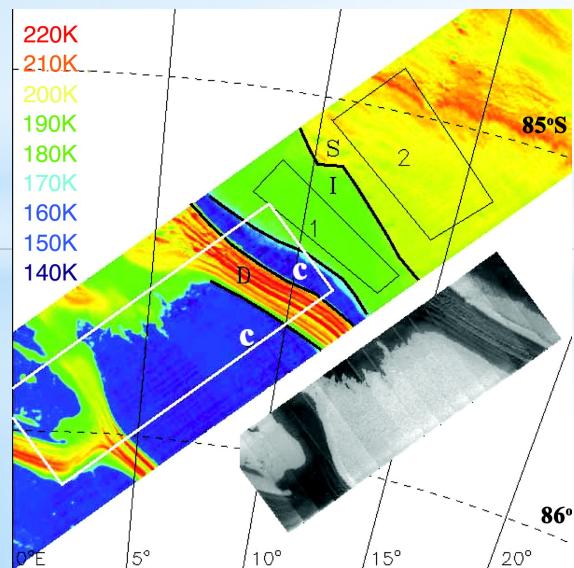
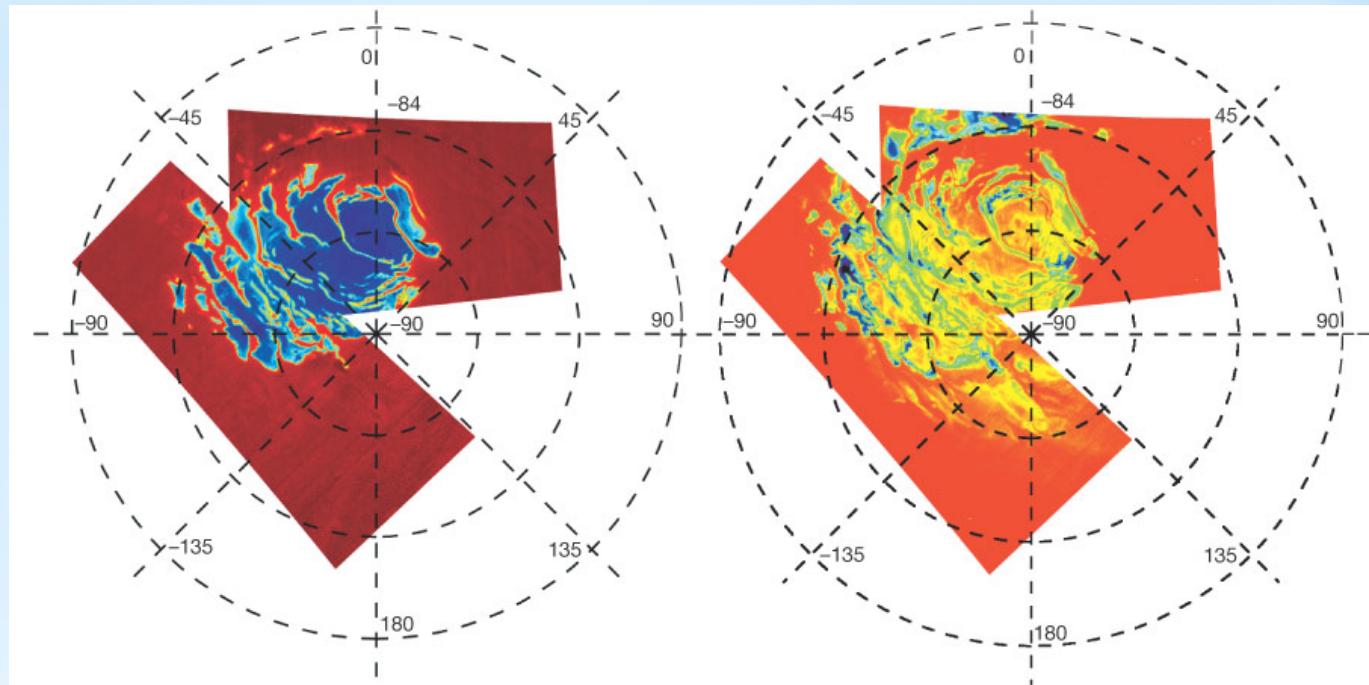
# ◆ Multiple years – Climatological Model



Putzig and Mellon 2007.



Thermal Inertia

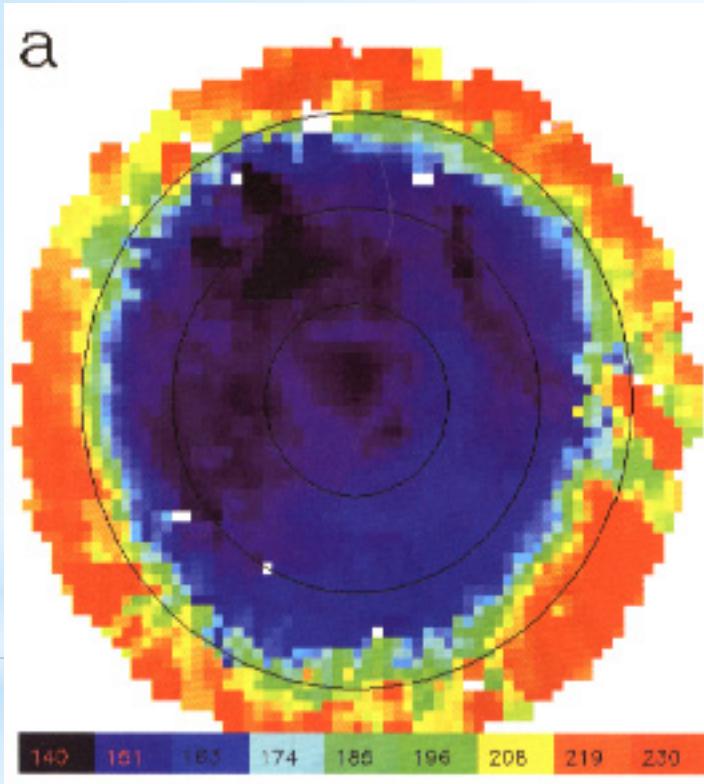


# Water Ice Outlier

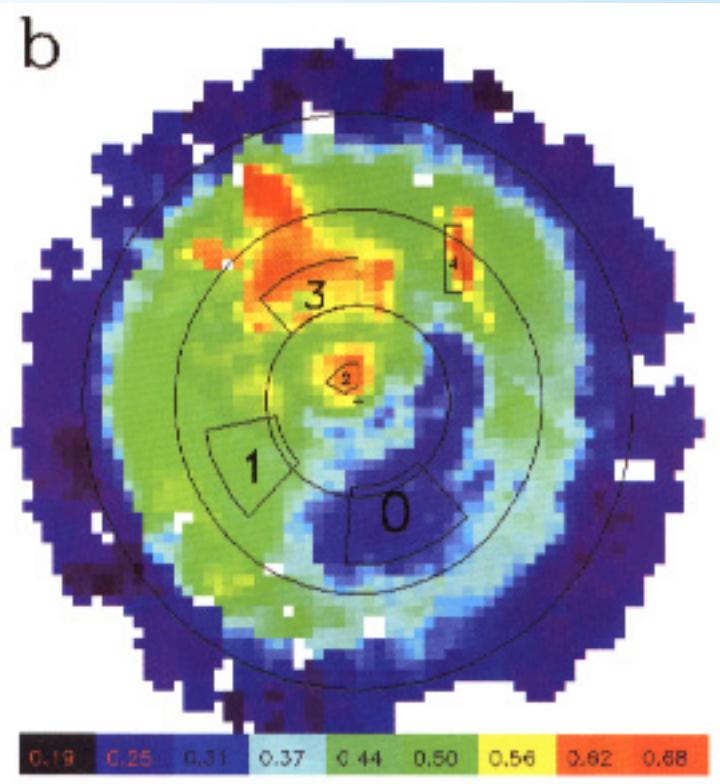
**Titus et al. Science 2003  
Bibring et al. Nature 2004**

# Cryptic Terrain

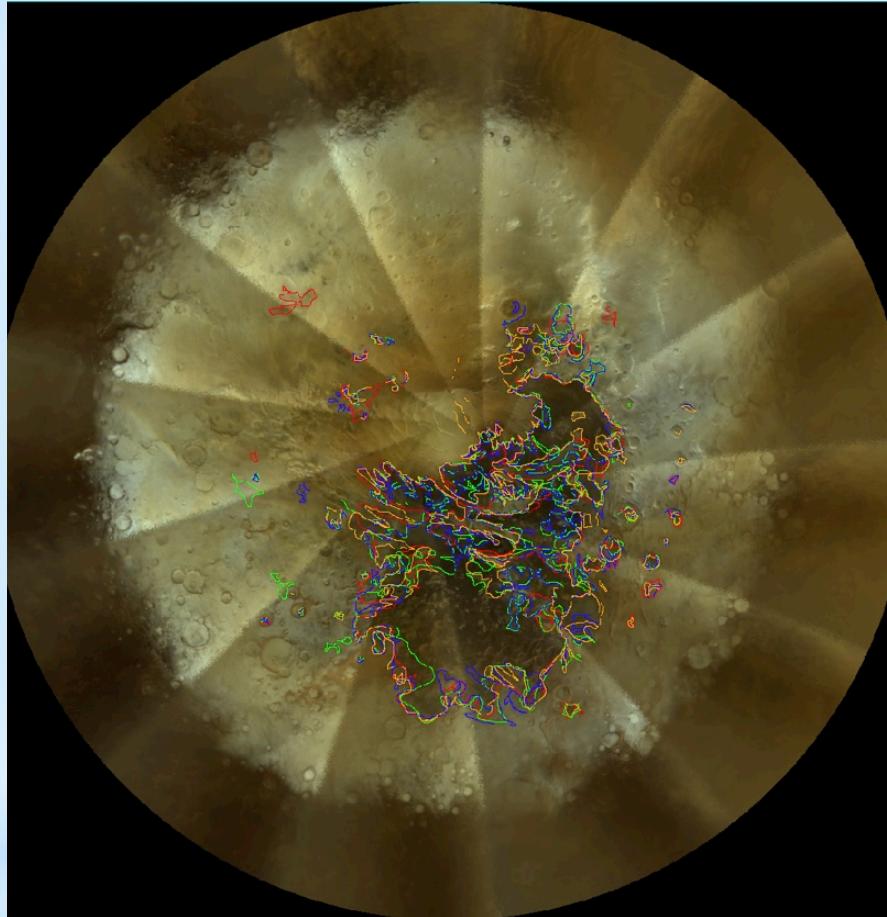
Temperature



Albedo



Kieffer et al. 2000. Cryptic region transparent slab ice.



- Cryptic terrain occurs in same general location but is not linked to elevation.
  - Anti-Arkyre hemisphere
  - Lack of ground ice or other substrate properties of the PLD, including small scale local topography additional influences.
  - Once initiated, surface dust, strong albedo-temperature influence.
  - MARCI finds several smaller patches of “cryptic” terrain outside the classic boundaries.
  - Cryptic terrain develops and leads to isolation of the Mtns of Mitchel.

## ◆ MARCI Observations

Calvin et al. 2017

- Examples from the NEX-SAG Report
  - HiRISE follow on
  - MARCI follow on with SWIR for ices
  - SWIR mapper (6m/pixel)
  - TIR imaging (30m/pixel)
  - Lidar w/ Compositional ability



Potential Future Orbital  
Instruments

- Examples from 2020 rovers
  - Mast mounted imaging + spectroscopy (Supercam, with IR for distinguishing ices)
  - Drill imager (MA\_MISS) or Spectrometer ( $\mu$ -Omega, Lucina, Raman, Tunable Laser Spectroscopy)
- Of interest- CO<sub>2</sub> in clathrates or complexes with water, trace other ices, organics

Spacecraft	Instrument	Best Spatial Footprint	Spectral Resolution	Wavelength range in $\mu\text{m}$
Orbit	Mariner 6, 7	200 km	1-2%	1.9 to 14.4
	Mariner 9	126 km	2.4 cm <sup>-1</sup>	5 to 50
	Phobos-2	22 km	40 nm	0.76 - 3.1
	MGS	3 km	10 cm <sup>-1</sup>	6 to 50
	Mars Express	300 m	14 nm	0.35 to 5.1
	PFS (atmosphere)	12km	1.3 cm <sup>-1</sup>	1.2 to 45
In Situ	MRO	18 m	6.55 nm	0.36 - 3.9
	MER	~15 cm	10 cm <sup>-1</sup>	5 to 29
	MSL	0.3 mm	< 0.3nm	0.24 - 0.8
	Exo Mars (planned)	5mm @ 20 $\mu\text{m}$	20 cm <sup>-1</sup>	0.9 - 3.1
	Lucina	50 $\mu\text{m}$	7 cm <sup>-1</sup>	2.6 to 67
Mars 2020	LIBS (elemental chemistry)	1cm @100 $\mu\text{m}$	4 cm <sup>-1</sup>	1.8 to 12

 Potential Landed /  
In-situ Instruments